Regular Article

Tales from the prehistory of Quantum Gravity

Léon Rosenfeld's earliest contributions

2 3

Giulio Peruzzi^a and Alessio Rocci

Department of Physics and Astronomy "G. Galilei", via Marzolo 8, 35131 Padova, 4 5

Italy

- Received 24 March 2017 / Received in final form 24 September 2017 6
- Published online (Inserted Later) 7
- © EDP Sciences, Springer-Verlag 2018 8

Abstract. The main purpose of this paper is to analyse the earliest 9 work of Léon Rosenfeld, one of the pioneers in the search of Quan-10 tum Gravity, the supposed theory unifying quantum theory and general 11 relativity. We describe how and why Rosenfeld tried to face this prob-12 lem in 1927, analysing the role of his mentors: Oskar Klein, Louis 13 de Broglie and Théophile De Donder. Rosenfeld asked himself how 14 quantum mechanics should *concretely* modify general relativity. In the 15 context of a five-dimensional theory, Rosenfeld tried to construct a 16 unifying framework for the gravitational and electromagnetic interac-17 tion and wave mechanics. Using a sort of "general relativistic quantum 18 mechanics" Rosenfeld introduced a wave equation on a curved back-19 ground. He investigated the metric created by what he called 'quantum 20 phenomena', represented by wave functions. Rosenfeld integrated Ein-21 stein equations in the weak field limit, with wave functions as source 22 of the gravitational field. The author performed a sort of semi-classical 23 approximation obtaining at the first order the Reissner-Nordström met-24 ric. We analyse how Rosenfeld's work is part of the history of Quantum 25 Mechanics, because in his investigation Rosenfeld was guided by Bohr's 26 correspondence principle. Finally we briefly discuss how his contribu-27 tion is connected with the task of finding out which metric can be 28 generated by a quantum field, a problem that quantum field theory on 29 curved backgrounds will start to address 35 years later. 30

'A study of history of science [...] shows that the natural attitude of a scientist is to be inspired by their predecessors, but always taking the liberty of doubting when there are reasons for doubt.'

Oskar Klein

1 Introduction 35

31

32

33

34

In the physics community, the word Quantum Gravity (QG) is today associated with 36 the task of quantizing gravity, directly or indirectly, in order to unravel a quantum 37

e-mail: giulio.peruzzi@unipd.it

structure of space and time. Despite many approaches, e.g. String Theory, Super-38 gravity (N = 8), Loop Quantum Gravity, non-commutative geometry and so on, a 39 consistent theory is still lacking. From the point of view of History and Philosophy 40 of Science: 'QG, broadly construed, is the physical theory (still "under construction") 41 incorporating both the principles of general relativity (GR) and quantum theory' 42 [emphasis added] [Rickles & Weinstein, 2016]. "Broadly construed" means that all 43 the attempts in this direction have contributed to our modern understanding of the 44 difficulties in constructing a consistent theory of QG, even those approaches that 45 did not quantize the gravitational interaction. To name one, quantum field theory 46 (QFT) on curved backgrounds increased our knowledge on the physics of Black Holes 47 [Hawking, 1975]. Furthermore, from a point of view of the integrated History and 48 Philosophy of Science (&HPS), the fact that the theory is still under construction 49 represents a unique opportunity for studying the process of a theory's formation from 50 the inside (in Kuhnian words "a revolution in progress"). 51

Usually the history of QG starts in 1930 with the first attempts to reconcile 52 the budding quantum field theory with gravity made by Léon Rosenfeld [1930a,b] 53 (cf. English translation [Léon Rosenfeld, 2017] and the accompanying commentary 54 [Salisbury & Sundermeyer, 2017]). In the first paper the author tried to find out what 55 would be the gravitational field produced by light in a weak-field approximation. This 56 paper marked the beginning of what is today called the *covariant approach*. In this 57 work the quantization procedure was applied to the electromagnetic field only, the 58 metric field being an operator because it is a function of the Maxwell field. In the 59 second paper, conversely, he tried to apply the quantization procedure directly to 60 the gravitational interaction, employing a tetrad gravitational field rather than the 61 conventional metric. This paper marked the beginning of the today called *canonical* 62 approach. Before Rosenfeld's attempts, soon after the birth of GR in 1915, researchers 63 tried to apply the theory of gravity to the microscopic world. The best known example 64 is Einstein's claim of 1916. When he discovered that a mass should emit gravitational 65 waves, Einstein pointed out the need to modify GR [Einstein, 1916]. Of course what 66 he had in mind was Bohr's old move that classical electrodynamics was not applicable 67 in his model of orbiting electrons. In a similar way GR had to be modified with respect 68 to its application to the microscopic world. Einstein's suggestion was not an isolated 69 episode. Recent developments in the history of QG show that in the fifteen years 70 before Rosenfeld's attempts many authors tried to reconcile the old quantum theory 71 or quantum mechanics (QM) with gravity [Stachel, 1999; Rickles, 2005, 2013; Hagar, 72 2014; Rocci, 2015a,b]. For this reason the period between 1915 and 1930 could be 73 called a prehistory era. 74

Exploring this time frame, the term "Quantum Gravity" must be necessarily 75 interpreted in a broad sense, because in the period between 1916 and 1930 the quan-76 tization procedure was a concept under construction. As far as we know, before 1930, 77 there were no attempts that tried to quantize the gravitational field directly. Before 78 going on, we therefore briefly summarize the evolution of the quantization procedure 79 during this period [Mehra & Rechenberg, 2001]. Between 1916 and 1924, the con-80 struction of atomic models was one of the main tasks of the old quantum theory. The 81 quantization procedure of the atomic model was performed by applying the Epstein-82 Sommerfeld-Wilson rules. After 1925, with the birth of QM, the investigation of the 83 atomic phenomena was pursued by wave mechanics (WM) and matrix mechanics 84 (MM). In the first formulation of QM, electrons are represented by normalized wave 85 functions. WM was born by using Hamilton Jacobi (HJ) analogy between particle and 86 waves [Schrödinger, 1926]. The quantization procedure consisted in writing a wave 87 equation and in imposing the boundary condition on wave functions. The second 88 formulation of QM focused on observable quantities. MM was born by attempting to 89 formulate a new theoretical technique for the determination of the intensities of quan-90 tum transitions, using the anharmonic oscillator as a toy model [Blum et al., 2017]. 91

The classical position and its conjugated momentum in the Hamiltonian formulation 92 were treated as "q-numbers", that today are known as operators. The name "q-93 numbers" stands for quantum numbers, in contrast with "c-numbers", i.e. the usual 94 classical variables, like e.g. classical position and momentum of a particle [Darrigol, 95 1992. The quantization procedure consisted in imposing the commutation relations 96 between these q-numbers. In 1926 Schrödinger pointed out the equivalence between 97 the two formulations, but WM remained the preferred point of view in attempting 98 to generalize Schrödinger approach in the context of both Special and General Rel-99 ativity [Rocci, 2015b]. In 1927 many new concepts were introduced: the description 100 of spin with two components wave functions, its statistical interpretation, the uncer-101 tainty relations. At the end of 1927 Oskar Klein and Pascual Jordan introduced for 102 the first time the quantum commutation relations for the scalar field operators, but 103 the general approach was developed by Heisenberg and Pauli at the end of 1929. 104

Rosenfeld was a protagonist of this early period as well. As stated in the intro-105 duction of a recent biography of Rosenfeld [Jacobsen, 2012], the Belgian physicist 106 is a blank sheet in the history of science literature, 'but he was at the centre of 107 modern physics as one of the pioneers of quantum field theory and quantum elec-108 trodynamics in the late 1920s and the 1930s' (Jacobsen, 2012; p. 1). In spite of 109 the fact that he initiated two of the major research areas in the history of QG, the 110 covariant and the canonical approaches, Rosenfeld never considered his early work 111 as an important contribution [Kuhn & Heilbron, 1963]. The aim of this paper is to 112 offer a historical analysis "in context" of the papers published by Rosenfeld at the 113 beginning of his career: [Léon Rosenfeld, 1927a,b,c,d,e]. In particular we will focus 114 on the aspects concerning the conciliation between GR and the WM, that produced 115 a first attempt to find the metric generated by "charged quantum matter", using a 116 wave-mechanical approach. Rosenfeld was persuaded, at that time, that he had found 117 a quantum modification of the flat metric, using the correspondence principle. He per-118 formed a semi-classical approximation in order to compare his quantum metric with 119 the external Reissner-Nordström (RN) metric. Aside from the fact that this attempt 120 is important by itself, it contained the seeds for his following work [Léon Rosenfeld, 121 1930a, nevertheless Rosenfeld later become one of the opponents to any quantization 122 of the gravitational field without any experimental evidence for the necessity to do 123 it [Léon Rosenfeld, 1963]. 124

The paper is organized as follows. In Section 2 we briefly introduce Rosenfeld's 125 life and we put it in the context of the prehistory of QG. In Section 3 we review the 126 work of the authors that influenced the professional training of the young Rosenfeld 127 in 1927: Oskar Klein, Louis de Broglie and Théophile De Donder. In particular we 128 will focus on the analogies and on the differences among these authors. In Section 4 129 we present Rosenfeld's attempt to reconcile GR with WM. At the beginning we shall 130 focus on his first paper, discussing how Klein, de Broglie and De Donder influenced 131 Rosenfeld's work. Then we shall review the papers written by Rosenfeld in 1927, 132 where a general relativistic version of Bohr's correspondence principle emerged. We 133 shall also analyse the role played by Klein, and indirectly by Bohr, in suggesting the 134 first use of the correspondence principle in the context of QG. At the beginning of 135 Section 4 we shall focus on what Rosenfeld wanted to achieve. In the last part of 136 the section, i.e. 4.3, we briefly present a modern interpretation of his approach and a 137 perspective on how the analysed papers would influence Rosenfeld's subsequent work 138 on the search of a quantum theory of gravity. In Section 5 we summarize the basic 139 stages of our paper without entering into technical details. 140

In the Appendices, we describe with more details some calculations left out in the main text.

¹⁴³ 2 The prehistory of QG and the young Rosenfeld

The prehistory of QG can be naturally divided into two parts. The first period from 144 1915 to 1924, was dominated by attempts to understand the role of GR in constructing 145 planetary models of atoms [Jaffé, 1922; Jeffery, 1921; Lodge, 1921; Vallarta, 1924]. 146 With the birth of QM in 1925–26 a new era began, because the classical concept of 147 trajectory had become problematic in the atomic realm. In particular, the second 148 period of the prehistory of QG from 1925 to 1930, was dominated by WM and by 149 attempts which tried to generalize Schrödinger's approach in the context of Special 150 Relativity (SR) and GR. In fact, between the two alternative formulations of QM, 151 MM and WM, the second formulation was the preferred one by the authors of the 152 period who tried to find a unique framework describing quantum phenomena and the 153 gravitational interaction [Rocci, 2015b]. In this respect, as we will see, Léon Rosenfeld 154 was not an exception. 155

The career of the young Belgian physicist had started with the accidental reading 156 of Schrödinger's communications [Schrödinger, 1926], as he recollected during an 157 interview with Thomas S. Kuhn and John L. Heilbron in 1963 Kuhn & Heilbron, 158 1963]. After completing his studies, Rosenfeld left the University of Liège and moved 159 to Paris at the end of 1926 to meet Louis de Broglie, where, as he recollected in the 160 interview, he spent most of his time learning what he had missed at Liège Kuhn & 161 Heilbron, 1963]. Rosenfeld himself stressed that he attended a course on relativity in 162 Liège and that the lecturer was an opponent of the new theory. In Paris, he attended 163 many lectures, e.g. Langevin's lectures at the College de France, and he studied a 164 lot of books, including Eddington's book on GR [Eddington, 1923]: 'I was anxious to 165 do some research, and then the only research I did was in just combining my freshly 166 acquired knowledge of relativity with wave mechanics [...]' [Kuhn & Heilbron, 1963]. 167

A key ingredient of this second period in the prehistory of QG is the enlargement 168 of the four-dimensional space-time by the introduction of a fifth space-like dimension 169 in order to look for a unified picture of the gravitational force, the electromagnetic 170 interaction and the quantum behaviour of particles, described by a wave function. The 171 idea was not new. The founding father of this approach is Theodore Kaluza [1921] who 172 had noted that a five-dimensional theory of "pure gravity", i.e. without any matter 173 content but with the electromagnetic potentials represented by specific components 174 of the metric field, seems to offer a unified framework to describe the usual four-175 dimensional gravitational and electromagnetic interactions.¹ In 1927 many authors 176 tried to harmonize Kaluza's picture with WM, and started explicitly from the German 177 physicist's 1921 paper.² The most well-known contribution was Oskar Klein's³ work, 178 who developed his ideas from 1926 to 1927. Less known contributions were the papers 179 written by Louis de Broglie [1927b] and Léon Rosenfeld [1927a,b,c,d,e]. During the 180 year spent in Paris, Rosenfeld started to interact frequently with de Broglie, discussing 181 for example the problem of spin. It was the Belgian physicist who drew de Broglie's 182 attention on the five-dimensional approach. As a consequence the French physicist 183 published a paper, in 1927, on this topic [Louis de Broglie, 1927b; Kuhn & Heilbron, 184 1963]. During Kuhn's interview Rosenfeld also recollected that he was anxious to 185 apply his new acquired knowledge to relativity, and that the first goal he wanted to 186

¹More precisely Gunnar Nordström also tried a similar approach before Kaluza [Nordström, 1914], but the Norwegian mathematician described the gravitational interaction using a scalar field instead of a tensor field.

²Kaluza's approach was completely classical. He was afraid that quantum theory could invalidate his five-dimensional approach, as he explicitly stated at the end of his paper ([Kaluza, 1984]; p. 8).

 $^{^{3}}$ The modern multidimensional approach used by e.g. supergravity and string theory is called Kaluza-Klein approach in honour of these two authors, but the modern approach is different from that of the Fathers. For a review of the modern approach and a comparison with the old one see Duff et al. [1986].

achieve was to develop 'the wave equation in five dimensions' [Kuhn & Heilbron, 1963]. 187 188 On this subject Rosenfeld published two notes during his stay at the Ecole Normal in Paris: Léon Rosenfeld [1927a] and Léon Rosenfeld [1927b]. Why did Rosenfeld 189 decide to embark on a five-dimensional adventure? What attracted him? What was 190 Rosenfeld's point of view at that time? In the case of Klein's work the answer was 191 well known, because the Swedish physicist himself answered the question. As we will 192 see, Klein, de Broglie and Rosenfeld constructed their five-dimensional approaches 193 starting from different perspectives and we will try to make clear what considerations 194 suggested to each of the three authors how to develop a five-dimensional picture. 195

Another important role for the young Rosenfeld was played by Théophile De Don-196 der. Like Rosenfeld, De Donder was a Belgian researcher, older and more experienced. 197 De Donder was an enthusiastic supporter of Einstein's theory. As we will see, soon 198 after the birth of QM he tried to explain the existence of atomic stable orbits with 199 the help of GR, but he always followed a classical approach [De Donder, 1926a,c; 200 De Donder & van den Dungen, 1926b]. As Rosenfeld recollected: 'I published a note 201 which I sent to him to be presented to the Belgian Academy. De Donder was the least 202 critical person you can imagine, he was enthusiastic about it. So he asked me then 203 to come to Brussels, he wanted to have me in Brussels; I wanted to go abroad a bit 204 more, but I worked for a month with him in Brussels.' [Kuhn & Heilbron, 1963]. As we 205 shall see, one of the main consequences of the Rosenfeld-De Donder collaboration in 206 1927 was the physical interpretation of the assumptions made by Rosenfeld in his first 207 paper, with the introduction of Bohr's correspondence principle in the context of QG, 208 contained in Léon Rosenfeld [1927c,e] and De Donder [1927b]. In October 1927 the 209 fifth Solvay conference took place in Brussels and on that occasion De Donder tried 210 to attract attention to Rosenfeld's work. This Solvay conference is well known to his-211 torians of Physics, because it indicates the start of the famous Einstein-Bohr debate. 212 The young Belgian physicist was not officially admitted to attend the conference, but 213 de Donder invited Rosenfeld to follow him. At the conference Rosenfeld met Max 214 Born for the first time and asked him about the possibility of a stay in Göttingen. 215 Born's positive answer permitted Rosenfeld to attend Hilbert's, Born's and Pascual 216 Jordan's lectures (Jacobsen, 2012), p. 18), and it would open the doors to his future 217 collaborations with Pauli, Jordan and many others. All these facts showed the crucial 218 role played by De Donder in Rosenfeld's life. 219

In the next section we will start with a brief summary of the history of Klein's work and its intersection with de Broglie's contribution to the construction of a five-dimensional Universe. Section 3 will end with an introduction of De Donder four-dimensional approach, based on the lectures he gave at MIT in 1925, in order to understand, in Section 4, how De Donder also influenced Rosenfeld's early work.

²²⁵ 3 Oskar Klein's, Louis de Broglie's and Theophile De Donder's ²²⁶ role

227 3.1 The five-dimensional universe: Klein's approach

Klein's investigation of the five-dimensional Universe started in 1926 with the purpose of unifying gravity, electromagnetism and WM [Pais, 2000]. As Klein himself recollected in Klein [1991], he was attracted by two facts. First, he knew that the Hamilton-Jacobi (HJ) equation offers a link between particle dynamics and the propagation of a wave front, in the limit of geometrical optics, suggesting a concrete realization for the wave-particle duality. Secondly, by writing the relativistic HJ equation for a particle moving in a combined gravitational and electromagnetic field, he

noticed that the electric charge would play the role of an extra momentum compo-235 nent: '[...] I gave a lecture course on electromagnetism, towards the end of which 236 I derived the general relativistic Hamilton-Jacobi equation for an electric particle 237 moving in a combined gravitational and electromagnetic field. Thereby, the simi-238 larity struck me between the ways the electromagnetic potentials and the Einstein 239 gravitational potentials enter into this equation, the electric charge in appropriate 240 units – appearing as the analogue to a fourth momentum component, the whole 241 looking like a wave front equation in a space of four dimensions. [emphasis added]^{'4} 242 ([Klein, 1991]; p. 108).⁵ In the summer of 1925 he became 'immediately very eager 243 to see how far the mentioned analogy reached' ([Klein, 1991]; p. 109) and he started 244 to investigate the five-dimensional Riemann geometry to describe the gravitational 245 and electromagnetic interactions in a unified framework, trying also to write a five-246 dimensional wave equation. In the long wavelength limit, the wave equation resembles 247 the eikonal equation for the paths of light rays in geometric optics. These paths 248 follow geodesic lines through a Riemannian space: Klein identified them with five-249 dimensional null-geodesics which reduce, on his assumptions, to four-dimensional 250 trajectories for charged massive particles moving in a combined electromagnetic and 251 gravitational field. Klein's original idea was to follow an analogy with light in five 252 dimensions, even if he wanted to relate five-dimensional geometry with the stationary 253 states of massive particles. Carrying on this work, the Swedish Physicist convinced 254 himself that his approach was only a first step towards the formulation of a theory 255 able to reconcile GR with WM. But this conclusion was contained only in his last 256 paper of the period [Klein, 1927b], a work that Rosenfeld would never cite. 257

Now we briefly retrace the steps followed by Klein in his first paper [Klein, 1926a, 1984]. Klein introduced the following five-dimensional line element⁶:

$$d\sigma^2 = \gamma_{\bar{\mu}\bar{\nu}} dx^{\bar{\mu}} dx^{\bar{\nu}} \,, \tag{1}$$

assuming that the metric tensor did not depend on the new fifth space-like component⁷ x^5 . Then it follows that the allowed coordinate transformations were restricted to the following set:

$$\int x^{\mu} = f^{\mu} \left(x^{0'}, \, x^{1'}, \, x^{2'}, \, x^{3'} \right) \tag{2a}$$

$$\left(x^{5} = x^{5'} + f_{5}\left(x^{0'}, x^{1'}, x^{2'}, x^{3'}\right) \right).$$
(2b)

([Klein, 1984]; p. 11). After noting the invariance of γ_{55} under the coordinate transformations (2a) and (2b), Klein decided to set $\gamma_{55} = \alpha$, where α is a constant. In modern Kaluza-Klein theories γ_{55} is not a constant, it is a real scalar field depending on the

⁷Kaluza called this hypothesis the *cylinder condition*. Using modern language, this means that translations in the fifth direction are isometries and hence that the five-dimensional space-time admits a space-like Killing vector field, namely $\frac{\partial}{\partial x^5}$. Neither Klein nor de Broglie or Rosenfeld mentioned this fact explicitly in their papers.

⁴It is worth noting that in the original paper Klein did not emphasize the role of the electric charge explicitly. Rosenfeld followed a similar reasoning in constructing his wave equation, but stated it explicitly: see the remark after equation (55).

⁵The original reasoning runs backward with respect to the path followed by Klein in the paper, where the author presented his model in an axiomatic way.

⁶In our paper we consider many authors who introduced different notations. We decided to adopt the following conventions. Barred indices refer to the five-dimensional World, $\bar{\mu} = 0, 1, 2, 3, 5$, where the zero component corresponds to a time-like dimension. We use the mostly-plus signature, i.e. $\eta_{\bar{\mu}\bar{\nu}} = diag(-1, +1, +1, +1)$. The unbarred Greek indices correspond to the usual four-dimensional space-time, $\mu = 0, 1, 2, 3$, and Latin indices refer to the three-dimensional spatial coordinates, i = 1, 2, 3. We use International System of Units.

transverse dimensions, called a dilaton field. As O'Raifeartaigh & Straumann [2000] and other authors [Overduin & Wesson, 1997] pointed out, Klein's choice is inconsistent, as we shall explain below after equation (8). Klein rewrote the line element (1) in the following form:

$$d\sigma^2 = \alpha d\theta^2 + ds^2 \,, \tag{3}$$

where

$$d\theta = dx^{5} + \frac{\gamma_{5\mu}}{\alpha} dx^{\mu} \quad ; \quad g_{\mu\nu} = \gamma_{\mu\nu} - \frac{\gamma_{5\mu}\gamma_{5\nu}}{\alpha} \quad ; \quad ds^{2} = g_{\mu\nu} dx^{\mu} dx^{\nu} \quad . \tag{4}$$

Citing Kramers' paper on stationary gravitational fields in four dimensions [Kramers, 268 1922], Klein noted that $d\theta$, equation (4), is invariant under the coordinate transfor-269 mations (2a) and (2b). In fact, following Kramers and remembering that $\alpha = \gamma_{55}$, 270 the invariance of $d\theta$ is transparent if we rewrite it in the following way: $d\theta =$ 271 $dx^5 + \frac{\gamma_{5\mu}}{\gamma_{55}} dx^{\mu} = \frac{1}{\gamma_{55}} \gamma_{5\bar{\mu}} dx^{\bar{\mu}}$. As a consequence, Klein noted that the four components 272 $\gamma_{5\mu}$ transform as a four-vector of the four-dimensional space-time. Following Kaluza, 273 Klein assumed that they would be proportional to the electromagnetic potentials 274 $A^{\nu} = (V; \vec{A})$, introducing another parameter β : 275

$$\frac{\gamma_{5\mu}}{\alpha} = \beta A_{\mu} , \qquad (5)$$

where we defined $A_{\mu} = g_{\mu\nu}A^{\nu}$. We note that $d\theta$ defined in equation (4) is not an exact form and that it can be rewritten as: $d\theta = dx^5 + \beta A_{\mu}dx^{\mu}$. Using $d\theta^2$ invariance and $d\sigma^2$ invariance, it follows that ds^2 is invariant under the coordinate transformations (2a) and (2b). As a consequence $g_{\mu\nu}$ can be interpreted as a four-dimensional metric. After having introduced the five-dimensional curvature scalar \tilde{R} , defined in Appendix B, Klein varied the five-dimensional Einstein-Hilbert action as usual in GR, with respect to the metric $\gamma_{\mu\bar{\nu}}$:

$$\delta_{\gamma} S_5 = \delta_{\gamma} \int_{\Omega} \tilde{R} \sqrt{-\gamma} d^5 x = \int_{\Omega} d^5 x \frac{\delta\left(\tilde{R} \sqrt{-\gamma}\right)}{\delta \gamma_{\bar{\mu}\bar{\nu}}} \delta \gamma_{\bar{\mu}\bar{\nu}} , \qquad (6)$$

where the symbol $\sqrt{-\gamma}$ represents the square root of the negative of the determinant of the metric and the integral is carried out over a closed region Ω , where boundary values of $\gamma_{\bar{\mu}\bar{\nu}}$ are kept fixed. From the principle of stationary action the five-dimensional Einstein equations follow:

$$\delta_{\gamma} \mathcal{S}_5 = 0 \quad \Rightarrow \quad \tilde{R}_{\bar{\mu}\bar{\nu}} - \frac{1}{2} \gamma_{\bar{\mu}\bar{\nu}} \tilde{R} = 0 \quad .$$
 (7)

It is worth noting that neither Klein nor any of the other authors we analysed considered the 55 component of equation (7), because they fixed $\alpha = \text{constant before varying}$ the action. Thanks to all assumptions he made, equation (7) are formally equivalent to the four-dimensional Einstein-like equations coupled to the four-dimensional

267

Maxwell-like equations⁸:

$$\begin{cases} R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = \frac{\alpha\beta^2}{2}T^{em}_{\mu\nu} \\ \partial_{\mu}\left(\sqrt{-g}F^{\mu\nu}\right) = 0 \quad , \end{cases}$$
(8a)
(8b)

where g is the determinant of $g_{\mu\nu}$ defined in equation (4). Choosing to set⁹ 287 $\alpha\beta^2 = \frac{16\pi G}{c^4}$, where G, and c are the Newton constant and the speed of light respectively, Klein justified the identification of $g_{\mu\nu}$ and $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$ with 288 289 our four-dimensional metric and with the electromagnetic tensor respectively. The 290 electromagnetic stress-energy tensor that appears in (8a) is defined by: $T_{\mu\nu}^{em} =$ 291 $F_{\mu}^{\ \alpha}F_{\nu\alpha} - \frac{1}{4}g^{\mu\nu}F_{\alpha\beta}F^{\alpha\beta}$. The condition $\alpha\beta^2 = \frac{16\pi G}{c^4}$ implies $\alpha > 0$. This means that Klein introduced a space-like extra dimension motivated by the need to obtain the 292 293 four-dimensional Einstein equations coupled with Maxwell's equations. Indeed, a 294 space-like coordinate only, i.e. a positive α constant in (8a), produces the correct 295 coupling between electromagnetic and gravitational interactions. In this sense our 296 four-dimensional World is a "projection" of a five-dimensional Universe. 297

As indicated Klein's model is inconsistent, if α is constant. Indeed, if the dilaton is 298 a non trivial scalar function $\alpha(x)$, the 55 component of equation (7) is not trivial and it 299 has the form $\Box \sqrt{\alpha} \sim (\sqrt{\alpha})^3 F_{\alpha\beta} F^{\alpha\beta}$, where the four-dimensional operator \Box , when acting on a scalar function $\alpha(x)$ is defined by $\Box \alpha = g^{\mu\nu} \nabla_{\mu} \partial_{\nu} \alpha$ for a curved four-300 301 dimensional space-time, where ∇_{μ} represents the covariant derivative. This means 302 that a non-zero constant dilaton would imply the too restrictive condition $F_{\alpha\beta}F^{\alpha\beta} =$ 303 0, i.e. that the modulus of the electric field should be proportional to the modulus 304 of the magnetic field. As reported in Overduin & Wesson [1997], this inconsistency 305 was noted by Pascual Jordan [1947] and Yves Thiry [1948] in 1947 and in 1948 306 respectively: all the authors of the period we are considering imposed the constancy 307 of the dilaton, including de Broglie and Rosenfeld, and they were not aware of this 308 inconsistency. 309

In order to reconcile this framework with WM, Klein's idea was to write a five-310 dimensional wave equation in a curved space-time, which was then to be connected 311 with the classical four-dimensional Lorentz equation for a charged particle in the 312 presence of gravitational and electromagnetic fields, in the so called geometrical optics 313 limit. The connection between the two equations, considered by all the authors that 314 we shall analyse, is as follows.¹⁰ In a geometrical optics approximation, the wave 315 equation reduces to the classical HJ equation with a particular Hamiltonian function. 316 After a Legendre transformation, the associated Lagrangian produces five equations 317 of motion. The four equations transverse to the fifth coordinate can be reduced to 318 the Lorentz equation for a charged massive particle. The Lagrangian approach shows 319 that, in five dimensions, charged particles follow a geodesic motion. Klein himself 320 explained this procedure in the introduction of his paper: 'the equations of motion 321 for the charged particles [..] take the form of equations of geodesic lines. If we explain 322 these equations as wave equations because the matter is supposed to be a kind of wave 323 propagation, we are almost naturally led to a partial differential equation of second 324 order, which may be regarded as a generalization of the ordinary wave equation." 325 (Klein, 1984; p. 10). This justifies Klein's idea stated above to connect wave equation 326

 $^{^8 \}mathrm{See}$ Appendix D.3 for a detailed explanation of the formal equivalence in the context of Rosenfeld's work.

 $^{^{9}}$ In his following papers Klein would set $\alpha = 1.$ In de Broglie's and Rosenfeld's paper both constants are present.

 $^{^{10}}$ For a short review with some mathematical details see Appendix A. For a detailed technical explanation of Klein's approach see e.g. [O'Raifeartaigh & Straumann, 2000].

³²⁷ with geodesic lines and it also clarifies why WM had a prominent role in his approach ³²⁸ in unifying GR with QM.

In order to write an equation that generalizes Schrödinger's equation, Klein followed an analogy with light. The equation he found resembles a massless Klein-Gordon (KG) equation,¹¹ what the author called 'our equations for the light wave' ([Klein, 1984]; p. 17). The Swedish physicist was forced to introduce a symmetric tensor $a_{\bar{\mu}\bar{\nu}}$, whose contravariant components are fixed by the request to connect the five-dimensional wave equation with the four-dimensional Lorentz equation for massive charged particles, as we shall see below. Klein's wave equation reads:

$$a^{\bar{\mu}\bar{\nu}} \left(\delta^{\bar{\sigma}}_{\bar{\nu}} \frac{\partial}{\partial x^{\bar{\mu}}} - \Gamma^{\bar{\sigma}}_{\bar{\mu}\bar{\nu}} \right) \partial_{\bar{\sigma}} \Psi = a^{\bar{\mu}\bar{\nu}} \nabla_{\bar{\mu}} \partial_{\bar{\nu}} \Psi = 0 , \qquad (9)$$

where he introduced the covariant derivative $\nabla_{\bar{\mu}}$ using the Christoffel symbols $\Gamma^{\bar{\sigma}}_{\bar{\mu}\bar{\nu}}$, 336 because Klein considered a wave function living on a curved five-dimensional Rieman-337 nian manifold. This means that Klein's wave function is different from Schrödinger's 338 wave function, which lives in configuration space. With this respect, Klein's Ψ resem-339 bles a classical scalar field. From a modern point of view, the introduction of $a^{\bar{\mu}\bar{\nu}}$ 340 sounds strange, because the covariant derivative is usually contracted with the con-341 travariant components of the metric $\gamma^{\bar{\mu}\bar{\nu}}$, which are different from $a^{\bar{\mu}\bar{\nu}}$, as we shall 342 see below. It is worth noting that Klein did not start from a variational principle to 343 obtain his wave equation. He simply wrote a light-like wave equation. The hypothesis 344 that the wave function would be periodic with respect to the fifth coordinate x^5 per-345 mits to "project" equation (9) to obtain the KG wave equation.¹² See Appendix C 346 for an explanation of the use of periodicity condition in the context of de Broglie's 347 work. 348

How did Klein justify the analogy with light? In Klein [1991] the author recol-349 lected: '[...] for some time I had played with the idea that waves representing the 350 motion of a free particle had to be propagated with constant velocity, in analogy with 351 light waves – but in a space of four dimensions – so that the motion we observe is a 352 projection on our ordinary three-dimensional space of what is really taking place in 353 four-dimensional space. [emphasis added]' ([Klein, 1991]; p. 108). The introduction of 354 the symmetric tensor $a^{\mu\bar{\nu}}$ served this specific purpose. Klein's conviction was enforced 355 by the fact that in the long wavelength limit equation (9) reduces to the eikonal 356 equation for light rays. As a consequence, Klein imposed that in the semi-classical 357 limit the four-dimensional motion of charged particles with mass m in the presence 358 of a gravitational and electromagnetic field should be described by five-dimensional 359 null-geodesics of the following differential form: 360

$$d\hat{\sigma}^2 = a_{\bar{\mu}\bar{\nu}}dx^{\bar{\mu}}dx^{\bar{\nu}} = \frac{1}{m^2c^2}d\theta^2 + ds^2$$
(10)

([Klein, 1984]; p. 17) and showed that the correspondent geodesic equation is equivalent to the four-dimensional Lorentz equation. It seems that Klein introduced a different metric for the microscopic world, $a_{\bar{\mu}\bar{\nu}}$, whose components can be obtained

¹¹Given a scalar field ϕ of mass m, the KG equation is $\Box \phi = \frac{m^2 c^2}{\hbar^2} \phi$.

¹²Klein and all the authors we consider in the present paper were convinced, at that time, that the relativistic wave equation for the electron would be the KG equation, instead of Dirac's equation. It is worth remembering that Pauli matrices were introduced in the same year [Pauli, 1927] and that the Dirac's equation would be published one year later [Dirac, 1928].

from equation (10), namely:

$$a_{\mu\nu} = g_{\mu\nu} + \frac{e^2}{m^2 c^4} A_{\mu} A_{\nu} \qquad a_{\mu5} = \frac{e^2}{m^2 c^3 \beta} A_{\mu} \qquad a_{55} = \frac{e^2}{m^2 c^4 \beta^2} \quad , \tag{11}$$

and which is quite unlike the space-time metric $\gamma_{\bar{\mu}\bar{\nu}}$, cf. equation (11) with (4) and (3), but he made no comments on this choice. It is worth noting that the particle's mass m and its charge e are hidden in the expressions of $a_{\bar{\mu}\bar{\nu}}$ tensor.

To show the correspondence between five-dimensional null-geodesics and fourdimensional motion of charged particles, Klein considered the corresponding Lagrangian picture, by projecting the equations of motions obtained by varying the Lagrangian $L = \frac{1}{2} a_{\bar{\mu}\bar{\nu}} \frac{dx^{\bar{\mu}}}{d\hat{\lambda}} \frac{dx^{\bar{\nu}}}{d\hat{\lambda}}$, where $\hat{\lambda}$ is an arbitrary parameter. One of the five resulting Euler-Lagrange equations states that the momentum conjugated to the coordinate x^5 is conserved, while the other four equations are equivalent to the Lorentz equation for an electron¹³ (charge q = -e):

$$mc\left(\frac{d}{d\tau}\left(g_{\mu\nu}u^{\nu}\right) - \frac{1}{2}\partial_{\mu}g_{\rho\nu}u^{\rho}u^{\nu}\right) = -\frac{e}{c}\left(\partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}\right)u^{\nu},\qquad(12)$$

where the four-dimensional proper time τ is defined by $d\tau = \sqrt{-ds^2}$, and the fourvelocity of the particle is defined by $u^{\mu} = \frac{dx^{\mu}}{d\tau}$. The analogy with light forced Klein to look for a correspondence between five-dimensional null-geodesics and four-dimensional paths: this conclusion would be criticized by de Broglie.

Before going on, it is worth noting that equation (12) can be obtained, as Klein 379 did, without fixing the constant¹⁴ β introduced in (5). In his first paper, Klein decided 380 to set $\beta = \frac{e}{c}$ and consequently the value of α must be $\alpha = \frac{16\pi G}{e^2 c^2}$. In his second paper [Klein, 1926b], a brief communication to *Nature*, it seems that Klein had changed 381 382 his mind about the role of null-geodesics. In fact he explicitly referred to 'the equa-383 tion of geodetics' ([Klein, 1926b]; p. 516) of the line element¹⁵ $d\sigma^2$. Furthermore, 384 he suggested to start from the new Lagrangian $L' = \frac{m}{2} \gamma_{\bar{\mu}\bar{\nu}} \frac{dx^{\bar{\mu}}}{d\tau} \frac{dx^{\bar{\nu}}}{d\tau}$, where the $a_{\bar{\mu}\bar{\nu}}$ tensor has disappeared, and the mass and the presence of the proper time τ indicate 385 386 that Klein did not refer to null-geodesics.¹⁶ This brief communication is important, 387 because Klein noted that the quantization of the momentum along the periodic fifth 388 dimension¹⁷ of finite size l could have been connected with the quantization of the 389 electric charge. In fact the momentum's quantization along the fifth dimension forces 390 the size l to assume a precise value: 391

$$l = \frac{hc\sqrt{2\kappa}}{e},\tag{13}$$

¹⁴See Appendix C for technical details in the context of de Broglie's work.

 $^{^{13}\}mathrm{Technical}$ details of the equivalence are given in Appendix A.

¹⁵In this brief communication Klein introduced a different notation and decided to set $\alpha = 1$ from the beginning and consequently $\beta = \sqrt{\frac{16\pi G}{c^4}}$: this simply means that now the fifth coordinate has a dimension of length.

¹⁶From a modern point of view, even in the massive case, the Lagrangian L' should be written by introducing the arbitrary parameter $\hat{\lambda}$ The proper time τ can be introduced because the ratio $\frac{d\hat{\lambda}}{d\tau}$ is constant, as we shall show in Appendix C, discussing de Broglie's approach. We suppose that Klein underlined implicitly that he did not consider null-paths any more.

¹⁷The momentum connected with the quantization of the electric charge is p_5 , the momentum conjugated to the fifth dimension, namely $p_5 = \frac{\partial L'}{\partial (dx^5/d\tau)}$.

where $\kappa = \frac{8\pi G}{c^4}$. As we will see, as far as we know, neither de Broglie nor Rosenfeld fixed explicitly either of both parameters and they also did not make explicit considerations on the size of the fifth dimension.

395 3.2 De Broglie's contribution

As mentioned in the introduction, during his stay in Paris Rosenfeld drew de Broglie's attention to Klein's approach. From de Broglie's point of view, the analogy with light
was not the correct perspective to describe the path of massive particles. In order to explain the conclusion reached by de Broglie, we emphasize again that Klein, de
Broglie and Rosenfeld developed the five-dimensional Universe for different reasons.

De Broglie's paper analyses the features of the five-dimensional approach from 401 two distinct perspectives: the classical and the quantum point of view. In the first 402 part of de Broglie's paper, the author described how the most attractive advantage 403 of the classical five-dimensional approach would reside in the fact that it allowed to 404 geometrize all the forces known at that time, i.e. the gravitational and the electro-405 magnetic forces. The author made an analogy between Einstein's approach and the 406 five-dimensional construction. De Broglie interpreted Einstein's theory as a geomet-407 rical description of the gravitational force and Kaluza's approach as an extension 408 of this geometrical description to Maxwell's theory¹⁸: 'The main consequence of the 409 introduction of the equivalence principle is that the metaphysic notion of force in the 410 theory of gravitation disappears. The path followed by a point particle in a gravita-411 tional field can be defined, thanks to Einstein's conceptions, as the geodesic line of 412 the space-time. [...] The success of this beautiful interpretation of the gravitational 413 field temptingly suggests to throw out the concept of force from the Physics, in order 414 to replace it with the concept of geometry.' ([Louis de Broglie, 1927b]; p. 65). 415

In the second part of the paper, de Broglie introduced the description of the 416 quantum behaviour of matter using wave/particle duality. From this perspective, 417 there are no forces associated to the particles' wave function, hence neither geomet-418 rical description nor analogy with light was needed. De Broglie explicitly stated that 419 With the present state of our knowledge it seems that all the forces of which we 420 are aware can be reduced to only two: the gravitational and electromagnetic forces.' 421 ([Louis de Broglie, 1927b] p. 65). It is worth noting that the quantum force concept 422 emerged with the introduction of quantum fields. Unlike Klein, de Broglie introduced 423 a wave equation describing quantum particles' dynamics, i.e. the KG equation, in four 424 dimensions: in the geometrical optics approximation the wave's rays would follow the 425 classical trajectories for massive particles. Hence a five-dimensional generalization of 426 the KG equation would not require any analogy with light. It is important to stress 427 that de Broglie did not use any variational principle to describe the wave's dynamics. 428 With this premise in mind we first consider de Broglie's approach in more detail. 429

⁴³⁰ De Broglie briefly reviewed Klein's approach and introduced the line element (1)
 ⁴³¹ with Klein's Ansatz that now we rewrite here for convenience:

where

$$d\sigma^2 = \alpha d\theta^2 + ds^2, \qquad (14)$$

$$d\theta = dx^5 + \beta A_\mu dx^\mu \quad ; \quad g_{\mu\nu} \quad = \quad \gamma_{\mu\nu} - \frac{\gamma_{5\mu}\gamma_{5\nu}}{\alpha} \quad ; \quad ds^2 = g_{\mu\nu}dx^\mu dx^\nu \quad (15)$$

 $^{^{18}}$ Here and in the following, we present an English translation of some parts of the original paper, written in French.

(We adapted de Broglie's notation changing the symbols he used). Let the values of α and β be unfixed for the moment. De Broglie's choice shall be analysed after equation (23).

At this point, de Broglie's and Klein's paths separate. As we said, de Broglie did not consider any analogy with light, hence he studied the geodesic equations in five dimensions for massive particles. Like Klein, the key idea is that our world would be a projection onto a four-dimensional manifold of what happens in the fivedimensional Universe. The four-dimensional geodesic equation is obtained by the following variational principle¹⁹:

$$\delta S_4 = 0 \quad \Rightarrow \quad \delta \int_O^M d\tau = 0 \,, \tag{16}$$

where O and M are 'two fixed points of the world line' ([Louis de Broglie, 1927b]; p. 69). De Broglie considered its natural generalization in five dimensions:

$$\delta S_5 = 0 \quad \Rightarrow \quad \delta \int_O^M d\hat{\tau} = 0 \,, \tag{17}$$

where we introduced the notation $d\hat{\tau} = \sqrt{-d\sigma^2}$. The geodesic equations following from (17) are equivalent to the five-dimensional equations obtained by Klein with the help of the $a_{\bar{\mu}\bar{\nu}}$ tensor he introduced in his first paper,²⁰ and their four-dimensional projection reproduce equation (12). In order to obtain the correct Lorentz equations, de Broglie set

$$\alpha \frac{d\theta}{d\tau} = -\frac{e}{\beta c} \frac{1}{mc} , \qquad (18)$$

underlining the importance of this equation. Indeed, from de Broglie's point of view, equation (18) suggests a geometrical interpretation of the ratio $\frac{e}{m}$. Let's consider, following de Broglie, 'a coordinate line x^5 ' ([Louis de Broglie, 1927b]; p. 68) and using $d\tau = \sqrt{-ds^2}$ and $d\hat{\tau} = \sqrt{-d\sigma^2}$ we rewrite equation (14) as follows:

$$d\hat{\tau}^2 = d\tau^2 + |\alpha| \, d\theta^2 \,. \tag{19}$$

We use $|\alpha|$, because de Broglie set $\alpha < 0$, a choice that we shall discuss after equation (23). 'Let us represent, on a point P of this coordinate line, a part of a plane π inclined with respect to the x^5 direction, which represents a little portion of the fourdimensional hypersurface $x^5 = const$. passing through the point P. Let \overline{PQ} be an element of a world line of length $d\hat{\tau}$ and let \overline{PS} and \overline{PR} be its projections along the x^5 direction and orthogonal to the x^5 direction respectively. From equation (19) it follows that

$$\overline{PS} = \sqrt{|\alpha|} d\theta \; ; \qquad \overline{PR} = d\tau \; . \tag{20}$$

¹⁹Because of our mostly-plus signature, the four-dimensional action for a point particle involves the proper time τ .

 $^{^{20}}$ See Appendix C for a detailed explanation of the original derivation. As we said, Klein was certainly aware of this fact, because he changed his own approach to the geodesics in the brief communication to *Nature*. It is worth noting that de Broglie never cited Klein's *Nature* paper.

[...] the tangent of the angle \widehat{QPR} , namely $\frac{\sqrt{|\alpha|}d\theta}{d\tau}$, is proportional to the ratio $\frac{e}{m}$ 459 where e and m are the charge and the mass of the particle of which \overline{PQ} is the 460 element of the world line. Hence the world line of every moving object makes the 461 same angle with the direction x^5 at each point, which angle is straight if the electric 462 charge is zero.' ([Louis de Broglie, 1927b]; p. 68).²¹ This result supported de Broglie's 463 conviction that the five-dimensional Universe could provide a geometrical description 464 for all of the known physical concepts. Rosenfeld would continue to use this idea, as 465 we shall see in the discussion after equation (59). 466

⁴⁶⁷ De Broglie asked himself what the exact form of the action S_5 to be varied would ⁴⁶⁸ be in order to obtain a five-dimensional generalization of the four-dimensional massive ⁴⁶⁹ particle's action. De Broglie stressed that he wanted to obtain, in the case of zero ⁴⁷⁰ charge, the usual action $S_4 = -mc \int_O^M d\tau$ ([Louis de Broglie, 1927b]; p. 70) and he ⁴⁷¹ proposed that the five-dimensional particle's action should be²²:

$$S_5 = -\mathcal{I} \int_O^M d\hat{\tau} \,, \tag{21}$$

472 where the quantity \mathcal{I} satisfies the following relations

$$\mathcal{I}\alpha \frac{d\theta}{d\hat{\tau}} = -\frac{e}{c\beta}, \qquad \mathcal{I}\frac{d\tau}{d\hat{\tau}} = mc,$$
 (22)

⁴⁷³ and has the following form:

$$\mathcal{I} = \sqrt{m^2 c^2 - \frac{e^2}{\alpha \beta^2 c^2}} \,. \tag{23}$$

The invariant \mathcal{I} needs some comments, connected with de Broglie's choice of α 's and β 's values. De Broglie implicitly set

$$\alpha\beta^2 = -\frac{16\pi G}{c^4} , \qquad (24)$$

from the beginning of his paper. As a consequence, $\mathcal{I}_{dB} = \mathcal{I}\left(\alpha\beta^2 = -\frac{16\pi G}{c^4}\right) = \text{is}$ a real constant:

$$\mathcal{I}_{dB} = \sqrt{m^2 c^2 + \frac{e^2 c^2}{16\pi G}},$$
(25)

and comparing S_4 and S_5 , de Broglie suggested that it should be interpreted as the modulus of the five-dimensional momentum $P_{\bar{\mu}}$ for charged particles, defined in analogy with the four-dimensional momentum $p_{\mu} = mcg_{\mu\nu}\frac{dx^{\nu}}{d\tau}$ for uncharged particles in four dimensions, namely $P_{\bar{\mu}} = \gamma_{\bar{\mu}\bar{\nu}}\mathcal{I}_{dB}\frac{dx^{\bar{\nu}}}{d\hat{\tau}}$. To be more explicit, referring to the geometrical picture discussed above, de Broglie asserted that relations (22) should

²¹With the choice $\alpha > 0$, the ratio would define the hyperbolic tangent of the angle.

 $^{^{22}}$ We skip over some technical details. See Appendix C for de Broglie's original proof that S_5 reduces to S_4 in the case of null charge.

be interpreted as the tangent and orthogonal components of the five-dimensional 483 momentum $P_{\bar{\mu}}$ with respect to the fifth direction x^5 ([Louis de Broglie, 1927b]; p. 484 70, note (1)). We will return to this interpretation discussing Rosenfeld work, see 485 the discussion after equation (59). Equation (24) means that unlike Klein, de Broglie 486 imposed that the fifth dimension would be a time-like coordinate, because from equa-487 tion (24) it follows $\gamma_{55} = \alpha < 0$. De Broglie made no explicit comment on the time-like 488 character of the fifth dimension. As we shall see, Klein noted that this choice was 489 inconsistent with other demands of the model. Rosenfeld would be strongly influ-490 enced by de Broglie's ideas, but he was aware of this inconsistency. After having 491 specified this fundamental difference between the two approaches, let us now return 492 to de Broglie's considerations. 493

After having established that the Lorentz equation (12) can be obtained by varying²³ S_5 , de Broglie declared: '*The notion of force has been banned completely from Mechanics.*' ([Louis de Broglie, 1927b]; p. 70), emphasizing his original aim. As a consequence he proposed the following wave equation as a generalization of Schrödinger wave equation, instead of (9), namely

$$\gamma^{\bar{\mu}\bar{\nu}}\nabla_{\bar{\mu}}\partial_{\bar{\nu}}\Psi = \frac{4\pi^2}{h^2}\mathcal{I}_{dB}^2\Psi, \qquad (26)$$

where now the covariant derivative is correctly contracted with the metric. Equation (26) could resemble a KG equation in five dimension, where $\frac{\mathcal{I}_{dB}}{c}$ plays the role of the mass in five dimensions, because it is a real quantity. It is worth noting that the identification of Ψ as a wave function prevents the identification of \mathcal{I}_{dB} with a mass term in the sense of modern field theory. Using the fact that the action S_5 can be rewritten as follows

$$S_{5} = -\int_{O}^{M} \frac{e}{c\beta} dx^{5} + \frac{e}{c} \int_{O}^{M} A_{\mu} dx^{\mu} - mc \int_{O}^{M} d\tau , \qquad (27)$$

⁵⁰⁵ de Broglie could show that equation (26) is equivalent to the four-dimensional ⁵⁰⁶ KG equation for massive particles, which reduces to Schrödinger equation in the ⁵⁰⁷ non-relativistic limit. In order to demonstrate his claim, de Broglie introduced the ⁵⁰⁸ geometrical optics approximation, writing the five-dimensional wave function Ψ as

$$\Psi = Ce^{\frac{i}{\hbar}S_5} = f(x, y, z, t)e^{\frac{i}{\hbar}\frac{ex^5}{c\beta}}$$
(28)

([Louis de Broglie, 1927b]; p. 72), where C is a constant and S_5 is the five-dimensional 509 action defined in (27). It is worth noting that De Broglie considered S_5 as an Hamilto-510 nian action. This means that he interpreted the five-dimensional action as a "Jacobi 511 function". As we will see, De Donder will be more explicit on this fact. At this 512 point, De Broglie expressed his opinion on the analogy with light introduced by the 513 Swedish physicist: 'O. Klein writes the equation (26) without the second member, 514 and he concludes that the world-lines must be null-geodesics; it is in our opinion that 515 the second term of (26) is fundamental and that the world-lines are still geodesics, 516 but not null-geodesics' ([Louis de Broglie, 1927b], p. 72; we modified the number of 517 the cited equation in order to fit with our numerical order). 518

⁵¹⁹ Before going on we return to the question of the fifth dimension's size, which ⁵²⁰ was never calculated by de Broglie. Indeed, the author commented on the size of the ⁵²¹ fifth dimension like this: 'The variations of the fifth coordinate completely escape ⁵²² our senses [...] two points that differ only for the value of the fifth coordinate are

²³See Appendix C.

⁵²³ indistinguishable from our point of view' ([Louis de Broglie, 1927b]; p. 67). But ⁵²⁴ from these observations, de Broglie inferred, like Klein, that the components of the ⁵²⁵ metric $\gamma_{\bar{\mu}\bar{\nu}}$ must be independent from the fifth coordinate and that 'the only humanly ⁵²⁶ possible transformations have the following form:

$$x'^{\mu} = f^{\mu} \left(x^0, \, x^1, \, x^2, \, x^3 \right) \quad , \tag{29}$$

⁵²⁷ ([Louis de Broglie, 1927b]; p. 67). If de Broglie would have chosen $\alpha\beta^2 = 2\kappa$, i.e. a ⁵²⁸ space-like dimension, he would have been able to read off the size of the compact ⁵²⁹ dimension. Indeed, after noting that²⁴ $\tilde{x}^5 = \sqrt{\alpha}x^5$ has dimensionality of [length]¹, ⁵³⁰ the dependence on the fifth dimension in (28) can be rewritten as

$$\frac{i}{\hbar}\frac{ex^5}{c\beta} = \frac{i}{\hbar}\frac{e}{c\sqrt{\alpha\beta}}\sqrt{\alpha}x^5 = \frac{i}{\hbar}\frac{e}{c\sqrt{2\kappa}}\sqrt{\alpha}x^5 = i\frac{\tilde{x}^5}{\tilde{l}}, \qquad (30)$$

where $\tilde{l} = \frac{\hbar c \sqrt{2\kappa}}{e}$ is Klein's length (13) divided by 2π , showing that Klein's length determines the periodicity.

⁵³³ De Broglie was very impressed by equation (26) and he concluded his paper with ⁵³⁴ the following remark: 'For studying the problem of matter and of its atomic structure ⁵³⁵ deeply, it would be necessary to perform a systematic analysis of the five-dimensional ⁵³⁶ Universe's point of view that seemed to be more promising than Weyl's approach. If ⁵³⁷ we understand how to interpret correctly the role played by the constants e, m, c,⁵³⁸ \hbar and G in equation (26), we will have finally grasped one of the most mysterious ⁵³⁹ secret of Nature.' ([Louis de Broglie, 1927b] p. 73).

Klein's answer to the question of null-geodesics arrived immediately [Klein, 1927a]. 540 He noted that in equation (26) de Broglie used the metric $\gamma^{\mu\nu}$ instead of his "arti-541 ficial" tensor $a^{\bar{\mu}\bar{\nu}}$: inserting the components of $a^{\bar{\mu}\bar{\nu}}$ in (26), Klein showed that the 542 equations (26) and (9) were equivalent. The fact is not surprising, because the parti-543 cle's mass is hidden in the expression of the $a^{\bar{\mu}\bar{\nu}}$ tensor.²⁵ Klein also noted that the 544 condition on the parameters $\alpha \beta^2 = 2\kappa$ was incompatible with the choice of a time-like 545 fifth dimension.²⁶ But he concluded the brief communication with a positive com-546 ment on de Broglie's assertion: '...this error has no influence on de Broglie's result 547 [emphasis added]²⁷ ([Klein, 1927a]; p. 243). It is worth noting that in his subsequent 548 papers Klein would have stressed the need to introduce a space-like fifth dimension²⁸ 549 ([Klein, 1927b]; p. 206, footnote *). Notwithstanding, after de Broglie's paper, Klein 550 abandoned explicitly the analogy with light. 551

552 3.3 De Donder's lectures on gravitation

⁵⁵³ Neither Klein nor de Broglie tried to obtain their wave equation, in the works we ⁵⁵⁴ analysed so far, using a unified variational principle. In fact they introduced only

²⁴Remember that de Broglie choose a negative value for α . We suppose that for this reason he never noted this fact.

²⁵See discussion after equation (10).

 $^{^{26}\}mathrm{In}$ Appendix B we will analyse Klein's claims in more detail.

 $^{^{27}}$ Klein assertion was referred to the fact that irrespective of the nature of the fifth coordinate, after having used the periodicity condition, the term with the Newton constant in (26) disappears and it reduces to the KG equation. See Appendix C, the discussion after equation (C.19) for a detailed explanation.

²⁸See Appendix B for technical details.

the particle's Lagrangian in order to describe the classical particle's dynamics.²⁹ The 555 Belgian physicist Théophile De Donder was an early supporter of variational princi-556 ples, developing the purely formal parts of the calculus of variations and analysing 557 e.g. the effect of transformations of coordinates and parameters upon what he called 558 "invariants" and upon other expressions which occur in the theory of the variational 559 calculus [De Donder, 1930]. As we shall see, De Donder's "invariants" would corre-560 spond to our modern Lagrangian density. He tried also to derive WM from a unified 561 variational principle. He did not consider multidimensional world, because he was 562 satisfied to write a unified Lagrangian involving the gravitational field, the Maxwell 563 field and a Lagrange function for the quantum particle. De Donder tried to present 564 a coherent framework for relativistic Lagrangian dynamics in the context of curved 565 spaces, and he was one of the first to note the role of the HJ equation as constraints 566 in this context. In his first paper, Rosenfeld mainly followed De Donder's approach 567 to introduce the wave function in the five-dimensional Universe, as we shall see later. 568 During the Spring of 1926, De Donder gave a series of lectures at the MIT. In these 569 lectures, which would be published the following year [De Donder, 1927a], the Belgian physicist gathered together all the results he had just published in the *Comptes* 571 *Rendus* journal. The lectures contain all the original references, with an advantage: 572 *Comptes Rendus* publications were often brief communications, whereas the lectures 573 gave a complete overview of De Donder's point of view. For this reason we will refer 574 to his MIT lectures. We stress that this paragraph is a brief analysis of the ideas that 575 influenced Rosenfeld. A deeper understanding of De Donder's methods goes beyond 576 the goals of the present paper. 577

The Belgian physicist tried explicitly to apply GR to the microscopic world. At 578 the end of the first lecture, the general introduction, De Donder wrote: 'We then 579 say a few words about the mysterious quantum. To shed some light on this obscure 580 physical entity, we shall deduce at first from relativistic electrodynamics expressed by 581 means of points in space-time, the dynamics of an atomic or molecular system of any 582 number of degrees of freedom. We shall then devise a general method of quantization 583 in space-time, which we shall apply to the quantization of the point electron and 584 to that of *continuous* systems: It will be shown that this quantization is a logical 585 consequence of our gravific theory [...]³⁰ ([De Donder, 1927a]; p. 8). 586

This comment is important for two reasons. First, it emphasized again that the 587 problem of reconciling quantum physics and GR was considered early in the history 588 of quantum physics. Secondly, De Donder developed his approach during the birth of 589 QM and it is a "spurious" approach in the following sense. Before 1925 the quantiza-590 tion of a system was performed using Epstein-Sommerfeld-Wilson rules and a system 591 like 'the point electron', as De Donder referred to, would follow a classical trajec-592 tory. He agreed with this interpretation and in this sense, from our point of view, 593 his approach belongs to the old quantum theory. But De Donder knew Schrödinger 594 papers and he explicitly stated that he was looking for new quantization rules that 595 should be compatible with the curved space-time of Einstein theory. These rules 596 would have to reproduce, in his opinion, the general relativistic generalization of 597 Schrödinger's equation.³¹ This means that with the phrase 'general method of quan-598 tization in space-time' De Donder intended a procedure to obtain a wave equation 599 for the wave function ψ , living on a curved background. As far as we know, De Don-600 der never referred to ψ as a field. For this reason we could say that De Donder was 601 looking for a "General Relativistic Quantum Mechanics" (GRQM). 602

 $^{^{29}}$ As we shall note in the next section, Klein's last paper would contain a five-dimensional variational principle to derive WM ([Klein, 1927b]; p. 201), which is slightly different from Rosenfeld's variational principle.

 $^{^{30}}$ De Donder used the old term 'gravific theory' instead of 'gravitational theory'.

 $^{^{31}\}mathrm{Once}$ again the reference was to the KG equation.

In WM a key ingredient of the quantization procedure was the imposition of 603 boundary conditions for the wave function. As far as we know, De Donder never 604 considered any boundary conditions explicitly. As we will see, his method was based 605 on a unified variational principle, but De Donder's ψ was treated, from our point of 606 view, classically. This means also that, from the modern field theoretic point of view, 607 he did not consider any quantum feature of the fields. Lastly, it is worth noting that 608 De Donder was not alone in believing that quantization rules could be derived in the 609 context of some unknown classical theory. Einstein, for example, would look for a 610 classical field theory (Einheitliche Feldtheorie) for the rest of his life [Pais, 1982]. We 611 do not know why De Donder was convinced of this idea, but because of the absence 612 of a discussion on the wave function's boundary conditions, as we shall discuss after 613 equation (45), the unified variational principle seemed not to require any modification 614 of GR. For this reason, De Donder thought that the quantization rules should have 615 been a consequence of GR principles, as he stated in the introduction cited 616 above. This attitude is consistent with the claim that De Donder belongs to the group 617 of authors who were convinced of GR supremacy. This conviction is confirmed by the 618 last sentence of the general introduction to his MIT lectures: 'Once more relativity 619 unfolds the great physical drama of the universe clad in an immutable form bearing 620 the stamp of eternal laws.' ([De Donder, 1927a]; p. 8). This means also that from 621 a modern point of view, in his approach De Donder did not consider any quantum 622 effect on the gravitational field. This fact was common to almost all the pre-1930 623 works: as far as we know Rosenfeld's approach was the only exception. 624

We introduce some technical details in order to understand how De Donder tried to harmonize WM with GR. The tenth lecture is dedicated to the 'Relativistic Quantization', and it started from the classical dynamics of a charged particle in GR, i.e. the 'point-electron'. The dynamics is described by the Euler-Lagrange equations obtained using the following Lagrangian³² ([De Donder, 1927a]; p. 90):

$$L_{DD}(x; u) = \frac{mc}{2} g_{\mu\nu} u^{\mu} u^{\nu} - \frac{e}{c} A_{\mu} u^{\mu}, \qquad (31)$$

where $u^{\mu} = \frac{dx^{\mu}}{d\tau}$, τ is the proper time, and the tangent vector satisfies the following constraint:

$$g_{\mu\nu}u^{\mu}u^{\nu} = -1.$$
 (32)

⁶³² Using L_{DD} , De Donder was able to define the conjugate momenta as $p_{\mu} = \frac{\partial L_{DD}}{\partial u^{\mu}} =$ ⁶³³ $mcg_{\mu\nu}u^{\nu} - \frac{e}{c}A_{\mu}$, and the Hamiltonian $H = p_{\mu}u^{\mu} - L_{DD}$ reads:

$$H = \frac{1}{2mc} \left(p_{\mu} + \frac{e}{c} A_{\mu} \right) \left(p^{\mu} + \frac{e}{c} A^{\mu} \right) \,. \tag{33}$$

The constraint $g_{\mu\nu}u^{\mu}u^{\nu} = -1$ is equivalent to the relation $H = -\frac{1}{2}mc$, i.e. the reduced HJ equation for a point particle, which De Donder called 'Jacobian equation'. Finally, by using equation (33), the constraint assumes the following form ([De

 $^{^{32}}$ The "Lagrangian" used by De Donder had the dimensions of a Lagrangian divided by a velocity and the same happens for the following "Hamiltonian" (33), but we will call them Lagrangian and Hamiltonian as well.

⁶³⁷ Donder, 1927a]; p. 91, Eq. (10)):

$$g^{\mu\nu}\left(\frac{\partial S}{\partial x^{\mu}} + \frac{e}{c}A_{\mu}\right)\left(\frac{\partial S}{\partial x^{\nu}} + \frac{e}{c}A_{\nu}\right) + m^{2}c^{2} = 0 \qquad , \qquad \frac{\partial S}{\partial x^{\mu}} = p_{\mu}, \qquad (34)$$

where S is the Jacobi function of classical mechanics. Before going on, we point out that De Donder was aware of the following fact. Using $S_4 = -mc \int_O^M d\tau$ as action for the free point-particle, the Lagrangian approach could be performed introducing an arbitrary parameter λ and rewriting S_4 as follows:

$$S_4 = \int_O^M L d\hat{\lambda} = \int_O^M \sqrt{-\gamma_{\bar{\mu}\bar{\nu}}} \frac{dx^{\bar{\mu}}}{d\hat{\lambda}} \frac{dx^{\bar{\nu}}}{d\hat{\lambda}} d\hat{\lambda} .$$
(35)

In this case, a Legendre transform would produce a null Hamiltonian, i.e. the constraint H = 0.

At this point De Donder introduced a wave function associated to the electron, namely $\psi(\tau, x)$, a function of the spatial coordinates x and of the proper time τ . In the MIT lectures, the author made no explicit discussion neither on the mathematical feature of the wave function nor on its physical interpretation. He implicitly identified it with Schrödinger's wave function, when considering a single electron. In fact, De Donder imposed the following Ansatz for the wave function ([De Donder, 1926a]; p. 91):

$$\psi = e^{kS} \quad \text{i.e.} \quad S = \frac{1}{k} \log(\psi) ,$$
(36)

where the Jacobi function $S(\tau, x)$ depends on the spatial coordinates and on the 651 proper time. At the beginning k is an unknown constant, but in the end, in order 652 to match his wave equation with Schrödinger equation, he would choose $k = \frac{i}{\hbar}$. 653 De Donder made no comment on the fact that with this choice both ψ and the 654 log-function in equation (36) turn into complex functions. As a consequence of the 655 fact that he left k undetermined, he would not use the complex conjugate as we 656 shall do in equation (42). De Donder will use the correct notation in his book on 657 Variational Calculus [De Donder, 1930]. If $k = \frac{i}{\hbar}$, the Ansatz (36) corresponds to 658 the correct geometrical optics approximation. It is worth noting that this procedure 659 is very similar to Klein's approach. In fact, this procedure was the common way 660 to introduce a wave equation for a "quantum" particle in the mid 1920s. Unlike 661 Klein, from De Donder's point of view it was not necessary to unify all forces with 662 a five-dimensional Lagrangian. Indeed, De Donder was satisfied with a unified action 663 principle. Unlike Klein, he looked from the beginning for an action principle in four 664 dimensions, with the help of relativistic Hamiltonian dynamics. 665

After having introduced the Jacobi function $S(\tau, x)$, in order to obtain the reduced HJ equation $H = -\frac{1}{2}mc$, the reducibility condition reads:

$$\frac{\partial S}{\partial \tau} = \frac{1}{2}mc. \tag{37}$$

Integrating (37), De Donder wrote the Jacobi function in the following form:

$$S = \frac{1}{2}mc\tau + S_0\left(x^0, x^1, x^2, x^3\right), \qquad (38)$$

that will play an important role for Rosenfeld, as we shall see in the next section.

Thanks to definition (36) and using equation (37), the author was able to write ([De Donder, 1926a]; p. 91):

$$\frac{\partial S}{\partial \tau} = \frac{\hbar}{i} \frac{1}{\psi} \frac{\partial \psi}{\partial \tau} , \qquad (39)$$

$$\frac{\partial S}{\partial x^{\mu}} = \frac{\hbar}{i} \frac{\partial_{\mu} \psi}{\psi} , \qquad (40)$$

$$\psi = \frac{\hbar}{i} \frac{\frac{\partial \psi}{\partial \tau}}{\frac{\partial S}{\partial \tau}} = \frac{\hbar}{i} \frac{2}{mc} \frac{\partial \psi}{\partial \tau} \,. \tag{41}$$

The conjugated wave function $\overline{\psi}$ satisfies the conjugated version of equations (39), (40) and (41).

Inserting (40) and (41) into (34), the HJ equation (34) can be rewritten in the following form:

$$J(\psi) \equiv -g^{\mu\nu} \left(\frac{mc}{2}\partial_{\mu}\psi + \frac{e}{c}A_{\mu}\frac{\partial\psi}{\partial\tau}\right) \left(\frac{mc}{2}\partial_{\nu}\overline{\psi} - \frac{e}{c}A_{\nu}\frac{\partial\overline{\psi}}{\partial\tau}\right) - m^{2}c^{2}\frac{\partial\psi}{\partial\tau}\frac{\partial\overline{\psi}}{\partial\tau} = 0.$$
(42)

In De Donder's approach equation (42) defines a functional $J(\psi)$, that is an invariant under all changes of variables, x^0, \ldots, x^3 ([De Donder, 1927a]; p. 92). The J functional plays a fundamental role for the author. From his point of view, with the introduction of the wave function ψ , the classical HJ equation (34) becomes a constraint for the new functional $J(\psi)$, i.e.

$$J\left(\psi\right) = 0\,,\tag{43}$$

and using this new functional De Donder was able to introduce what the author calls
 the relativistic quantization rule for curved space-time. After defining the following
 functional derivative:

$$\frac{\delta}{\delta\psi}J(\psi) = \frac{\partial J}{\partial\psi} - \partial_{\mu}\frac{\partial J}{\partial\partial_{\mu}\psi} + \cdots, \qquad (44)$$

the quantization rule reads: 'the variational derivative of the left-hand member of the Jacobian equation (43), with respect to ψ , shall vanish. Explicitly:

$$\frac{\delta}{\delta\psi}\left(\sqrt{-g}J\right) = 0 \,\,,\tag{45}$$

⁶⁸⁶ ([De Donder, 1927a]; p. 92).

Before going on, let us consider De Donder's variational principles in more detail. Lecture 5 of the MIT lectures is dedicated to 'The Fundamental Equations of the Gravific Field'. In order to obtain Einstein equations, De Donder considered the ⁶⁹⁰ following variational principle ([De Donder, 1926a]; p. 47):

$$\frac{\delta\left[\left(aR+b+\mathcal{L}_{m}\right)\sqrt{-g}\right]}{\delta g^{\mu\nu}}=0, \qquad (46)$$

where the functional derivative is defined as in equation (44) with ψ replaced by the 691 metric, R is the four-dimensional curvature scalar, a and b are arbitrary constants 692 (incidentally, the constant b plays the role of the Cosmological Constant A, but De 693 Donder did not comment on this fact), \mathcal{L}_m is an unspecified Lagrangian density for 694 the matter part of the theory, and the functional $(aR+b)\sqrt{-g}$, i.e. the Lagrangian 695 density, is named 'the characteristic gravific function' ([De Donder, 1926a]; p. 47). 696 It seems that in these years De Donder preferred to introduce a variational principle 697 using Lagrangian densities instead of action functionals. De Donder himself stressed 698 this fact as follows, advocating a precise justification of the choice he made: 'The vari-699 ational principle, as we have presented it, is evidently a generalization of Hamilton's 700 principle, that is, equivalent to placing 701

$$\delta \int_{\Omega} \left(aR + b + \mathcal{L}_m \right) \sqrt{-g} d^4 x = 0 , \qquad (47)$$

Ω being a region of space-time at the boundaries of which the variations must vanish.
It is in order to avoid the use of four-dimensional space that we have preferred the above presentation.' [emphasis added] ([De Donder, 1926a]; p. 47). In his following works devoted on the developments of variational principles and their applications
[De Donder, 1930], the author will use both forms. Let us now consider again De Donder's approach to quantization procedure.

Why did De Donder call equation (45) 'a quantization rule'? The functional 708 derivative (44), introduced by De Donder, produces the usual equations of motion 709 for a charged scalar field and he showed that it reduces to the Schrödinger's equation 710 in the non relativistic limit and in the approximation of an electrostatic field. It is 711 worth noting that De Donder's ψ would not have the correct dimensionality to be 712 interpreted as the Schrödinger's wave equation, but De Donder made no comments 713 on this fact. For this reason he considered equation (45) as a quantization rule. In 714 this sense, for us, De Donder's approach belongs to the WM point of view: like Klein 715 he believed that writing a wave equation was a sufficient condition to describe the 716 quantum behaviour of a system. 717

Why did De Donder assert in his general introduction that this quantization rule 718 would be 'a logical consequence of our gravitational theory'? In order to answer this 719 question, firstly we note that from a modern point of view, De Donder's approach 720 is of course a classical approach, because it is equivalent to a classical variational 721 principle for a field theory, though De Donder interpreted the "field" ψ as a wave 722 function. The absence of the integral in (45) was compensated by an ad hoc choice of 723 the functional derivative defined in (44). Secondly we remember that the first authors 724 that tried to quantize scalar fields were Klein and Jordan in 1927 [Jordan & Klein, 725 1927]. This means that the concept of quantum field was not already born and like 726 other authors De Donder was convinced that writing a wave equation for a system 727 was sufficient to quantize it. De Donder was convinced that GR could explain where 728 the quantization rules come from, because he obtained Schrödinger's wave equation 729 through the use of a variational principle, like Einstein's equations are obtained, only 730 from different action. Lastly, it is worth noting that by applying variational methods 731 without imposing commutation relations for the fields, the apparatus of GR seems not 732 to require any modification. For these reasons, De Donder made the following remark, 733 in order to emphasize his interpretation of the approach: 'We have thus shown that 734

⁷³⁵ the quantization of the point electron can be deduced from Einstein's gravitational ⁷³⁶ theory by means of an absolute extremal.' ([De Donder, 1927a]; p. 95).

⁷³⁷ Before going on, we make the following remark on De Donder's functional. Unlike ⁷³⁸ Klein, who considered a real scalar field in five dimensions, De Donder wrote a sort ⁷³⁹ of Lagrangian density for a charged scalar field. More precisely, using relation (41) ⁷⁴⁰ the J functional reads:

$$J(\psi) = \frac{m^2 c^2}{4} \left[-g^{\mu\nu} \left(\partial_\mu \psi + \frac{i}{\hbar} \frac{e}{c} A_\mu \psi \right) \left(\partial_\nu \overline{\psi} - \frac{i}{\hbar} \frac{e}{c} A_\nu \overline{\psi} \right) - \frac{m^2 c^2}{\hbar^2} \overline{\psi} \psi \right].$$
(48)

The expression in the squared brackets resembles the Lagrangian density of a complex scalar field in the presence of an electromagnetic and a gravitational field, but neither ψ nor J would have the correct dimensionality to be interpreted as a scalar field and a density Lagrangian respectively. Unlike Klein's functional, De Donder's functional (48) would have the correct sign in order to be interpreted as a Lagrangian density Rocci, 2013].

747 **4 Rosenfeld's contributions**

Rosenfeld merged De Donder's and de Broglie's ideas using Klein's approach. He 748 explicitly cited all the authors we discussed in the preceding section. Like De Donder, 749 he considered the relativistic Jacobi function approach. Like de Broglie, he explicitly 750 inserted a mass term in the KG equation. Like Klein, he was aware of the fact that 751 the fifth dimension's character should be space-like. But the principal purpose of 752 Rosenfeld was to try to understand concretely how quantum effects should modify 753 the classical view in the presence of a gravitational field, at least in the weak field 754 approximation. 755

All of Rosenfeld's papers on this topic, [Léon Rosenfeld, 1927a,b,c], are authored 756 by Rosenfeld alone: to what extent were de Broglie and De Donder active collabora-757 tors in these articles? The influence of de Broglie and De Donder is stated explicitly 758 by the author himself. At the end of the introduction of his first paper, Rosenfeld 759 wrote: 'This work was completed under the direction of Mr. L. de Broglie and Mr. 760 Th. De Donder, who have never ceased to assist me with their advice, and have been 761 kind enough to communicate to me their works, even manuscripts; I am happy to 762 be able to express my deep appreciation to them here.' ([Léon Rosenfeld, 1927a]; p. 763 305). From the observations that we make in the rest of this paper, we can infer that 764 De Donder had an active part in Rosenfeld's paper. In particular, we shall see how 765 Rosenfeld followed De Donder's approach to introduce the wave equation in the con-766 text of a curved space-time, which permitted him to find a natural explanation of De 767 Donder's interpretation of the quantum wave amplitude. Furthermore, we shall infer 768 what precisely de Donder found attractive in Rosenfeld's five-dimensional Universe. 769 In his second and third communications, Rosenfeld supported with a physical expla-770 nation his first paper. Stimulated by De Donder's influence, Rosenfeld recognized that 771 he was using Bohr's correspondence principle. Unlike Rosenfeld, De Donder thought 772 that Rosenfeld's work was a proof of a new version of the correspondence principle. 773 which could be derived from Einstein's theory, and stressed that this principle should 774 have been a cornerstone or the 'gravitational wave mechanics' ([De Donder, 1927b]; 775 p. 506), i.e. a theory reconciling WM with Einstein's theory. 776

Rosenfeld's first paper [Léon Rosenfeld, 1927a] is a long and technical work and it
does not contain any physical interpretation of the choices he made. For this reason,
in Section 4.1 we shall pay more attention to the technical details of the Rosenfeld's
approach, explaining his results from the author's point of view. The second and the

third papers are shorter than his first contribution. In these articles the author clarified his technical choices from the physical point of view. We will analyse Rosenfeld's comments in Section³³ 4.2. At the end, in Section 4.3, we shall emphasize how these first articles influenced Rosenfeld's future work and we shall interpret the author's results from a modern point of view.

786 4.1 The quantum origin of a space-time metric

In the introduction to his first paper [Léon Rosenfeld, 1927a], written during his stay in Paris at the "Ècole normale supérieure", Rosenfeld formulated his main goals³⁴:

'The first part of this work is dedicated to the systematic study of the 789 five-dimensional universe considered by O. Klein, Th. De Donder and L. 790 de Broglie. We will show how the model of the five-dimensional universe is 791 satisfactory [...]. Generalizing Gordon's and Schrödinger's papers, we will 792 show how the introduction of the Ψ function of de Broglie-Schrödinger 793 permits us to combine in a unique variational principle, into the five-794 dimensional universe, the gravitational force, the electromagnetic force 795 and the quantum phenomena (the Ψ equation). [...] Finally, a formula will 796 be established to calculate the gravitational and electromagnetic potentials, 797 for a field slightly different from the Minkowskian field, as a function of 798 Ψ . The calculation will be developed for the case of a stationary charge 799 and for the case of a charge moving with constant speed. Comparing the 800 values obtained with the classical potentials, we find that the amplitude of 801 the Ψ function representing the charge must have a constant value inside 802 a finite volume and it must be zero outside of that volume: these results 803 can be well understood with the beautiful interpretation of the Ψ function 804 recently proposed by Mr. De Donder; quite to the contrary it appears to 805 be irreconcilable with the opinion of Mr. de Broglie, who believed that the 806 charge would be a point singularity of the Ψ function. [emphasis added]' 807 ([Léon Rosenfeld, 1927a]; p. 304-5). 808

We shall investigate only the first case proposed by Rosenfeld, i.e. the case of a 809 stationary massive charge, represented by a wave function, in order to investigate the 810 gravitational field produced by a quantum particle. Rosenfeld would consider a weak-811 field approximation, what he called 'a field slightly different from a Minkowskian 812 field³⁵. Rosenfeld would find that the quantum particle should be represented by 813 a localized wave function, which is non zero inside a finite volume, instead of a 814 point-like object, in contrast with de Broglie's point of view. This fact would enforce 815 De Donder's interpretation of the wave function's amplitude as representing a sort 816 of internal quantum force of matter. We will not discuss this interpretation, which 817 was based on the application of Rosenfeld-De Donder's approach to multi-particle 818 systems, because for this case Rosenfeld did not investigate the gravitational field. 819

Why did Rosenfeld consider a five-dimensional framework? The answer seems now almost trivial: the author studied Klein's work with de Broglie and was fascinated by its capability to describe in a unified framework GR and Maxwell's theory.

What was Rosenfeld's starting point? The answer is connected with his knowledge of De Donder's and de Broglie's works. Indeed, following De Donder, Rosenfeld

 $^{^{33}}$ The fourth of Rosenfeld's communication is an attempt to unify the preceding works.

 $^{^{34}}$ We present an English translation of some parts of the original paper, written in French, and then we comment on it. We omit the references of the original work.

³⁵Minkowskian field is the English translation of the French expression "champ de Minkowski" which was well understood and commonly used in that period as the vacuum space. See e.g. Solomon [1938] or Lichnerowitz in Pauli [1993].

started from the classical description of a single charged particle, and following Klein and de Broglie, he considered a five-dimensional space-time, with the usual coordinates $(x^0, x^1, x^2, x^3, x^5)$. The classical particle was described by a five-dimensional Jacobi function \bar{S} , namely

$$\bar{S}(x) = -\frac{e}{c\beta}x^5 + S_0\left(x^0, x^1, x^2, x^3\right), \qquad (49)$$

in analogy with De Donder's four-dimensional Jacobi function (38), that we rewrite here for convenience, namely:

$$S = \frac{1}{2}mc\tau + S_0\left(x^0, x^1, x^2, x^3\right) \,. \tag{50}$$

⁸³¹ Rosenfeld explicitly defined the fifth coordinate putting:

$$x^{5} = -\frac{mc^{2}\beta}{2e}\tau' \text{ ([Léon Rosenfeld, 1927a]; Eq. (5), p. 306),}$$
(51)

⁸³² specifying that ' β is a *universal constant.*' ([Léon Rosenfeld, 1927a]; p. 306). From ⁸³³ our point of view, the introduction of the fifth coordinate simply follows from the ⁸³⁴ comparison between De Donder's Jacobi function, equation (50), and de Broglie's five-⁸³⁵ dimensional Hamiltonian action for the charged particle, equation (27). Indeed, to ⁸³⁶ obtain equation (27), it is sufficient in (50) to set $S_0 = -\int_O^M \frac{e}{c} A_\mu dx^\mu - mc \int_O^M d\tau$. ⁸³⁷ About the size of the fifth dimension, Rosenfeld shared de Broglie's view. He observed

that from equation (49) it follows the invariance of x^5 with respect to the general transformation of coordinates $f(x^0, x^1, x^2, x^3)$ and concluded: 'Its invariance with respect to the transformations that we are able to perform explains why this fifth dimension escapes direct observations.' ([Léon Rosenfeld, 1927a]; p. 307). Like de Broglie, Rosenfeld did never discuss explicitly the size of the fifth dimension, though he would have been able to extract it.³⁶

The dynamics of classical charged particles is described by the HJ equation and Rosenfeld introduced his five-dimensional analogously. Following the author we note first that the new Jacobi function \bar{S} satisfies³⁷

$$\partial_5 \bar{S} = -\frac{e}{c\beta} \,. \tag{52}$$

Secondly, Rosenfeld used Klein's five-dimensional metric $\gamma_{\bar{\mu}\bar{\nu}}$ defined in the previous section, see equations (14) and (15), with the same convention, i.e. imposing the following choice for α and β : $\alpha\beta^2 = 2\kappa$. Lastly, with the help of the components of the inverse metric $\gamma^{\bar{\mu}\bar{\nu}}$, namely

$$\gamma^{\mu\nu} = g^{\mu\nu} , \qquad \gamma^{55} = \frac{1}{\alpha} + \beta^2 A_{\mu} A^{\mu} , \qquad \gamma^{5\mu} = -\beta A^{\mu} , \qquad (53)$$

the author is able to show how De Donder's four-dimensional HJ equation (34), namely

$$g^{\mu\nu} \left(\partial_{\mu}S_0 + \frac{e}{c}A_{\mu}\right) \left(\partial_{\nu}S_0 + \frac{e}{c}A_{\nu}\right) + m^2 c^2 = 0 , \qquad (54)$$

 $^{^{36}}$ See the discussion after equation (29).

³⁷Note that the combination $\frac{e}{c\beta}x^5$ has the dimension of an action.

can be rewritten in the following compact form ([Léon Rosenfeld, 1927a]; p. 307):

$$\gamma^{\bar{\mu}\bar{\nu}}\partial_{\bar{\mu}}\bar{S}\partial_{\bar{\nu}}\bar{S} = -\left(m^2c^2 - \frac{e^2c^2}{16\pi G}\right).$$
(55)

It is worth noting that equation (52) is the same relation that induced Klein to introduce a fifth coordinate: it suggests indeed that the electric charge could play the role of an extra momentum component, as recollected by Klein (see the beginning of Sect. 3.1), and permits to translate in the five-dimensional language the relativistic HJ equation for a particle moving in a combined electromagnetic and gravitational field.

⁸⁶⁰ Choosing $\alpha\beta^2 = 2\kappa$, Rosenfeld implicitly imposed $\alpha > 0$. As noted in the previous ⁸⁶¹ section, this means that, like Klein, Rosenfeld correctly introduced a space-like fifth ⁸⁶² dimension. Hence, the quantity \mathcal{I}^2 , see equation (23), assumes the following form:

$$\mathcal{I}_{Ros}^2 = m^2 c^2 - \frac{e^2 c^2}{16\pi G} , \qquad (56)$$

and it differs from de Broglie's \mathcal{I}_{dB} , see equation (25), because of the presence of 863 the minus sign. For an electron, the quantity \mathcal{I}_{Ros}^2 is negative: indeed Rosenfeld did not use the symbol \mathcal{I}_{Ros}^2 , but he explicitly wrote its square root, cf. equation (57) 864 865 below. Hence, we introduced it in order to compare Rosenfeld's and de Broglie's 866 work. As we shall see in a moment, Rosenfeld did not discuss the square root of 867 the expression \mathcal{I}_{Ros} , but he underlined that it has a geometrical meaning as follows. 868 Parametrizing the five-dimensional path with $\hat{\tau}$ and the particle's four-dimensional 869 world line with the proper time τ , Rosenfeld wrote: 'It is easy to calculate the five-870 dimensional trajectory's slope on the space-time. Indeed, if \bar{S} is a complete integral 871 of equation (55), along the trajectory, from (55) it follows that 872

$$\gamma^{\bar{\mu}\bar{\nu}}\partial_{\bar{\nu}}\bar{S} = \sqrt{m^2c^2 - \frac{e^2c^2}{16\pi G}} \cdot \frac{dx^{\bar{\mu}}}{d\hat{\tau}}, \qquad (57)$$

and from (52), (54) and (53) it follows that

$$\gamma^{\mu\bar{\nu}}\partial_{\bar{\nu}}\bar{S} = mc\frac{dx^{\mu}}{d\tau}.$$
(58)

874 This means that the slope reads:

$$\frac{d\hat{\tau}}{d\tau} = \sqrt{1 - \frac{1}{2\kappa\mu^2}} \tag{59}$$

and therefore it is determined only by the ratio μ ; this geometric interpretation of 875 the ratio μ was on the ground of de Broglie's reasoning.³⁸ [emphasis added] ([Léon 876 Rosenfeld, 1927a]; p. 308). The ratio μ is defined by $\mu = -\frac{mc^2}{c}$ and it encodes the 877 characteristics of the particle, because it involves the particle's mass and charge. 878 The emphasis added at the end of the citation underscores de Broglie's influence on 879 Rosenfeld's approach. Firstly, Rosenfeld's equation (59) is equivalent to de Broglie's 880 equation (22). Secondly, in the previous section we said that from de Broglie's point 881 of view $P_{\bar{\nu}} = \partial_{\bar{\nu}} \bar{S}$ should be interpreted as the five-dimensional generalization of 882

 $^{^{38}}$ See Landau & Lifshitz [1951] for an explanation of the four-dimensional case. Inserting equation (57) into (55), it can be verified that (57) is a complete integral of (55).

 $p_{\mu} = mcg_{\mu\nu}\frac{dx^{\nu}}{d\tau}$. Rosenfeld referred to the fact that equations (57) and (58) made explicit this connection,³⁹ because they implied that $\gamma^{\mu\bar{\nu}}P_{\bar{\nu}} = g^{\mu\nu}p_{\nu}$. Furthermore, Rosenfeld agreed explicitly with de Broglie's idea that the particle's five-dimensional geodesics would be inclined with respect to the hyperplane that locally describes the four-dimensional hypersurface $x^5 = const$. See de Broglie's comments after equation (19).

After having introduced the five-dimensional Universe and its unified description of the gravitational and electromagnetic interaction, the author introduced what he called the 'de Broglie-Schrödinger wave function' ([Léon Rosenfeld, 1927a]; p. 311). Following de Broglie and De Donder, equations (28) and (36), Rosenfeld's general Ansatz for the five-dimensional wave function reads:

$$\Psi(x) = \mathcal{A}(x^0, x^1, x^2, x^3) e^{k\bar{S}}, \qquad (60)$$

where \bar{S} is the Jacobi function (49), k is a constant and the amplitude A is in general a 894 complex function of the form $\mathcal{A} = A + iB$. Like De Donder, Rosenfeld made the choice 895 $k = \frac{i}{\hbar}$ and then he considered the case of real constant amplitude, in order to compare 896 his five-dimensional functional with De Donder's J functional. But Rosenfeld assigned 897 the value of k ab initio, therefore, as we pointed out in the discussion after equation 898 (36), both De Donder and Rosenfeld considered wave functions as complex objects. The periodicity condition is still contained in Rosenfeld's Ansatz (60), because the 900 wave function is periodic in the fifth coordinate, see equation (49). In the case of real 901 constant amplitude A, from equation (60) it follows: 902

$$\frac{\partial \bar{S}}{\partial x^{\bar{\mu}}} = \frac{\hbar}{i} \frac{\partial_{\bar{\mu}} \Psi}{\Psi} \,. \tag{61}$$

Inserting (61) into the HJ equation (55), Rosenfeld obtained the five-dimensional generalization of De Donder's functional equation (43), i.e. $\mathcal{L} = 0$, where the new functional is

$$\mathcal{L}\left(\Psi,\overline{\Psi}\right) = -\gamma^{\bar{\mu}\bar{\nu}}\partial_{\bar{\mu}}\overline{\Psi}\partial_{\bar{\nu}}\Psi - \frac{\mathcal{I}_{Ros}^{2}}{\hbar^{2}}\overline{\Psi}\Psi, \qquad (62)$$

the symbol $\overline{\Psi}$ is the complex conjugate of the five-dimensional wave function and 906 we used for this quantity the symbol \mathcal{I}_{Ros} , equation (56), for brevity. This means 907 that from Rosenfeld's point of view the constant amplitude case corresponded to the 908 classical limit. Indeed, the author underlined: 'In the general case, i.e. when \mathcal{A} is an 909 arbitrary function, \mathcal{L} is no longer null along a trajectory.' ([Léon Rosenfeld, 1927a]; 910 p. 312). As a consequence \mathcal{L} is able to play a central role for the quantum dynamics. 911 Following De Donder, the quantum picture would be described by a variational 912 principle involving (62): Rosenfeld applied De Donder's functional derivative (44) on 913 $\mathcal{L}\sqrt{-g}$ and obtained, by varying with respect to $\overline{\Psi}$ and Ψ independently, the following 914 wave equations: 915

$$\gamma^{\bar{\mu}\bar{\nu}}\nabla_{\bar{\mu}}\partial_{\bar{\nu}}\Psi = \frac{\mathcal{I}_{Ros}^2}{\hbar^2}\Psi \qquad \text{and} \qquad \gamma^{\bar{\mu}\bar{\nu}}\nabla_{\bar{\mu}}\partial_{\bar{\nu}}\overline{\Psi} = \frac{\mathcal{I}_{Ros}^2}{\hbar^2}\overline{\Psi}, \tag{63}$$

and that should be, as Rosenfeld wrote, 'a generalization of the de Broglie-Schrödinger's equation' ([Léon Rosenfeld, 1927a]; p. 312), i.e. equation (26). Having introduced a complex wave function ab initio, Rosenfeld wrote explicitly a wave equation both for Ψ and for $\overline{\Psi}$. The author's functional \mathcal{L} is formally equivalent to the

 $^{^{39}}$ In Appendix D.1 we clarify the connection among equations (57), (58) and (59).

Lagrangian density of a complex scalar field, but as for all of the authors of this 920 period, Ψ is treated as a wave function. This approach has been conceived in a period 921 that lies between the birth of QM and the birth of QFT, when scholars were look-922 ing for a "relativistic quantum mechanics". For this reason we could say that, like 923 De Donder, Rosenfeld was looking for GRQM. The wave equation obtained by vary-924 ing Ψ in (62) is formally equivalent to the five-dimensional wave equation suggested 925 by de Broglie (26). Rosenfeld used De Donder's variational derivative, but he was 926 aware of the fact that this procedure is equivalent to the variational principle used 927 in a modern field theory, obtained varying the integral of the Lagrangian density 928 and imposing that the variations of the fields should be zero at the boundary of the 929 domain of integration. Indeed, Rosenfeld claimed that \mathcal{L} should be the generalization 930 of the Lagrangian considered by Gordon [1927], where Gordon himself suggested to 931 consider the wave function and his complex conjugated as independent variables with 932 vanishing variations at the boundary. Unlike Klein's functional, Rosenfeld's \mathcal{L} func-933 tional had the correct sign to be interpreted as a Lagrangian density [Rocci, 2013]. 934 This follows from the fact that Rosenfeld was influenced by De Donder's approach 935 presented above. Unlike De Donder, Rosenfeld considered a general form for the wave 936 functions, admitting that its amplitude A could be a non-constant function of the 937 four-dimensional coordinates. Rosenfeld noted that in the constant-amplitude case he obtained De Donder's results, which are connected with the classical HJ equation 939 (55) as suggested by De Donder himself. 940

How did Rosenfeld reconcile GR with QM? Like De Donder, after having used the wave-particle duality via the Hamiltonian dynamics, Rosenfeld supposed that, in the case of non-constant amplitude, \mathcal{L} should be the correct generalization of Schrödinger's Lagrangian [Schrödinger, 1927] in the sense of GRQM. Finally, Rosenfeld introduced a variational principle, based on the following five-dimensional action⁴⁰

$$\mathcal{S}_{tot}\left(\gamma,\Psi,\overline{\Psi}\right) = \int d^5x \sqrt{-g} \left[-\tilde{R} + 2\kappa \mathcal{L}\right], \qquad (64)$$

where $2\kappa = \frac{16\pi G}{c^4}$. Rosenfeld did not specify the domain of integration, we suppose 947 that the integral should be performed over an arbitrary portion Ω of the five-948 dimensional space-time. By varying the action with respect to the metric like in 949 equation (6), he obtained the five-dimensional Einstein equations coupled with the 950 complex field Ψ , which are formally equivalent to a system with the four-dimensional 951 Maxwell equations coupled to the scalar field and the four-dimensional Einstein equa-952 tions coupled to the electromagnetic and the scalar fields. By varying the action with 953 respect to $\overline{\Psi}$ and Ψ , using De Donder's functional derivative, Rosenfeld obtained the 954 KG equation (63) for Ψ and $\overline{\Psi}$, respectively, as before, because the curvature's scalar 955 depends neither on the wave function nor its complex conjugate. This is the unified 956 framework that should reconcile, from Rosenfeld's point of view, GR with WM. 957

Did the five-dimensional formalism offer any additional insights beyond these that 958 De Donder could have deduced in his four-dimensional context? As Rosenfeld stressed, 959 the main advantage offered by the five-dimensional Universe was the opportunity to 960 write a unified variational principle ([Léon Rosenfeld, 1927a]; p. 304). It is worth 961 noting that the neutron would be discovered five years later [Chadwick, 1932]. This 962 means that all known elementary particles were charged particles and the unified 963 picture offered by the five-dimensional Universe seemed to be a way to describe the 964 known physical phenomena. As we shall see in Section 4.2, Rosenfeld's approach 965

⁴⁰In equation (64) the determinant of the four-dimensional metric g appears, instead of γ . In Rosenfeld's approach, the two determinants are related by the relation $\gamma = \alpha g$ as explained in Appendix D.2. This means that the presence of g does not affect the equations obtained by varying (64).

966 permitted also to incorporate and, in a certain sense, to justify some of De Donder's 967 ideas.

It is not clear whether Rosenfeld considered his approach as a result or as a point of departure. But it is evident that he tried, for the first time, to investigate the geometry created by the wave function Ψ . In fact, the equations obtained by varying action (64) with respect to the metric are:

$$\tilde{R}_{\bar{\mu}\nu} - \frac{1}{2}\gamma_{\bar{\mu}\nu}\tilde{R} = \kappa T_{\bar{\mu}\nu} , \qquad (65)$$

where Einstein's and Maxwell's equations are coupled to the complex scalar field via the stress-energy tensor $T_{\bar{\mu}\nu}$, defined by Rosenfeld as

$$T_{\bar{\mu}\bar{\nu}} = -\frac{2}{\sqrt{-g}} \frac{\delta\left(\sqrt{-g}\mathcal{L}\right)}{\delta\gamma^{\bar{\mu}\bar{\nu}}} , \qquad (66)$$

⁹⁷⁴ which has the usual form:

$$T_{\bar{\mu}\nu} = \partial_{\bar{\mu}} \overline{\Psi} \partial_{\nu} \Psi + \partial_{\nu} \overline{\Psi} \partial_{\bar{\mu}} \Psi + \gamma_{\bar{\mu}\nu} \mathcal{L} .$$
 (67)

Rosenfeld made no comments on the fact that in general the r.h.s. of equation (65) is 975 a complex quantity. It is worth noting that the author investigated a particular case, 976 i.e. when the wave function's amplitude is real. Hence, the energy momentum tensor 977 is a real quantity. Introducing the wave function on the right side of equation (65). 978 Rosenfeld considered implicitly the wave function as representing the material part 979 creating gravity. In this first paper, a long and technical paper, Rosenfeld did not 980 justify this choice, which seems to be in contrast with the probabilistic interpretation 981 of the wave function, from a modern point of view. As we shall see in the next section, 982 the author would clarify his choice in the following work, where he referred explicitly 983 to Bohr's correspondence principle. 984

Like Klein, Rosenfeld did not consider the 55 component of the equations of motion: the Belgian physicist explicitly stated that this equation can be neglected, because the constancy of γ_{55} implies $\delta\gamma_{55} = 0$ ([Léon Rosenfeld, 1927a]; p. 314)⁴¹.

Before going on, we compare briefly Rosenfeld's approach with that of his men-988 tors. Though Rosenfeld started out generalizing De Donder's approach, the unitary 989 variational principle is presented starting with the action functional (64) instead of De 990 Donder's invariants, i.e. density Lagrangians. It is worth noting that in the same year 991 Klein published independently a similar action, using a real scalar field. Klein cou-992 pled matter and geometry exactly like Rosenfeld did ([Klein, 1927b]; p. 207). Unlike 993 Rosenfeld, in Klein [1927b], Klein will express explicitly some perplexities about this 994 kind of approach, observing that a unified action principle, e.g. that based on (64), 995 was only a starting step towards a unified theory that reconciles WM with GR ([Klein, 996 1927b]; p. 190, footnote (*) at the end of the introduction). In contrast, Rosenfeld, 997 and De Donder with him, seemed to be convinced that the five-dimensional unified 998 action principle would have some interesting features. Thanks to this conviction, the 999 Belgian physicist investigated the quantum character of the metric produced by a 1000 quantum object, represented by the wave function Ψ . 1001

In order to face this problem, Rosenfeld considered the weak-field approximation for the gravitational field, introduced by Einstein in 1916 to study the problem of gravitational waves, because it permitted to integrate the Einstein equations. In this approximation the metric can be written in the following form ([Léon Rosenfeld,

⁴¹As we said in the previous section, this is not correct.

1006 1927a]; p. 319):

$$\gamma_{\bar{\mu}\nu} = \eta_{\bar{\mu}\nu} + h_{\bar{\mu}\nu} \,, \tag{68}$$

where $\eta_{\bar{\mu}\nu}$ is the five-dimensional Minkowski metric and $h_{\bar{\mu}\nu}$ represents the perturbation of the flat metric, which satisfies the condition $|h_{\bar{\mu}\nu}| \ll 1$. Rosenfeld contracted (65) with $\gamma^{\nu\bar{\mu}}$ to obtain an expression for the five-dimensional curvature scalar \tilde{R} , namely⁴²

$$\tilde{R} = -\kappa \left[\gamma^{\nu \bar{\mu}} T_{\bar{\mu}\nu} + \frac{F_{\sigma\lambda} F^{\sigma\lambda}}{2} - \gamma^{\mu\rho} A_{\rho} \nabla_{\lambda} \left(\gamma_{\mu\sigma} F^{\sigma\lambda} \right) \right] \,. \tag{69}$$

After having inserted (69) into equation (65), Rosenfeld used the Ansatz (68) for the metric and he considered linear terms only obtaining:

$$\Box h_{\bar{\mu}\nu} = -\kappa \left[T_{\bar{\mu}\nu} - \frac{1}{2} \eta_{\bar{\mu}\nu} \eta^{\lambda\bar{\sigma}} T_{\bar{\sigma}\lambda} \right]$$
$$= -\kappa \bar{T}_{\bar{\mu}\nu}, \tag{70}$$

where the \Box operator acts only on the usual four dimensions, because the metric 1013 does not depend on the fifth coordinate. In this approximation we are considering 1014 the gravitational field strength far away from the source, i.e. the particle's wave 1015 function, and the second and third term in the r.h.s. of equation (69) can be ignored 1016 in the case of a stationary charge.⁴³ The stress-energy tensor appearing in (70) has 1017 the same form of equation (67), but the curved metric $\gamma_{\mu\nu}$ has been substituted by 1018 the flat metric ([Léon Rosenfeld, 1927a]; p. 319). In particular, in this approximation 1019 the indices are raised and lowered by $\eta_{\mu\nu}$. Rosenfeld was now able to integrate (70), 1020 and obtained, using Rosenfeld's original notation⁴⁴ ([Léon Rosenfeld, 1927a]; p. 319, 1021 Eq. (71)): 1022

$$h_{\bar{\mu}\nu} = -\frac{\kappa}{2\pi} \int \left\{ \bar{T}_{\bar{\mu}\nu} \right\}_{t-\frac{r}{c}} \frac{dxdydz}{r} , \qquad (71)$$

where, according to Rosenfeld, r represents the radial distance and the symbol $\{u\}_{t-\frac{r}{c}}$ means that the function u has been calculated using the variable $t-\frac{r}{c}$: for this reason the (71) components are often called retarded potentials. In order to consider the case of a stationary mass, the author chooses the following form⁴⁵ for the Jacobi function \bar{S} ([Léon Rosenfeld, 1927a]; p. 320):

$$\bar{S} = -\frac{e}{c\beta}x^5 + mcx^0, \qquad (72)$$

that appears in (60), where now the amplitude A is a real function of the fourdimensional coordinates. Using this Ansatz, Rosenfeld was able to calculate explicitly

 $^{^{42}}$ See Appendix D.4 for a detailed explanation.

⁴³Rosenfeld did not write explicitly equation (70), he referred to a 'well known procedure' ([Léon Rosenfeld, 1927a]; p. 319) and wrote directly equation (71).

⁴⁴Rosenfeld did not specify that the integration is carried over a three-dimensional hypersurface Σ at the retarded time. In Appendix D.5 we express equation (71) in a modern notation. In the rest of our paper we will continue to use Rosenfeld's original notation.

⁴⁵Remember that in our notation the combination $\frac{e}{c\beta}x^5$ has the dimensions of an action.

the retarded potentials. Introducing the following functions⁴⁶ of \mathbf{x} and t:

$$\mathcal{F} = \frac{2mc^2}{\hbar^2} \int \left\{ A^2 \right\}_{t-\frac{r}{c}} \frac{dxdydz}{r} , \qquad (73)$$

$$\mathcal{W}_{\mu\nu} = \int \left\{ \partial_{\mu} A \partial_{\nu} A \right\}_{t-\frac{r}{c}} \frac{dxdydz}{r} , \qquad (74)$$

$$\mathcal{G} = \int \left\{ \partial_{\mu} A \partial^{\mu} A \right\}_{t - \frac{r}{c}} \frac{dx dy dz}{r} , \qquad (75)$$

 $_{1031}$ the perturbations of the flat metric are therefore⁴⁷:

$$h_{5i} = 0$$
, $i = 1, 2, 3$, (76)

$$h_{50} = -\alpha\beta \left(\frac{e}{4\pi}\frac{\mathcal{F}}{c^2}\right) \,, \tag{77}$$

$$h_{\mu\nu} = \frac{8G}{c^4} \mathcal{W}_{\mu\nu} \qquad \qquad \mu \neq \nu , \qquad (78)$$

$$h_{\mu\mu} = \frac{2mG}{c^4} \mathcal{F} + \frac{8G}{c^4} \mathcal{G} .$$
⁽⁷⁹⁾

It is worth noting that in (77) and in (79) the Planck constant appears via tha definition of \mathcal{F} (73). In this sense, Rosenfeld's result represents a quantum correction of the flat metric. This is not surprising, because these corrections are generated by the wave function Ψ . In this sense, the result is the first attempt to describe a quantum metric using WM and GR. As far as we know, this is the first time that a quantum metric appears in the history of QG.

Rosenfeld did not emphasize this feature of the metric he found. As we have 1038 said, in his first paper Rosenfeld did not make explicit comments on the physical 1039 meaning of the calculations performed. As we shall see, in his following papers he 1040 would advocate Bohr's correspondence principle in explaining his use of the wave 1041 function as the source of gravitational field. From this perspective, it is easier to 1042 understand why Rosenfeld was more interested in analysing the metric in the case of 1043 a constant amplitude. Indeed, he considered a sort of semi-classical limit, confronting 1044 his "quantum metric" with its classical analogue. In this limit, equations (76), (77), 1045 (78) and (79) should match the metric produced by a classical source of mass m and 1046 charge e, sitting at the origin **O** of the coordinates, at least in the weak-field limit, 1047 known today as the RN solution. The classical metric is presented in Appendix D.6, 1048 equation (D.21). At asymptotically large distances from the source it can be written 1049 as $\gamma_{\bar{u}\bar{\nu}}^{RN} = \eta_{\bar{\mu}\bar{\nu}} + h_{\bar{u}\bar{\nu}}^{RN}$, where the components of the perturbations of the flat metric 1050 are: 1051

$$h_{5i} = 0$$
, $i = 1, 2, 3$, (80)

$$h_{50} = \alpha \beta A_0$$
 where $A_0 = \eta_{00} A^0 = V = -\frac{c}{4\pi m}$, (81)

$$h_{\mu\nu} = 0 \qquad \mu \neq \nu , \qquad (82)$$

$$h_{\mu\mu} = \frac{2mG}{c^2 r_0} , \qquad (83)$$

 $^{^{46}}$ The integration domain is the same as in equation (71).

⁴⁷In equation (79) we used explicitly that α and β satisfy the constrain $\alpha\beta^2 = 2\kappa$, like in Klein's approach.

where, according to Rosenfeld, r_0 represents 'the distance between the origin **O** and 1052 an arbitrary point [of the five-dimensional space-time]' ([Léon Rosenfeld, 1927a]; p. 1053 321). Equations (82) and (83) represent the components of the RN metric in the 1054 weak field approximation expressed using isotropic Cartesian coordinates,⁴⁸ while 1055 (80) and (81) coincide with $\gamma_{5\mu}$ components (5) in the case of a stationary charge. As 1056 we shall see in a moment, in considering the matching between classical metric and 1057 "quantum metric" in the semiclassical limit, Rosenfeld did not consider a point-like 1058 charge, hence $r_0 = r_0(\vec{x})$ should be a sort of "mean distance" from the charged body, 1059 sitting at the origin of the coordinates. 1060

In order to match (76)–(79) with (80)–(83), $\mathcal{W}_{\mu\nu}$ and \mathcal{G} must be zero and, as a consequence, the two following conditions must hold:

$$\partial_{\mu}A = 0 \quad , \tag{84}$$

$$\mathcal{F} = \frac{c^2}{r_0} \,. \tag{85}$$

Equation (84) follows directly from the condition $\mathcal{W}_{\mu\nu} = 0$, while equation (85) can 1063 be obtained comparing (81) with (77). Rosenfeld discussed both these relations: 'The 1064 first condition tells us that a fixed charge can be represented by a wave with stationary 1065 phase and *constant* amplitude.' ([Léon Rosenfeld, 1927a]; p. 322). As stated above, 1066 though Rosenfeld did not emphasize this fact, the constancy of the amplitude, i.e. 1067 condition (85), emerged as a condition to ensure that the quantum description could 1068 contain, at least as a limiting case, the classical description, which in this context 1069 corresponds to the classical five-dimensional RN metric (80)-(83). Besides this, the 1070 wave function of a fixed charge should have a fixed energy $\mathcal{E} = mc^2$, and because 1071 of Heisenberg's uncertainty principle it should spread over the whole space. In a 1072 semi-classical approximation the wave packet is highly localized. Rosenfeld used a 1073 "localized wave function" instead, in the sense that Rosenfeld's wave function is non-1074 zero only inside an arbitrary volume V. Indeed Rosenfeld continued: 'The second 1075 condition is satisfied [...] if we imagine that the amplitude is non-zero inside a finite 1076 volume centred around **O**.' ([Léon Rosenfeld, 1927a]; p. 322). Finally, using the mean 1077 value theorem, the author defined formally the "mean distance" $^{49} r_0$ ([Léon Rosenfeld, 1078 **1927**a]; p. 322): 1079

$$\frac{V}{r_0} = \int \frac{dxdydz}{r} \ . \tag{86}$$

As usual, Rosenfeld did not specify the domain of integration. We suppose that it is the region where the wave function is non-zero, i.e. the volume V. By using definition (86) and the definition of \mathcal{F} , equation (73), in the constant amplitude approximation the condition (85) reads:

$$\mathcal{F} = \frac{2mc^2}{\hbar^2} \int \left\{ A^2 \right\}_{t-\frac{r}{c}} \frac{dxdydz}{r} = \frac{c^2}{r_0} ,$$
$$\frac{2m}{\hbar^2} A^2 \int \frac{dxdydz}{r} = \frac{1}{r_0} ,$$
$$\frac{2mA^2}{\hbar^2} \frac{V}{r_0} = \frac{1}{r_0} ,$$

 $^{^{48}}$ See Appendix D.6 for a detailed discussion.

⁴⁹See Appendix D.6 for a definition of the mean distance using modern notation.

1084 i.e.

$$\frac{2mA^2V}{\hbar^2} = 1.$$
(87)

This condition is consistent from the point of view of dimensional analysis. To under-1085 stand it, let us consider action (64). The presence of the four-dimensional Einstein 1086 coupling κ produces a consequence for the length dimensions of the wave function Ψ . 1087 We remember that the curvature scalar has dimensions $|\tilde{R}| = (length)^{-2}$ for every 1088 space-time dimension and we observe that from (64) it follows that $\kappa \mathcal{L}$ and \tilde{R} have 1089 the same dimensions. As a consequence, the squared wave function amplitude A^2 has 1090 the following dimensions $[A^2] = \frac{(length)(mass)}{(time)^2}$ as it should, because of equation (87). 1091 It is worth noting that from Rosenfel's point of view, the wave function of a particle is 1092 not a point singularity: its amplitude is non zero in a finite volume V. This fact is in 1093 contrast with de Broglie's point of view as Rosenfeld anticipated in the introduction 1094 of his paper. 1095

In this paper, Rosenfeld did not make any particular comment on (87) and on 1096 the whole calculation: he would discuss the physical meaning of the whole apparatus 1097 in the next papers, that we will briefly analyse in the following section. However, 1098 for us, Rosenfeld's calculation acquired a fundamental importance. Indeed, with this 1099 derivation the author showed for the first time how in the semi-classical limit GRQM 1100 is able to reproduce the RN metric in the weak-field approximation. In particular the 1101 condition (87) found by Rosenfeld can be interpreted as the normalization condition 1102 for the wave function. In this pre-second-quantized picture, the normalization condi-1103 tion of the wave function can be imposed using the definition of the Hamiltonian⁵⁰ 1104 ([Landau et al., 1971]) *H*: 1105

$$H = \int d^3x T_{00} \,, \tag{88}$$

where T_{00} is the 00 component of the total stress-energy tensor (67). The integration is 1106 carried out over the three-spatial volume for the following reason. The stress-energy 1107 tensor defined by Rosenfeld is a four-dimensional object, because of the unusual 1108 coupling between matter and geometry in the action (64). The presence of the four-1109 dimensional constant κ means that the stress-energy tensor's components represent 1110 an energy density with respect to the three-dimensional volume, instead of a four-1111 dimensional volume. Rosenfeld was aware of this peculiarity, even if he did make 1112 no specific comment, because he noted that equation (65) imply a relation for the 1113 four-dimensional curvature scalar,⁵¹ namely 1114

$$R = -\kappa \left[\gamma^{\nu \bar{\mu}} T_{\bar{\mu}\nu} - \gamma^{\mu\rho} A_{\rho} \nabla_{\lambda} \left(\gamma_{\mu\sigma} F^{\sigma\lambda} \right) \right] , \qquad (89)$$

that permitted him to define a four-dimensional mass density⁵² ([Léon Rosenfeld, 1927a]; p. 318, Eq. (63)), i.e. the quantity between the squared brackets on the r.h.s. of (89). For a stationary charge, in the weak field limit, the four-dimensional density mass defined by Rosenfeld in (89) coincides with T_{00} . Moreover, for a localized wave packet the Hamiltonian must correspond to the rest energy $\mathcal{E} = mc^2$ of the classical particle. In the case of a constant amplitude, the T_{00} value can be easily read off using

 $^{^{50}\}mathrm{In}$ the weak-field limit, at the first order, the metric is flat.

 $^{^{51}}$ See Appendix D.4 for a detailed explanation.

⁵²Remember that in GR the trace of the stress-energy tensor is proportional to the curvature scalar and it is the energy density at first order in v/c.

equations (62), (60), (67), and the normalization condition for the wave function reads:

$$\int d^3x \frac{2m^2 c^2 A^2}{\hbar^2} = mc^2 \quad \Rightarrow \quad \frac{2m^2 c^2 A^2}{\hbar^2} V = mc^2 \quad \Rightarrow \quad \frac{2mA^2 V}{\hbar^2} = 1, \qquad (90)$$

where V is the three-volume of the localized wave packet. The normalization condition 1123 is precisely Rosenfeld condition (87). This normalization condition can be obtained 1124 also by considering the conserved current $j^{\bar{\mu}}$. In the weak field approximation the 1125 continuity equation is $\partial_{\bar{\mu}} j^{\bar{\mu}} = 0$. Using the wave function Ansatz (60) with a real 1126 constant amplitude A, namely $\Psi = A \exp\left[\frac{i}{\hbar} \left(-\frac{e}{c\beta}x^5 + mcx^0\right)\right]$, the continuity equa-1127 tion reads $\frac{\hbar}{i} \frac{\partial \rho}{\partial t} = 0$, where the squared modulus of the "probability amplitude" ρ is 1128 $\rho = \frac{2m}{\hbar^2} A^2$. By integrating over a three-spatial volume, because of the unusual length 1129 dimensions of the scalar field Ψ , the normalization condition reads $\frac{2mA^2V}{\hbar^2} = 1$, that 1130 is the same result obtained using the stress-energy tensor. 1131

In the rest of his first paper, Rosenfeld tried to generalize his previous results 1132 to the case of a many-body system. This generalization process would continue in 1133 his following papers, where the author also analysed the role of the wave function 1134 amplitude A. Rosenfeld inspected the consequences produced by considering a non-1135 constant amplitude. In particular, he would be interested in its interpretation as a 1136 'potential of the internal forces' ([Léon Rosenfeld, 1927a]; p. 325) that should emerge 1137 when considering a continuous system. This idea was also shared by de Broglie, but 1138 was introduced by De Donder,⁵³ as Rosenfeld wrote: 'Recently, Mr. De Donder has 1139 introduced in WM two important concepts: the notion of *permanence* of a system and 1140 the interpretation of the amplitude A of the Schrödinger's function Ψ as a *potential* 1141 of the internal tensions of the system.⁵⁴ ([Léon Rosenfeld, 1927b]; p. 447). 1142

1143 4.2 The role of the correspondence principle in QG

As noted in our previous section, the first communication was sent to De Donder, 1144 who asked Rosenfeld to work with him during the summer of 1927. Even if they did 1145 not publish a joint paper, they cited each other in the communications published by 1146 the Bulletin de l'Académie royale de Belgique [De Donder, 1927b; Léon Rosenfeld, 1147 1927b.c]. Rosenfeld acknowledged De Donder explicitly at the end of the introduction: 1148 'My warmest thanks to Mr. De Donder, who did not quit to take an active interest 1149 in my work.' ([Léon Rosenfeld, 1927b]; p. 448). At the end of the third paper's 1150 introduction, Rosenfeld underscored again: 'Mr. De Donder played an essential role 1151 in this work, because he suggested to me the basic idea. I owe a lot to De Broglie, 1152 who kindly continued to have a correspondence with me of which I took greatest 1153 advantage.' ([Léon Rosenfeld, 1927c]; p. 574). The main result of Rosenfeld-De Donder 1154 collaboration was the introduction of Bohr's correspondence principle as a physical 1155 interpretation of Rosenfeld's previous mathematical treatment. As far as we know, 1156 this is the first time that Bohr's principle was invoked in searching for a theory 1157 that could reconcile WM with GR. In particular, Rosenfeld and De Donder posed 1158 this principle as one of the founding principles of this new theory, which De Donder 1159 called 'the gravitational wave mechanics' ([De Donder, 1927b]; p. 506). The purpose 1160 of this paragraph is to discuss the role of the correspondence principle, presenting 1161 Rosenfeld's following works: [Léon Rosenfeld, 1927b,c,e]. 1162

⁵³The original citations are not quoted.

⁵⁴We will not deepen the concept of "permanence".

In order to understand the role of the correspondence principle, we start pointing out that Rosenfeld was impressed by the fact that the stress-energy tensor (67) resembled the stress-energy tensor for a particles' system whose form was:

$$T_{\mu\nu} = \sigma_{(m)} g_{\mu\rho} g_{\nu\sigma} u^{\rho} u^{\sigma} + P_{\mu\nu} , \qquad \text{where} \qquad u^{\rho} = \frac{dx^{\rho}}{d\tau} , \qquad (91)$$

as it appears in De Donder's MIT lectures ([De Donder, 1927a] p. 52), and where $\sigma_{(m)}$ represents the mass density as measured by the observer u^{μ} . For a swarm of noninteracting particles $P_{\mu\nu} = 0$, for a perfect fluid with pressure p, $P_{\mu\nu} = p (u_{\mu}u_{\nu} + g_{\mu\nu})$ ([Misner et al., 1973]); p. 132), while if we consider the dissipative processes its form is more complicated. The resemblance between the stress-energy tensor of a scalar field and that of a particle's system emerges as follows. Rosenfeld considered the following Ansatz for the wave function and for the Jacobi function:

$$\Psi(x) = A(x^0, x^1, x^2, x^3) e^{\frac{i}{\hbar}\bar{S}}$$
(92)

$$\bar{S}(x) = -\frac{e}{\beta c} x^5 + S(x^0, x^1, x^2, x^3) , \qquad (93)$$

where now S has an unspecified form and A is an arbitrary real function. The author inserted (92) into equation (67), and the stress-energy tensor components read:

$$T_{\bar{\mu}\nu} = 2\frac{A^2}{\hbar^2}\partial_{\bar{\mu}}\bar{S}\partial_{\nu}\bar{S} + 2\partial_{\bar{\mu}}A\partial_{\nu}A + \gamma_{\bar{\mu}\nu}\mathcal{L}\,,\tag{94}$$

where the 55 component has been explicitly omitted, because Rosenfeld was not interested in the 55 component of five-dimensional Einstein equations. Using the inverse components of the metric, equation (53), Rosenfeld rewrote equation (58), that we rewrite here for convenience

$$\gamma^{\mu\bar{\nu}}\partial_{\bar{\nu}}\bar{S} = mc\frac{dx^{\mu}}{d\tau} , \qquad (95)$$

¹¹⁷⁹ in the following form:

$$g^{\mu\nu}\partial_{\nu}S = mcu^{\mu} + \frac{e}{c}A^{\mu} .$$
(96)

1180 Equations (96) and (93) imply that:

$$\partial_{\mu}\bar{S} = \partial_{\mu}S = g_{\mu\nu}mcu^{\nu} + \frac{e}{c}A_{\mu} , \qquad (97)$$

$$\partial_5 \bar{S} = -\frac{e}{\beta c} \,. \tag{98}$$

Inserting equations (97) and (98) in (94), the author obtained 55 ([Léon Rosenfeld, 1927b]; p. 454):

$$T_{\mu\nu} = \varrho_{(m)} g_{\mu\rho} g_{\nu\sigma} u^{\rho} u^{\sigma} + \Pi_{\mu\nu} \tag{99}$$

$$\beta T_5^{\ \nu} = \varrho_{(e)} u^{\nu} + \Lambda^{\nu} \,, \tag{100}$$

⁵⁵Equation (100) was obtained raising an index with the five-dimensional metric, $\gamma^{\bar{\rho}\bar{\mu}}T_{\bar{\mu}\bar{\nu}}$, and then choosing $\bar{\rho} = \rho$ and $\bar{\nu} = 5$.

where we define, following Rosenfeld, a "quantum" mass density $\rho_{(m)}$ and a "quantum" charge density⁵⁶ $\rho_{(e)}$:

$$\varrho_{(m)} = \frac{2m^2c^2}{\hbar^2}A^2 \qquad \qquad \varrho_{(e)} = -\frac{2em}{\hbar^2}A^2.$$
(101)

Equations (99) and (100) require some comments, because, from Rosenfeld's and De Donder's point of view they are the basis for invoking the correspondence principle. Firstly, the analogy between (91) and (99) is now evident, and this explains why $\rho_{(m)}$ could play the role of a mass density. In order to understand why $\rho_{(e)}$ represents a charge density, we remember that the Maxwell equations on curved space-time for a classical charged system are

$$\nabla_{\mu}F^{\nu\mu} = j^{\mu} \qquad \text{whith} \qquad j^{\mu} = \sigma_{(e)}u^{\mu} , \qquad (102)$$

where $\sigma_{(e)}$ represents the charge density of the system as measured by the observer u^{μ} . On the other hand, the Maxwell equations obtained by the five-dimensional Einstein equations coupled to the wave function stress-energy tensor (65) are⁵⁷:

$$\nabla_{\mu}F^{\nu\mu} = \beta T_5^{\nu} \,. \tag{103}$$

Therefore, it is evident that βT_5^{ν} could play the role of the density current j^{μ} and, as a consequence, equation (100) defines a charge density $\varrho_{(e)}$.

Secondly, this is the point where Bohr's principle comes into play. At the end of the 1196 introduction of his communication, Rosenfeld underscored that the identification of 1197 $\rho_{(m)}$ and $\rho_{(e)}$ with the mass and electric densities of quantum system is 'a particularly instructive aspect of the *correspondence principle*' ([Léon Rosenfeld, 1927b]; p. 448): 1198 1199 he stressed that this claim would deserve further analysis and that the connection 1200 between the above identification and the correspondence principle has been suggested 1201 by De Donder. At the end of the fifth section of the brief communication, Rosenfeld 1202 remarked that (we changed the original equation's numbers in order to fit our numer-1203 ical order): 'equations (99) and (100) show that $\varrho_{(m)}$ and $\varrho_{(e)}$ should be interpreted 1204 as a mass density and an electric density of the system, or, better(*), corresponding 1205 to the system [...]' ([Léon Rosenfeld, 1927b]; p. 454). Rosenfeld himself used the ital-1206 ics and in the footnote corresponding to the symbol (*) he underscored again that 1207 this remark had been suggested by De Donder. The term "corresponding" referred to 1208 the formal correspondence between a classical and a quantum system. Indeed, $\varrho_{(m)}$ 1209 and $\varrho_{(e)}$ depend on the wave function's amplitude. In the following papers, Rosenfeld 1210 would clarify how his approach is connected with Bohr's correspondence principle. 1211 Our last comment concerns the terms $\Pi_{\mu\nu}$ and Λ_{ν} . Their precise form will not be 1212 discussed here, but it is worth noting that they contain the contribution due to the 1213 fact that the amplitude is not constant. From Rosenfeld's and De Donder's point of 1214 view the $\Pi_{\mu\nu}$ tensor would represent the contribution of the internal forces of the 1215 system, while Λ_{ν} was called 'quantum current' ([Léon Rosenfeld, 1927e]; p. 665) by 1216 Rosenfeld, maybe because it has no classical analogue. 1217

In the third communication Rosenfeld dedicated an entire section to enunciate his principle of correspondence, explicitly referring to Bohr's principle, also describing what he had in mind as QG theory (we changed the original equation numbers in order to fit our numerical order):

¹²²² 'The wave mechanics obtained using the variational principle (64) realizes ¹²²³ formally the fusion between Gravity and quantum theory. To the *field*

⁵⁶Remember that ab inizio we decided to consider the case of q = -e.

⁵⁷See Appendix D.3 for technical details.

equations that describe gravitational and electromagnetic phenomena, we 1224 added the equation of quantization (26), that rules the quantum-energy 1225 exchanges. In this last equation intervenes the fundamental quantity Ψ , 1226 and the fusion between the two theories is represented by the fact that the 1227 five-dimensional matter tensor that is present in the gravitational field 1228 equation is defined using the fundamental quantity Ψ ; on the contrary, 1229 in a *pure* Einsteinian gravitational theory, this tensor is a function of 1230 different fundamental quantities of the system: the mass density $\sigma_{(m)}$ 1231 and the electric charge density $\sigma_{(e)}$.⁵⁸ ([Léon Rosenfeld, 1927c]; p. 574). 1232

Rosenfeld used different letters referring to the mass and charge densities, because he wanted to emphasize the difference between a classical system and the corresponding quantum system. The author continued:

1235

1257

'The new definition of the stress-energy tensor as a function of Ψ , (67), 1236 implies a modification of our conception for the role of the fundamental 1237 quantities $\sigma_{(m)}$ and $\sigma_{(e)}$. In the Einsteinian theory these quantities inter-1238 vene directly in in the field equations in order to fix the gravitational 1239 and the electromagnetic potentials, corresponding to a given distribu-1240 tion $(\sigma_{(m)}, \sigma_{(e)})$. In Wave Mechanics, these quantities do not intervene 1241 directly, but through [..] the quantity Ψ . [...] The material tensor $T^{\bar{\mu}\bar{\nu}}$ as 1242 a function of Ψ should not necessarily be identical to the material tensor 1243 of pure Gravity, which is defined as a function of $\sigma_{(m)}$ and $\sigma_{(e)}$. It seems 1244 desirable to analyse, thenceforward, as soon as possible, the behaviour of 1245 the $T^{\bar{\mu}\bar{\nu}}$ tensor, in order to emphasize all possible modifications to 1246 Gravity produced by the introduction of the quantum quantity 1247 Ψ ; this is the role of the *principle of correspondence*. [bold form added]⁵⁹ 1248 ([Léon Rosenfeld, 1927c]; p. 575). 1249

The bold text emphasizes clearly what was the physical meaning of the calculation presented in Section 4.1. From Rosenfeld's point of view, the introduction of the wave function was responsible for the modifications of the "pure", i.e. classical, GR, because even in the case of constant amplitude, it permits us to introduce two quantum quantities, corresponding to classical quantities $\sigma_{(m)}$ and $\sigma_{(e)}$: through the new stressenergy tensor, the new quantities $\varrho_{(m)}$ and $\varrho_{(e)}$, defined by (101), must be considered as the quantum source of gravitational and electromagnetic field. Indeed Rosenfeld continued:

'The comparison between $\varrho_{(m)}$ and $\varrho_{(e)}$, and $\sigma_{(m)}$ and $\sigma_{(e)}$ will show us 1258 how the quantum objects will modify the gravitational and the electro-1259 magnetic phenomena. It will be possible to enunciate a more precise and 1260 general correspondence principle; [...] there are some precise formulas that 1261 define, in a strict sense, the principle of correspondence and that estab-1262 lish the identification of the formal schema of wave mechanics with the 1263 gravitational schema of Th. De Donder, [...] showing how Wave Mechanics 1264 widens the picture of the pure Gravity, in order to incorporate quantum 1265 phenomena.' ([Léon Rosenfeld, 1927c]; p. 575). 1266

⁵⁸The term 'pure Einsteinian gravitational theory' seems to be referred to the classical theory obtained without the introduction of the "quantum field". We introduced Rosenfeld symbols $\sigma_{(m)}$ and $\sigma_{(e)}$ in equations (91) and (102) respectively.

⁵⁹The term *pure Gravity* can be interpreted as GR. See also footnote [58].

It is important to stress that, like Klein, de Broglie and De Donder, Rosenfeld 1267 never discussed the role of the boundary conditions of the wave function. Like De 1268 Donder he referred to the introduction of the wave function as the 'equation of 1269 quantization'. It is worth to remember that Heisenberg's uncertainty principle was 1270 introduced in February of the same year [Heisenberg, 1927]. This coincided with 1271 the fact that Rosenfeld considered it sufficient to introduce the wave function into 1272 Einstein's equations in order to describe correctly the coupling between gravity and 1273 quantum matter. 1274

Rosenfeld did not cite any of Bohr's papers, but the idea that the correspondence 1275 principle could be a theoretical argument to infer the behaviour of a quantum system 1276 with respect to the classical one is a consequence of Bohr's influence. Indeed, in the 1277 introduction of the third communication, Rosenfeld declares that his approach, i.e. 1278 the variational principle, is a 'formal theory' (Léon Rosenfeld, 1927c]; p. 573). Then 1279 he continued: 'To put a physical interpretation [on the formal theory], we let ourselves 1280 be guided by the *correspondence principle*, using the interpretation given by Klein 1281 [1927c] ...' ([Léon Rosenfeld, 1927c]; p. 573). 1282

In order to understand Bohr's role, we briefly analyse Klein's paper [Klein, 1927c]. 1283 Klein's work is a cornerstone of the history of QM. Before that article, matrix mechan-1284 ics was the only approach incorporating the correspondence principle, 60 as Heisenberg 1285 himself reported in his review of matrix mechanics' successes in 1926 (Mehra & 1286 Rechenberg, 2001f; p. xxxii). In this sense, the title of Klein's contribution was very 1287 revealing: Electrodynamics and Wave Mechanics from the point of view of the Cor-1288 respondence Principle. As reported in Mehra & Rechenberg [2001f], Bohr was aware 1289 of the content of Klein's work and he expressed an enthusiastic comment in a let-1290 ter to Schrödinger ([Mehra & Rechenberg, 2001f]; p. 176). In particular, Bohr was 1291 fascinated by the connection between Hamiltonian mechanics and HJ dynamics of 1292 wave rays, that generated Klein's relativistic WM. Paraphrasing Bohr's words, he 1293 was interested in the fact that thanks to this analogy it is possible, on the basis of 1294 WM, to build a corresponding theory. Klein's main purpose was to investigate the 1295 possibilities of exploiting relativistic WM for understanding atomic processes involv-1296 ing discontinuities. In Klein's paper, the correspondence principle intervenes when 1297 the author tries to modify Maxwell's equations. Schrödinger also expressed the idea 1298 that the wave function 'possesses the property to enter even the untouched [classical] 1299 Maxwell-Lorentz equations between the electromagnetic field vectors as a "source" 1300 of the latter' (Mehra & Rechenberg, 2001f; p. 43). 1301

In 1927 Schrödinger investigated also the effect on the stress-energy tensor 1302 obtained by a unified variational principle involving the Maxwell's Lagrangian and 1303 the complex scalar field Lagrangian, i.e. 'the de Broglie's wave' ([Schrödinger, 1927]; 1304 p. 265). Unlike Klein, de Broglie and Rosenfeld, Schrödinger declared explicitly that 1305 he would consider neither additional dimensions, nor gravitational field contribu-1306 tions. Indeed, Schrödinger's Lagrangian \mathcal{L}_S is the sum of Maxwell's Lagrangian, 1307 $\mathcal{L}_{em} = -\frac{1}{4}F_{\mu}F^{\mu\nu}$, and \mathcal{L}_{ψ} , the Lagrangian for material fields, which is related to 1308 De Donder's work, see (48), because Schrödinger cited De Donder's contribution: 1309

$$\mathcal{L}_{S} = \mathcal{L}_{em} + \mathcal{L}_{\psi} = -\frac{1}{4} F_{\mu} F^{\mu\nu} - \eta^{\mu\nu} \left(\partial_{\mu} \psi + \frac{i}{\hbar} \frac{e}{c} A_{\mu} \psi \right) \left(\partial_{\nu} \overline{\psi} - \frac{i}{\hbar} \frac{e}{c} A_{\nu} \overline{\psi} \right) + \frac{m^{2} c^{2}}{\hbar^{2}} \overline{\psi} \psi .$$
(104)

1310

⁶⁰Heisenberg referred to the fact that classical results can be obtained, in matrix mechanics approach, in the limit of high quantum numbers.

¹³¹¹ \mathcal{L}_S can be obtained after a dimensional reduction from Rosenfeld's Lagrangian ¹³¹² (64) in the limit of a flat background. But Schrödinger did not investigate the role of ¹³¹³ ψ as a source of the electromagnetic field, because he explicitly asserted that the KG ¹³¹⁴ Lagrangian \mathcal{L}_{ψ} did not describe any real field. In spite of this, Klein analysed this ¹³¹⁵ aspect, inspired by the idea to use the correspondence principle. First he manipulated ¹³¹⁶ his scalar relativistic equation to define the four-vector $j^{\mu} = (\rho; j^i)$, where⁶¹ ([Klein, ¹³¹⁷ 1927c]; p. 414, Eqs. (20)):

$$\rho = -\frac{e}{2mc^2} \left\{ -\frac{\hbar}{i} \left(\overline{\psi} \frac{\partial \psi}{\partial t} - \psi \frac{\partial \overline{\psi}}{\partial t} \right) + 2e\overline{\psi}\psi A^0 \right\}$$
(105)

$$j^{i} = -\frac{e}{2m} \left\{ \frac{\hbar}{i} \eta^{ij} \left(\overline{\psi} \partial_{i} \psi - \psi \partial_{i} \overline{\psi} \right) + 2 \frac{e}{c} \overline{\psi} \psi A^{i} \right\} .$$
(106)

Then he showed that using the usual optical geometric Ansatz $\psi = e^{\frac{i}{\hbar}S}$ for the wave function, in the semiclassical limit $\hbar \to 0$, equations (105) and (106) reduce to the components of the usual potentials for a relativistic scalar charged particle, namely:

$$\rho_{cl} = -\frac{e}{\sqrt{1 - (v^2/c^2)}} \tag{107}$$

$$j_{cl}^{i} = -\frac{ev^{i}}{\sqrt{1 - (v^{2}/c^{2})}} , \qquad (108)$$

where v^i is the three-velocity of the particle⁶² and v its modulus. Finally, using the correspondence principle, Klein interpreted equations (105) and (106) as the source for the electromagnetic field, in order to investigate the quantum modifications of the Maxwell equations, namely:

1324

$$\partial_i E^i = 4\pi\rho \tag{109}$$

$$\varepsilon^{ijk}\partial_j B_k - \frac{1}{c}\frac{\partial E^i}{\partial t} = \frac{4\pi}{c}j^i.$$
(110)

Klein solved the Maxwell equations (109) and (110) using the advanced and the 1325 retarded potentials, in order to write an expression for the electric and the magnetic 1326 fields as functions of ψ . Klein identified these electric and magnetic fields with the 1327 electromagnetic field produced by the bounded electron,⁶³ by means of the correspon-1328 dence principle ([Klein, 1927c]; p. 422, Eqs. (41). See also Eqs. (33), (28) and (18)). 1329 As we have seen, Rosenfeld followed the same path in order to obtain an expression 1330 for the metric components, explicitly referring to Klein's paper. In this sense, Rosen-1331 feld was the first author to introduce the correspondence principle in the context of 1332 QG. It is worth noting that in the five-dimensional picture the Maxwell equations are 1333 naturally coupled to the four-current, like Rosenfeld himself showed with relations 1334 (103). This seemed to be another advantage of the five-dimensional approach. 1335

⁶¹The symbols have the usual meaning. We remember that the electromagnetic potentials are $A^{\mu} = (A^0; A^i)$.

 $^{^{62}}$ The role of the analogy between Hamiltonian dynamics and the dynamics of wave's rays is fundamental to obtain these relations.

 $^{^{63}}$ Unlike Rosenfeld, Klein considered also the full quantum treatment, introducing the eigenfunctions expansion for the wave field.

1336 4.3 Back to the present

In his last paper of the year,⁶⁴ written in October 1927, Rosenfeld made a detailed and 1337 wider exposition of all the concepts introduced in his previous work. His idea was to 1338 formulate a sort of formal basis for the five-dimensional Universe as a unified frame-1339 work for GR and WM. The foundations of the whole building are three principles: a 1340 variational principle, i.e. equation (64); the principle of Schrödinger eigenfunctions, 1341 i.e. the usual 'boundary conditions that must be imposed on Ψ and $\overline{\Psi}$ in order to 1342 quantize the system' ([Léon Rosenfeld, 1927e]; p. 665); and the correspondence prin-1343 ciple, that the author formulated with the help of De Donder. Rosenfeld also cited a 1344 paper written by De Donder, where the latter tried to give a more precise formulation 1345 of the principle [De Donder, 1927b]. Unlike Rosenfeld, De Donder will not abandon 1346 this idea in the future. Indeed while Rosenfeld seemed to be convinced that quantum 1347 theory should modify GR, De Donder will continue to claim that GR and WM, were 1348 compatible theories [De Donder, 1930]. 1349

Rosenfeld confirmed the ideas proposed in the previous paper, claiming that the 1350 components of the new stress-energy tensor as a function of the wave function Ψ 1351 should play the role of 'quantum currents', i.e. quantum source for the right side of 1352 Maxwell and Einstein equations. The author wrote explicitly: 'The correspondence 1353 principle consists in stating that this analogy is not only a formal analogy, but also a 1354 physical analogy.' ([Léon Rosenfeld, 1927e]; p. 666). He also emphasized the particular 1355 nature of the correspondence principle: 'There exist *postulates* in the sense of the 1356 formal logic, whilst the correspondence principle is a physical principle [...]' ([Léon 1357 Rosenfeld, 1927e]; p. 667). Rosenfeld meant that the extension of the analogy from 1358 the formal plane to the physical plane is a sort of meta-sentence, and it was different, 1359 in this sense, from a formal sentence of the "basic language" of the equations, like 1360 e.g. the variational principle. 1361

Rosenfeld's approach, as well as de Broglie's proposal were briefly discussed at the 1362 Solvay conference. As stated above, in Section 2, Rosenfeld was not officially admitted 1363 to the conference, but De Donder invited him to follow him, in order to have the 1364 possibility to meet Pauli at the conference. The conference's proceedings showed once 1365 again how de Broglie, Rosenfeld and De Donder agreed on the meaning of the five-1366 dimensional Universe. De Broglie asserted that De Donder succeeded in harmonizing 1367 Einstein theory with WM (Bacciagaluppi & Valentini, 2009); p. 483); De Donder 1368 tried to draw attention to the MIT lectures we previously discussed, speculating on a 1369 connection between his correspondence principle and Bohr reflections (Bacciagaluppi 1370 & Valentini, 2009; p. 483). Subsequently De Donder stated that there is a connection 1371 between de Broglie's contributions, his work and Rosenfeld's ideas (Bacciagaluppi & 1372 Valentini, 2009; p. 499 and 519). De Donder will try again to discuss his approach 1373 ([Bacciagaluppi & Valentini, 2009]; p. 470; 471; 510), but the questions raised by De 1374 Donder and de Broglie will not be faced by the group of physicists. 1375

¹³⁷⁶ De Donder's approach to Hamiltonian dynamics discussed in Section 2 is peculiar, ¹³⁷⁷ because he introduced systematically the use of poly-momenta p_{μ} obtained starting ¹³⁷⁸ with a Lagrangian $\mathcal{L}(y^a, \partial_{\mu}y^a)$, which were functions of some variables y^a and its ¹³⁷⁹ derivatives, deriving it with respect to all of the derivatives, $p^a_{\mu} = \frac{\partial \mathcal{L}}{\partial \partial_{\mu} y^a}$, instead of ¹³⁸⁰ using the time derivative only as usual. This convention, sometimes called the De ¹³⁸¹ Donder-Weyl approach, and its generalization to a curved space-time has survived

⁶⁴In a brief communication to the *Comptes rendus* in June of the same year, Rosenfeld claimed that he was able to reproduce Epstein's description of 'the magnetic electron of Uhlenbeck and Goudsmith' ([Léon Rosenfeld, 1927d]; p. 1541), i.e. the spinning electron, using the five-dimensional apparatus described in the previous section. We will not go into the reasons that could explain Rosenfeld's claim, because we postpone this analysis to a future work.

¹³⁸² until recent years, as an alternative approach for the quantization of gravity, and it ¹³⁸³ is today known as pre-canonical quantization [Kanatchikov, 1998, 2014].

At the end of 1929, after his stay in Göttingen, Rosenfeld moved to Zürich where, 1384 stimulated by Pauli, tried to inspect what we today call the gravitational self-energy 1385 of a quantized electromagnetic field. In Léon Rosenfeld [1930a] he approached the 1386 problem in a way that resembles the work analysed here. Like in his previous work, 1387 he integrated again the linearised Einstein equations, this time coupled with Maxwell 1388 equations only. The quantized electromagnetic field played the role of the complex 1389 scalar field. Rosenfeld used the annihilation and creation operators approach for treat-1390 ing the electromagnetic field, hence the metric field $h_{\mu\nu}$ itself was described by an 1391 operator. In this sense he obtained again a sort of quantum metric, because it is gen-1392 erated by a quantum field. Rosenfeld did not cite the previous papers we analysed, 1393 but we must stress the importance played by his early work, because of the affinity 1394 of the path followed by the author. 1395

The term quantum metric could be understood in a complementary way. The 1396 quantum corrections to the classical gravitational field can be considered as the contri-1397 bution to the classical effects produced by the quantization of the gravitational field. 1398 In the mid thirties, Bronstein [1935] would quantize for the first time the gravitational 1300 field directly in the weak field limit, in order to understand quantum deviations from 1400 the classical Newton law. Only 37 years later, after the development of perturbation 1401 theory, Duff [1973] tried to understand the quantum corrections to the Schwarzschild 1402 metric. Duff used explicitly a classical source and he quantized directly the gravita-1403 tional field. At the tree level, in the weak field limit, he obtained the classical results, 1404 while the quantum corrections came from the one-loop corrections. 1405

Finally we address the following question: what is the physical meaning of Rosen-1406 feld's result from the modern point of view? Rosenfeld interpreted the particle's wave 1407 function as the source of the gravitational field. From a modern point of view, this 1408 approach treats the gravitational interaction as a classical phenomenon and the par-1409 ticle's description as fully quantized. This means that Rosenfeld's procedure gives a 1410 semi-classical result, even in the case of non constant amplitude. From a modern point 1411 of view, Rosenfeld's results can be obtained as non-relativistic limit of the so-called 1412 semi-classical Einstein equations, an approach formally suggested by Møller for the 1413 first time [Møller, 1962]. These equations are obtained by replacing the stress-energy 1414 tensor, i.e. the r.h.s. of Einstein equations, by the expectation value of the stress-1415 energy operator $\hat{T}_{\mu\nu}$ with respect to some quantum state $|\Psi\rangle$. In four dimensions 1416 they have the following form: 1417

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 8\pi G \langle \Psi \hat{T}_{\mu\nu} | \Psi \rangle .$$
(111)

The modern interpretation of equation (111) is connected with the character of the 1418 coupling between gravity and matter. This character has not yet been clarified and it 1419 is an open problem in the QG research area. It is equivalent to the question whether 1420 gravity should be quantized or not von Borzeszkowski et al., 1988. This is a long 1421 debate, see e.g. Carlip [2008] and Kiefer [2004], that divided the physicist community 1422 in two groups, initiated incidentally by Rosenfeld himself [Léon Rosenfeld, 1963]. On 1423 one side those who believe that the gravitational interaction must be quantized, on 1424 the other side those who believe that gravitational interaction must remain classical. 1425 As a consequence, for the first group equations (111) can be derived approximately 1426 from canonical QG as a kind of mean-field equation [Kiefer, 2004]. In this case, the 1427 metric obtained integrating the linearised Einstein equations following Rosenfeld's 1428 procedure is a sort of "mean metric" $\langle \Psi \hat{g}_{\mu\nu} | \Psi \rangle$, where the hat-symbol means that the 1429 metric should be an operator. This perspective is also shared by those who inves-1430 tigate the behaviour of QFT on a curved background [Birrel & Davies, 1982], that 1431

led to Hawking's results on black hole's entropy. For the second group the coupling 1432 1433 between quantum fields and classical gravity described by Einstein equations should be understood as a fundamental description of nature. As a consequence, they inter-1434 pret the l.h.s. of (111) as evaluated using the classical metric. From this perspective, 1435 a possible starting point for reconciling WM with gravity is the so called Schrödinger-1436 Newton equation⁶⁵ [Bahrami et al., 2014], where the source of the gravitational field 1437 is represented by the squared modulus of the wave function. We do not enter the 1438 debate whether which approach could be the fundamental one, because we believe 1439 that any extension of our conceptual framework for the description of nature would 1440 be of interest in itself. We observe that recently there has been a revival of Rosen-1441 feld's ideas coming from the second group of physicists. Modern authors, Giulini 1442 & Grossardt [2012] and Bahrami et al. [2014], with different scope, used some of 1443 the Rosenfeld's ideas, extended to the non-static case. More precisely, in Giulini & 1444 Grossardt [2012], the authors studied the coupling between KG field and gravity in 1445 the case of a non-static spherical symmetric space-time, in the limit of semi-classical 1446 and non-relativistic approximation from the four-dimensional point of view. Follow-1447 ing Kiefer's scheme for non-relativistic and semi-classical approximation, the authors 1448 investigated KG equation on a curved background, showing that it reduces itself, in 1449 this WKB-like scheme, to the Newton-Schrödinger equation, at a certain order of 1450 the WKB expansion. Einstein equations coupled with the KG stress-energy tensor 1451 reduces, in the same approximation, to the Poisson equation for the gravitational 1452 potential, where the wave function amplitude plays the role of the mass density. This 1453 means that, like in Rosenfeld's scheme, the wave function is the source of the metric. 1454 At the order chosen by the authors, the metric itself results as an expansion in terms 1455 of $\frac{\hbar}{c^2}$ powers and it depends on the wave amplitude of the field. In the weak field limit, 1456 the quantum-mechanical description can be derived from the field-theoretic approach 1457 with a well defined procedure, which allows one to use the wave function, instead of 1458 Fock's states [Robertson, 1972]. In Bahrami et al. [2014], the authors refined their 1459 analysis using the second-quantised formalism and hence they apply the procedure 1460 to find the quantum mechanical limit. Once again they find that the wave function 1461 is the source of the gravitational field, like in Rosenfeld's approach. 1462

¹⁴⁶³ 5 Summary and conclusions

In this paper we have described the earliest of Rosenfeld's contributions of 1927. From an historical point of view, Rosenfeld's work is interesting for various reasons. First, it contains many ingredients that the author will use in his future work. Second, it shows how Rosenfeld was influenced by his mentors: Oskar Klein, Louis de Broglie and Theophile De Donder. Third, it offers a connection between the history of QM and the history of QG.

We started considering the main results achieved by his mentors, at the time he 1470 started to write his first paper. Klein wrote a five-dimensional unified variational 1471 principle for the electromagnetic and the gravitational field. He introduced the rel-1472 ativistic wave equation on a curved background using the correspondence between 1473 Hamiltonian dynamics for point particles and the HJ equation in the geometrical 1474 optics limit. Following this correspondence, Klein tried to introduce a sort of massless 1475 KG equation, in analogy with light. De Broglie was pressed by Rosenfeld, who joined 1476 the French physicist in Paris, to investigate the five-dimensional Universe features. 1477 De Broglie showed that it is not necessary to consider null-geodesics, and that the 1478 four-dimensional geodesics can be represented as the projection of five-dimensional 1479

⁶⁵The Schrödinger-Newton equation was introduced by Roger Penrose to provide a dynamical description of the quantum wave function's collapse [Penrose, 1996].

geodesics. De Broglie built his five-dimensional Universe using an inconsistent time-1480 like extra dimension, as Klein himself would note in a following paper. De Donder, 1481 the third character of our story, introduced the Lagrangian approach involving the 1482 wave function, treating it as a field, again using the correspondence between Hamil-1483 tonian particle dynamics and the HJ equation for wave's rays. De Donder interpreted 1484 the introduction of a unified variational principle as the mathematical instrument 1485 responsible for the quantization of the system, because it produces the KG equation. 1486 He was convinced that no modifications of GR were needed for describing quantum 1487 phenomena. De Donder played a fundamental role in Rosenfeld's work. Rosenfeld 1488 sent De Donder his first paper, who presented it for publication at the Bulletin de 1489 l'Académie royale de Belgique journal. Even though we have not analysed any De 1490 Donder-Rosenfeld correspondence, a collaboration between these authors emerges 1491 clearly. Furthermore, De Donder invited Rosenfeld to the fifth Solvay conference, 1492 where De Donder tried to draw attention to Rosenfeld's work and where Rosenfeld 1493 met Einstein and the physicists of the Göttingen school. 1494

After having introduced Klein's, de Broglie's and De Donder's approaches, we 1495 considered Rosenfeld's work. In his first paper, Rosenfeld tried to walk one step 1496 ahead with respect to his mentors. He decided to put De Donder's action model in 1497 a five-dimensional context, building upon the work of Klein and de Broglie. His sec-1498 ond contribution, central in our analysis, was to address the task of understanding 1499 which metric can be generated by a quantum object, i.e. a localized electron's wave 1500 function. Rosenfeld tried also to understand which conditions must hold in order 1501 that WM and GR could reproduce in a semi-classical approximation a classical met-1502 ric in the weak field limit. Studying this problem he presented for the first time a 1503 quantum modification of the flat metric, because of the appearance of \hbar . In his fol-1504 lowing papers, thanks to De Donder's collaboration, Rosenfeld succeeded in giving 1505 a physical meaning to his mathematical treatment. De Donder recognized the idea 1506 of Bohr's correspondence principle in using the wave function's stress-energy tensor 1507 as a source of the gravitational field. In his third communication Rosenfeld himself 1508 explicitly recognized that his approach to QG was inspired to what Klein did in the 1509 context of Maxwell's equations. 1510

Thanks to De Donder, Rosenfeld started to interact with Pauli, Jordan, Bohr 1511 himself and many other physicists who will play, unlike de Broglie and De Donder, 1512 a fundamental role in constructing the new quantum theory of fields. After 1927, 1513 Rosenfeld will convince himself of the importance of quantizing these new objects 1514 and, stimulated by Pauli, he will study again the problem of a quantum metric, but 1515 using the new-born quantum theory of massless spin-1 fields Léon Rosenfeld, 1930a. 1516 From an historical point of view, this paper concluded what we called the prehistory 1517 era in the history of QG. 1518

Even if he never considered his early papers on QG an important work, Rosenfeld's contributions show how the search of a theory that could reconcile quantum phenomena with GR started early and that it also reached interesting results, that will continue to be valid in the context of quantum field theory on a curved spacetime. Even if Klein, de Broglie, De Donder and Rosenfeld were not a research group as in our modern meaning, in 1927 their works were related by a common purpose.

The problem of finding a quantum theory of gravity has never been limited, and is 1525 not limited today, to the quantization of gravitational interaction only. We now know 1526 that attempts to apply directly to the gravitational field quantization procedures, 1527 which have been successful in other contexts, have failed. From the beginning of the 1528 prehistory of QG, the authors that tried to face the problem of reconciling quantum 1529 phenomena with gravity interpreted the idea of QG in the broadest sense. From an 1530 historical point of view, the following statement is particularly true: 'In the broadest 1531 sense, a quantum theory of gravitation would represent an extension of our conceptual 1532

framework for the description of nature: any such extension would be interest in itself.' ([Ashtekar & Geroch, 1974]; p. 1213).

Acknowledgements: We express our gratitude to all anonymous referees who gave us the opportunity of improving the original manuscript. We are very grateful to Kurt Lechner for his invaluable comments and suggestions.

¹⁵³⁸ This work has been supported in part by the DOR 2016 funds of the University of Padua.

Appendix A Wave optics and null-geodesics in Klein's five-dimensional manifold

Klein's original idea was to write a wave equation in analogy with light in the context of his five-dimensional Universe. This appendix follows Klein's original approach
[Klein, 1926a].

In a curved five dimensional space-time, a relativistic generalization of Schrödinger equation is represented by the following equation:

$$a^{\bar{\mu}\bar{\nu}} \left(\delta^{\bar{\sigma}}_{\bar{\nu}} \frac{\partial}{\partial x^{\bar{\mu}}} - \Gamma^{\bar{\sigma}}_{\bar{\mu}\bar{\nu}} \right) \partial_{\bar{\sigma}} \Psi = a^{\bar{\mu}\bar{\nu}} \nabla_{\bar{\mu}} \partial_{\bar{\nu}} \Psi = 0 , \qquad (A.1)$$

where Ψ is the wave function and the covariant derivative $\nabla_{\bar{\mu}}$ is defined using the Christoffel symbols $\Gamma^{\bar{\sigma}}_{\bar{\mu}\bar{\nu}}$. As stated in the main text, Klein defined the Christoffel symbols using the space-time metric $\gamma_{\bar{\mu}\bar{\nu}}$, that we rewrite here for convenience:

$$d\sigma^2 = \alpha d\theta^2 + ds^2 \,, \tag{A.2}$$

where

$$d\theta = dx^5 + \beta A_{\mu} dx^{\mu} \quad ; \quad g_{\mu\nu} = \gamma_{\mu\nu} - \frac{16\pi G}{c^4} A_{\mu} A_{\nu} \quad ; \quad ds^2 = g_{\mu\nu} dx^{\mu} dx^{\nu} \quad . \tag{A.3}$$

Equation (A.1) resembles a massless equation for a scalar field, where the inverse of the space-time metric $\gamma^{\bar{\mu}\bar{\nu}}$ has been replaced by the tensor $a^{\bar{\mu}\bar{\nu}}$, whose covariant components are defined by equation (11). As stressed in Sections 3.1 and 3.2, this fact generated the ambiguity in Klein's approach, criticized by de Broglie. Following Klein's approach, we shall show how wave equation (A.1) is connected with fivedimensional null-geodesics that reduce to the four-dimensional equations of motion for charged massive particles in a combined electromagnetic and gravitational field.

In the geometrical optics limit a wave front propagates locally as a plane-fronted
 wave. Therefore, the Ansatz for the wave function is

$$\Psi(x) = Ae^{i\omega S(x)} \tag{A.4}$$

where ω is so large that only the term proportional to ω^2 in equation (A.1) need to be taken into account. The function S = S(x) is termed the eikonal and it characterizes the phase of the wave. Substituting (A.4) into the wave equation, the term with two derivatives is proportional to ω^2 and equation (A.1) reads:

$$a^{\bar{\mu}\bar{\nu}}\partial_{\bar{\mu}}S\partial_{\bar{\nu}}S = 0 \quad . \tag{A.5}$$

Last equation resembles the eikonal equation for light rays, that describes the propagation of the wave front S(x) of light rays. In the HJ approach, it can be derived ¹⁵⁶⁵ by a particular Hamiltonian, whose Hamilton equations describe the dynamics of ¹⁵⁶⁶ the particle associated to the wave by wave/particle duality. Klein we defined the ¹⁵⁶⁷ Hamiltonian as follows:

$$H = \frac{1}{2} a^{\bar{\mu}\bar{\nu}} p_{\bar{\mu}} p_{\bar{\nu}} \qquad \text{where} \qquad p_{\bar{\mu}} = \partial_{\bar{\mu}} S \quad . \tag{A.6}$$

Hence, equation (A.5) now reads:

$$H = 0 , \qquad (A.7)$$

and parametrizing the five-dimensional particle's world line with an arbitrary parameter $\hat{\lambda}$, the relativistic Hamilton equations are:

$$\frac{\partial H}{\partial p_{\bar{\mu}}} = \frac{dx^{\bar{\mu}}}{d\hat{\lambda}} \qquad ; \qquad -\frac{\partial H}{\partial x^{\bar{\mu}}} = \frac{dp_{\bar{\mu}}}{d\hat{\lambda}} . \tag{A.8}$$

The analogy between equation (A.5) and the usual eikonal equation suggests to consider null-geodesics for the differential form $a_{\bar{\mu}\bar{\nu}}dx^{\bar{\mu}}dx^{\bar{\nu}}$ as stated by Klein, where $a_{\bar{\mu}\bar{\nu}}$ represent the reciprocal quantities of $a^{\bar{\mu}\bar{\nu}}$. As emphasized in the main text, After a Legendre transformation, the Hamiltonian H is mapped into the following Lagrangian:

$$L = \frac{1}{2} a_{\bar{\mu}\bar{\nu}} \frac{dx^{\bar{\mu}}}{d\hat{\lambda}} \frac{dx^{\bar{\mu}}}{d\hat{\lambda}} , \qquad (A.9)$$

¹⁵⁷⁶ where the covariant components of the tensor $a_{\mu\bar{\nu}}$ are:

$$a_{\mu\nu} = g_{\mu\nu} + \frac{e^2}{m^2 c^4} A_{\mu} A_{\nu} \qquad a_{\mu5} = \frac{e^2}{m^2 c^3 \beta} A_{\mu} \qquad a_{55} = \frac{e^2}{m^2 c^4 \beta^2} \ . \tag{A.10}$$

Like all the quantities introduced by Klein, also the components of $a_{\bar{\mu}\bar{\nu}}$ do not depend on the fifth coordinate. As we emphasized in the main text, $a_{\bar{\mu}\bar{\nu}}$ and $\gamma_{\bar{\mu}\bar{\nu}}$ are quite different, cf. equation (A.10) and equation (A.3). As we said, it seems that Klein introduced a new metric for the microscopic world, $a_{\bar{\mu}\bar{\nu}}$, indeed the null-like character of the paths is referred to the tensor $a_{\bar{\mu}\bar{\nu}}$ instead of $\gamma_{\bar{\mu}\bar{\nu}}$. If following Klein we define $\mu = a_{55}$, hence $a_{\bar{\mu}\bar{\nu}}dx^{\hat{\mu}}dx^{\hat{\nu}} = \mu d\theta^2 + ds^2$. After having defined the tangent vector along the null-path, $V^{\mu} = \frac{dx^{\mu}}{d\hat{\lambda}}$, it should satisfy the condition $\mu \left(\frac{d\theta}{d\hat{\lambda}}\right)^2 + \left(\frac{ds}{d\hat{\lambda}}\right)^2 = 0$. The Hamilton equations are equivalent to the Euler-Lagrange equations:

$$\frac{d}{d\hat{\lambda}}\frac{\partial L}{\partial\left(dx^{\bar{\mu}}/d\hat{\lambda}\right)} - \frac{\partial L}{\partial x^{\bar{\mu}}} = 0.$$
(A.11)

We now skip some technical details, because a similar derivation is proposed in Appendix C, discussing de Broglie's approach. The equation for the fifth component is a conservation law, because the tensor $a_{\bar{\mu}\bar{\nu}}$ does not depend on the fifth coordinate x^5 . The conserved momentum p_5 reads $p_5 = \frac{\partial L}{\partial \left(dx^5/d\hat{\lambda}\right)} = \mu \frac{d\theta}{d\hat{\lambda}}$. This conservation law can be used to reduce equation (A.11), with $\bar{\mu} = 0, 1, 2, 3$, to:

$$mc\left(\frac{d}{d\hat{\lambda}}\left(g_{\mu\nu}V^{\nu}\right) - \frac{1}{2}\partial_{\mu}g_{\rho\nu}V^{\rho}V^{\nu}\right) = -\frac{e}{c}\left(\partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}\right)V^{\nu}.$$
 (A.12)

Klein now introduces the particle's proper time τ as follows. The constancy of p_5 and the condition for the null-like character of the paths imply that the ratio $\frac{d\tau}{d\hat{\lambda}}$ is constant along the path. Hence, in the projected four-dimensional equation (A.12), the arbitrary parameter can be substituted with the proper time, notwithstanding we started considering null-geodesics.⁶⁶ After some manipulation it can be shown that it is equivalent to the Lorentz equation for a charged massive particle of mass m and charge -e in a combined electromagnetic and gravitational field (see Appendix C):

$$mc\left(\frac{du^{\lambda}}{d\tau} + \Gamma^{\lambda}_{\varrho\nu}u^{\rho}u^{\nu}\right) = -\frac{e}{c}F^{\lambda}{}_{\nu}u^{\nu}, \qquad (A.13)$$

where now $u^{\mu} = \frac{dx^{\mu}}{d\tau}$ is the particle's four-velocity. We stress again the role of the tensor $a_{\bar{\mu}\bar{\nu}}$. The mass of the particle is hidden into its definition, equation (A.10). Therefore, the five-dimensional null-geodesics for the differential form $a_{\bar{\mu}\bar{\nu}}dx^{\bar{\mu}}dx^{\bar{\nu}}$ are connected with four-dimensional geodesics of charged massive particles.

Appendix B On the inconsistency of a time-like compactified dimension

One of the most important assertion we made in the text is that, unlike Klein, de Broglie considered a time-like fifth dimension. In order to understand the consequences of this choice we start again with the five-dimensional line element $d\sigma^2 =$ $\gamma_{\bar{\mu}\bar{\nu}}dx^{\bar{\mu}}dx^{\bar{\nu}}$. Using Klein notation, which we rewrite here for convenience, we define $\gamma_{5\mu} = \beta A_{\mu}$ and the components of the five-dimensional metric are:

$$\gamma_{\mu\nu} = g_{\mu\nu} + \alpha \beta^2 A_{\mu} A_{\nu} , \qquad \gamma_{55} = \alpha , \qquad \gamma_{5\mu} = \alpha \beta A_{\mu} . \tag{B.1}$$

This metric has the following signature: $(-; +; +; +; \epsilon)$, where $\epsilon = +$ if $\alpha > 0$, i.e. if the fifth coordinate describes a space-like dimension, and $\epsilon = -$ if $\alpha < 0$, i.e. in the case of a time-like coordinate. We remember that the line element can be rewritten as $d\sigma^2 = \alpha d\theta^2 + ds^2$, where $d\theta = dx^5 + \beta A_{\mu}dx^{\mu}$ and $ds^2 = g_{\mu\nu}dx^{\mu}dx^{\nu}$. The components of the inverse metric are:

$$\gamma^{\mu\nu} = g^{\mu\nu} , \qquad \gamma^{55} = \frac{1}{\alpha} + \beta^2 A_{\mu} A^{\mu} , \qquad \gamma^{5\mu} = -\beta A^{\mu} .$$
 (B.2)

¹⁶¹³ Using the Ansatz that the metric does not depend on the fifth coordinate, we have ¹⁶¹⁴ calculated the components of the five-dimensional Ricci tensor, defined by

$$\tilde{R}_{\bar{\mu}\bar{\nu}} = \partial_{\bar{\lambda}}\tilde{\Gamma}^{\bar{\lambda}}_{\bar{\mu}\bar{\nu}} - \partial_{\bar{\nu}}\tilde{\Gamma}^{\bar{\lambda}}_{\bar{\mu}\bar{\lambda}} + \tilde{\Gamma}^{\bar{\lambda}}_{\bar{\sigma}\bar{\lambda}}\tilde{\Gamma}^{\bar{\sigma}}_{\bar{\mu}\bar{\nu}} - \tilde{\Gamma}^{\bar{\lambda}}_{\bar{\sigma}\bar{\nu}}\tilde{\Gamma}^{\bar{\sigma}}_{\bar{\mu}\bar{\lambda}} .$$
(B.3)

⁶⁶It is worth remembering that the proper-time cannot be defined for null-geodesics.

¹⁶¹⁵ We need the following results:

$$\tilde{R}_{55} = \frac{\alpha^2 \beta^2}{4} F_{\mu\nu} F^{\mu\nu} , \qquad (B.4)$$

$$\tilde{R}_{5\sigma} = \alpha \beta \nabla_{\lambda} F_{\sigma}^{\ \lambda} + \frac{\alpha^2 \beta^3}{4} A_{\sigma} F_{\mu\nu} F^{\mu\nu} , \qquad (B.5)$$

$$g^{\mu\nu}\tilde{R}_{\mu\nu} = R + \frac{\alpha^2\beta^4}{4}A_{\sigma}A^{\sigma}F_{\mu\nu}F^{\mu\nu} - \frac{\alpha\beta^2}{2}F_{\mu\nu}F^{\mu\nu} + \alpha\beta^2A_{\mu}\nabla_{\lambda}F^{\mu\lambda}, \quad (B.6)$$

that lead to the following relation for the five-dimensional curvature scalar:

$$\tilde{R} = \gamma^{\bar{\mu}\bar{\nu}} \tilde{R}_{\bar{\mu}\bar{\nu}} = \gamma^{55} \tilde{R}_{55} + 2\gamma^{5\mu} \tilde{R}_{5\mu} + \gamma^{\mu\nu} \tilde{R}_{\mu\nu} = R - \frac{\alpha\beta^2}{4} F_{\mu\nu} F^{\mu\nu} .$$
(B.7)

Equation (B.7) shows that if the fifth dimension is space-like, $\alpha > 0$, we can identify $\alpha\beta^2 = 2\kappa$ and the electromagnetic kinetic term has the correct sign. On the contrary, if α is negative this identification is not possible. This is the inconsistency connected with a compactified time-like dimension. As written in the main text, Klein inferred from this fact the need to introduce a space-like compact dimension.

¹⁶²² Appendix C Geodesics in de Broglie-Rosenfeld approach

In this section we describe de Broglie's analysis of five-dimensional geodesics, with 1623 some details. After having introduced the five-dimensional metric, in the fifth para-1624 graph of his paper de Broglie considered all five-dimensional geodesics, not only 1625 null-geodesics as suggested by Klein, with the following motivation: 'Admitting the 1626 existence of a fifth dimension of the Universe, we could enunciate the following prin-1627 ciple: *IIIn the five-dimensional universe, the World-line of every point particle is a* 1628 geodesic;; ([Louis de Broglie, 1927b]; p. 69). Given O and M, 'two fixed points of 1629 the World-line' ([Louis de Broglie, 1927b]; p. 69), five-dimensional geodesics can be 1630 seen as world-lines of extremal "five-dimensional proper time" $d\hat{\tau} = \sqrt{-d\sigma^2}$: 1631

$$\delta \int_{O}^{M} d\hat{\tau} = 0.$$
 (C.1)

After introducing an arbitrary parameter $\hat{\lambda}$, the geodesic equation can be obtained equivalently by the following variational principle:

$$\frac{1}{2}\delta\int_{O}^{M} \left[\gamma_{\bar{\mu}\bar{\nu}}\frac{dx^{\hat{\mu}}}{d\hat{\lambda}}\frac{dx^{\hat{\nu}}}{d\hat{\lambda}}\right] d\hat{\lambda} = \frac{1}{2}\delta\int_{O}^{M} \left[\alpha\left(\frac{d\theta}{d\hat{\lambda}}\right)^{2} + g_{\mu\nu}\frac{dx^{\mu}}{d\hat{\lambda}}\frac{dx^{\nu}}{d\hat{\lambda}}\right] d\hat{\lambda} = 0 \quad \text{i.e.}$$
$$\frac{1}{2}\delta\int_{O}^{M} \left[\alpha\left(V^{5} + \beta A_{\mu}V^{\mu}\right)^{2} + g_{\mu\nu}V^{\mu}V^{\nu}\right] d\hat{\lambda}$$
$$= 0, \qquad (C.2)$$

where we used $d\sigma^2 = \gamma_{\hat{\mu}\hat{\nu}}dx^{\hat{\mu}}dx^{\hat{\nu}} = \alpha d\theta^2 + ds^2$ and where V^5 and V^{μ} are the five components of the five-velocity $V^{\bar{\mu}} = \frac{dx^{\bar{\mu}}}{d\hat{\lambda}}$. Now de Broglie identified the quantity into the square bracket as a Lagrangian L(x, V). Varying the action as a function of $x^{\bar{\mu}}$ and $V^{\bar{\mu}}$, de Broglie obtained the following Euler-Lagrange equations:

$$\frac{d}{d\hat{\lambda}}\frac{\partial L}{\partial V^5} = \frac{\partial L}{\partial x^5} , \qquad (C.3a)$$
$$\frac{d}{d\hat{\lambda}}\frac{\partial L}{\partial V^{\mu}} = \frac{\partial L}{\partial x^{\mu}} . \qquad (C.3b)$$

Remembering that there is no dependence from the fifth dimension, the equation (C.3a) produces a conserved quantity:

$$\frac{d}{d\hat{\lambda}}\alpha\left(V^5 + \beta A_{\mu}V^{\mu}\right) = 0 \qquad \text{i.e.} \qquad \pi_5 = \alpha \frac{d\theta}{d\hat{\lambda}} = \text{constant} , \qquad (C.4)$$

while equation (C.3b) read⁶⁷:

$$\frac{d}{d\hat{\lambda}}\left(\pi_5\beta A_{\mu} + g_{\mu\nu}V^{\nu}\right) = \frac{1}{2}\partial_{\mu}g_{\rho\sigma}V^{\rho}V^{\sigma} + \pi_5\beta\partial_{\mu}A_{\nu}V^{\nu} , \qquad (C.5)$$

and, rearranging the terms and inserting π_5 expression (C.4), its equivalent form is:

$$\frac{d}{d\hat{\lambda}}\left(g_{\mu\nu}\frac{dx^{\nu}}{d\hat{\lambda}}\right) = \frac{1}{2}\partial_{\mu}g_{\rho\sigma}\frac{dx^{\rho}}{d\hat{\lambda}}\frac{dx^{\sigma}}{d\hat{\lambda}} + \alpha\frac{d\theta}{d\hat{\lambda}}\beta F_{\mu\rho}\frac{dx^{\rho}}{d\hat{\lambda}} \,. \tag{C.6}$$

We can now introduce the proper-time $d\tau = \sqrt{-ds^2}$, because we are considering nonnull geodesics. The five-dimensional geodesic equation and the metricity condition imply that the covariant derivative of the $\gamma_{\bar{\mu}\bar{\nu}}V^{\bar{\mu}}V^{\bar{\nu}}$ would be zero. Hence the ratio $\frac{d\hat{\lambda}}{d\tau}$ is constant along the geodesic curve and in equation (C.6) $\hat{\lambda}$ could be substituted by τ . If we define the normalized four-dimensional vector $u^{\mu} = \frac{dx^{\mu}}{d\tau}$ and if we set, following de Broglie,

$$\alpha \frac{d\theta}{d\tau} = -\frac{e}{\beta c} \frac{1}{mc} , \qquad (C.7)$$

1648 equation (C.6) reduces to

$$mc\left(\frac{d}{d\tau}\left(g_{\mu\nu}u^{\nu}\right) - \frac{1}{2}\partial_{\mu}g_{\rho\nu}u^{\rho}u^{\nu}\right) = -\frac{e}{c}F_{\mu\nu}u^{\nu}.$$
 (C.8)

As claimed in the main text, the parameter β disappears and it remains undetermined. In order to obtain Lorentz equations we rewrite the first term of the l.h.s. of equation (C.8) as follows:

$$\frac{d}{d\tau} \left(g_{\mu\nu} u^{\nu} \right) = u^{\rho} \partial_{\rho} \left(g_{\mu\nu} u^{\nu} \right) = g_{\mu\nu} \frac{du^{\nu}}{d\tau} + \frac{1}{2} \left(\partial_{\rho} g_{\mu\nu} + \partial_{\nu} g_{\mu\rho} \right) , \qquad (C.9)$$

⁶⁷Remember that A_{μ} is a function of the four-dimensional coordinates.

Finally, we insert the previous equation in (C.8) and we contract it with the inverse 1652 components of the metric $q^{\lambda\mu}$ to get the Lorentz equations: 1653

$$mc\left(\frac{du^{\lambda}}{d\tau} + \Gamma^{\lambda}_{\mu\rho}u^{\rho}u^{\nu}\right) = -\frac{e}{c}F^{\lambda}{}_{\nu}u^{\nu}.$$
 (C.10)

Unlike Klein, de Broglie's purpose was to show how the five-dimensional Universe's 1654 approach permits to geometrize the electromagnetic force. He stressed: 'This means 1655 that with geometric meaning we have attributed to the [electromagnetic] poten-1656 tials and to the ratio e/m, the five-dimensional World-lines of point particles are all 1657 geodesics. The notion of force has been completely banned from Mechanics.' ([Louis 1658 de Broglie, 1927b; p. 70). This connection between geodesic lines and equation (C.8) 1659 convinced de Broglie that was not necessary to consider null-geodesics lines only. 1660

De Broglie's investigation of five-dimensional geodesic lines continued with the 1661 question of what would be the correct particle's action in five dimensions. The author 1662 proposed ([Louis de Broglie, 1927b]; p. 70): 1663

$$S_5 = -\mathcal{I} \int_O^M d\hat{\tau} \,, \tag{C.11}$$

where 1664

$$\mathcal{I} = m^2 c^2 - \frac{e^2}{\alpha \beta^2 c^2} \,, \tag{C.12}$$

because it reduces to the usual point particle action in the case of zero charge. In 1665 order to understand this fact, following de Broglie, we point out that we can rewrite 1666 S_5 as follows: 1667

$$S_5 = -\mathcal{I} \int_O^M d\hat{\tau} = \int_O^M \left(\mathcal{I} \alpha \frac{d\theta}{d\hat{\tau}} \right) d\theta - \int_O^M \left(\mathcal{I} \frac{d\tau}{d\hat{\tau}} \right) d\tau , \qquad (C.13)$$

where the second equality sign follows by inserting $1 = \left(\frac{d\hat{\tau}}{d\hat{\tau}}\right)^2 = \left(\frac{d\tau}{d\hat{\tau}}\right)^2 - \alpha \left(\frac{d\theta}{d\hat{\tau}}\right)^2$, 1668 as a formal consequence of $d\hat{\tau}^2 = d\tau^2 - \alpha d\theta^2$. We remember that the condition 1669 $\partial_5 \gamma_{\bar{\mu}\bar{\nu}} = 0$ is equivalent to assert, using modern language, that space-time would 1670 admit a Killing vector field, which is tangent to the fifth coordinate. The scalar prod-1671 uct between the Killing field and the velocity field is constant along the geodesic. This result implies that the ratio $\frac{d\theta}{d\hat{\tau}}$ must be constant. Hence, de Broglie chose: 1672

1673

$$\mathcal{I}\frac{d\tau}{d\hat{\tau}} = mc \tag{C.14}$$

and 1674

$$\mathcal{I}\alpha \frac{d\theta}{d\hat{\tau}} = -\frac{e}{c\beta} , \qquad (C.15)$$

which are consistent with equation (C.7). Finally, using $d\theta = dx^5 + \beta A_{\mu}dx^{\mu}$, S_5 1675 assumes the following form: 1676

$$S_5 = -\int \frac{e}{c\beta} dx^5 + \frac{e}{c} \int A_\mu dx^\mu - mc \int d\tau . \qquad (C.16)$$

It is now evident that S_5 reduces to $S_4 = -mc \int d\tau$ when we set e = 0. Indeed, when e = 0 the scalar product between the Killing field and the velocity field (C.15) (cf. de Broglie's comment on equation (20)) is zero, then the geodesic projects onto the four-dimensional space-time. As a consequence, de Broglie convinced himself that the invariant \mathcal{I}^2 should have been a five-dimensional generalization of the fourdimensional momentum⁶⁸ mc. At this stage, we are able to explain the form of the invariant \mathcal{I} . Equations (C.14) and (C.15) and the identity $d\hat{\tau}^2 = d\tau^2 \alpha d\theta^2$ imply that \mathcal{I} must have the following form:

$$\mathcal{I}^2 = m^2 c^2 - \frac{e^2}{\alpha \beta^2 c^2} \,. \tag{C.17}$$

As noted by Klein, de Broglie choose $-\alpha\beta^2 = 2\kappa$, because, if the fifth dimension is time-like, α is negative and the invariant \mathcal{I}^2 would be strictly positive. On the other hand, as we have said, the choice is not consistent with the request to obtain Maxwell Lagrangian in (B.7).

As we have said in the main text, Klein asserted that de Broglie's mistake did not affect his conclusions. Klein referred to the following fact. De Broglie proposed the five-dimensional wave equation:

$$\gamma^{\bar{\mu}\bar{\nu}}\nabla_{\bar{\mu}}\partial_{\bar{\nu}}\Psi = \frac{4\pi^2}{h^2}\mathcal{I}^2\Psi.$$
 (C.18)

It is worth noting that S_5 depends linearly from x^5 , as we can see integrating (C.16). Hence, using a geometrical optics Ansatz $\Psi = Ae^{\frac{i}{\hbar}S_5}$, the periodicity with respect to x^5 follows immediately.

¹⁶⁹⁵ This means that the wave function can be written as:

$$\Psi(x) = \psi\left(x^0, x^1, x^2, x^3\right) \cdot \exp\left(\frac{i}{\hbar} \frac{e}{c\beta} x^5\right) , \qquad (C.19)$$

where ψ is the four-dimensional wave function. Using (C.19) and the components of the inverse metric (B.2), we note, following Klein ([Klein, 1927a]; p. 243) that, Ψ satisfies

$$\gamma^{55}\partial_5^2\Psi = -\frac{1}{\hbar^2} \left(\frac{1}{\alpha} + \beta^2 A_\mu A^\mu\right) \frac{e^2}{c^2\beta^2}\Psi .$$
 (C.20)

This means that (C.18) can be rewritten, in a flat space-time, in the following way ([Louis de Broglie, 1927b]; p. 73):

$$g^{\mu\nu}\partial_{\mu}\partial_{\nu}\psi - \frac{2ie}{\hbar c}A^{\mu}\partial_{\mu}\psi - \frac{e^2}{\hbar^2 c^2}A_{\mu}A^{\mu}\psi = \frac{m^2 c^2}{\hbar^2}\psi.$$
(C.21)

We note that (C.21) corresponds to the KG equation for a complex scalar field in an external electromagnetic field, and can be written in the following compact way:

$$g^{\mu\nu} \left(\partial_{\mu} - \frac{i}{\hbar} \frac{e}{c} A_{\mu}\right) \left(\partial_{\nu} - \frac{i}{\hbar} \frac{e}{c} A_{\nu}\right) \psi = \frac{m^2 c^2}{\hbar^2} \psi , \qquad (C.22)$$

⁶⁸We remember that De Broglie's idea emerged comparing S_4 with S_5 .

if the Lorenz gauge, namely $\partial_{\mu}A^{\mu} = 0$, holds. As claimed by Klein, independently to 1703 the character of the fifth dimension, the term depending on $\alpha\beta^2$ in \mathcal{I}^2 definition (C.17) 1704 disappears, and equation (C.21) reduces to de Broglie's equation ([Louis de Broglie, 1705 1927b]; p. 73, Eq. (40)), where the case of null gravitational field is considered. 1706

Appendix D On Rosenfeld approach 1707

In this section we explain some technical details skipped in the main text. 1708

D.1 Five-dimensional versus four-dimensional momentum 1709

Equations (58) and (59) can be obtained as follows. First we note that if S_0 is a 1710 complete integral of the HJ equation in four dimensions, see equation (54), it fol-1711 lows that⁶⁹ $g^{\mu\nu} \left(\partial_{\nu}S_0 + \frac{e}{c}A_{\nu}\right) = mc\frac{dx^{\mu}}{d\tau}$. Then, using the inverse components of the metric tensor (53) and equation (52), we rewrite the l.h.s. of (58) as follows: 1712 1713

$$\gamma^{\mu\bar{\nu}}\partial_{\bar{\nu}}\bar{S} = \gamma^{\mu5}\partial_{5}\bar{S} + \gamma^{\mu\nu}\partial_{\nu}\bar{S} = (-\beta A^{\mu})\left(-\frac{e}{c\beta}\right) + g^{\mu\nu}\partial_{\nu}S_{0}$$
$$= g^{\mu\nu}\left(\partial_{\nu}S_{0} + \frac{e}{c}A_{\nu}\right) = mc\frac{dx^{\mu}}{d\tau}, \qquad (D.1)$$

and we have finally obtained equation (58). Since Rosenfeld introduced explicitly the 1714 quantity $\sqrt{m^2c^2 - \frac{e^2c^2}{16\pi G}}$, we used for this quantity the symbol \mathcal{I}_{Ros} for brevity. From 1715 equation (57) we get 1716

$$\gamma^{\mu\bar{\nu}}\partial_{\bar{\nu}}\bar{S} = \mathcal{I}_{Ros}\frac{dx^{\mu}}{d\hat{\tau}} = \mathcal{I}_{Ros}\frac{dx^{\mu}}{d\tau}\frac{d\tau}{d\hat{\tau}}\,,\tag{D.2}$$

and confronting equation (D.2) with (58) we get equation (59). 1717

D.2 Modern five-dimensional action 1718

In action (64) Rosenfeld choose an unusual coupling between matter and gravity. 1719 Rosenfeld's coupling is unusual for the following reason. In a modern five-dimensional 1720 approach, the action would be: 1721

$$\mathcal{S}_{tot}\left(\gamma, \Phi, \bar{\Phi}\right) = \int d^5 x \sqrt{-\gamma} \left[-\frac{1}{2\kappa_5} \tilde{R} + \tilde{\mathcal{L}} \right] \,, \tag{D.3}$$

where \mathcal{L} is the action for a complex scalar field Φ , that has the expected length dimen-1722 sion $[\Phi] = (length)^{-\frac{3}{2}}$, in natural units $\hbar = c = 1$. Using the determinant definition 1723 and (B.1) it can be proved that⁷⁰ 1724

$$\gamma = \epsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{\sigma}\lambda}\gamma_{\bar{\mu}0}\gamma_{\bar{\nu}1}\gamma_{\bar{\rho}2}\gamma_{\bar{\sigma}3}\gamma_{\bar{\lambda}5} = \alpha g .$$
 (D.4)

⁶⁹See Landau & Lifšhitz [1951]. ⁷⁰We define $\epsilon^{01235} = 1$.

Using $\kappa_5 = 2\pi \tilde{l}\kappa$, where $2\pi \tilde{l}$ is the "volume" of the compact dimension, (D.3) can be rewritten as follows:

$$\mathcal{S}_{tot}\left(\gamma, \Phi, \bar{\Phi}\right) = \frac{\sqrt{\alpha}}{2\kappa_5} \int d^5 x \sqrt{-g} \left[-\tilde{R} + \kappa \left(2\pi \tilde{l}\tilde{\mathcal{L}}\right)\right] \,. \tag{D.5}$$

Now the length \tilde{l} of the fifth dimensions can be adsorbed with the following field redefinition: $\Psi = \sqrt{2\pi \tilde{l}} \Phi$. This shows that the equations obtained by varying (D.5) are equivalent to Rosenfeld's equations of motion, but the new scalar field Ψ has length dimensions $[\Psi] = (length)^{-1}$ as a four-dimensional scalar field. As a consequence, as stated in the main text, the stress-energy tensor defined by Rosenfeld is a fourdimensional object.

1733 D.3 Einstein–Maxwell equations coupled with complex scalar field

¹⁷³⁴ The equations obtained by varying (D.5) with respect to the metric are:

$$\tilde{R}_{\bar{\mu}\bar{\nu}} - \frac{1}{2}\gamma_{\bar{\mu}\bar{\nu}}\tilde{R} = \kappa T_{\bar{\mu}\bar{\nu}} , \qquad (D.6)$$

and, as written in the main text, they are formally equivalent to the the fourdimensional Einstein equations, coupled to the electromagnetic and matter stressenergy tensor, and Maxwell equations. In order to understand this fact, firstly we write the expression for $\tilde{R}_{\mu\nu}$. After a lengthy calculation, from (B.3) it follows:

$$\tilde{R}_{\mu\nu} = R_{\mu\nu} + \frac{\alpha^2 \beta^4}{4} A_{\mu} A_{\nu} F_{\sigma\lambda} F^{\sigma\lambda} - \frac{\alpha \beta^2}{2} F_{\mu\lambda} F_{\nu}{}^{\lambda} + \frac{\alpha \beta^2}{2} \left(A_{\mu} \nabla_{\lambda} F_{\nu}{}^{\lambda} + A_{\nu} \nabla_{\lambda} F_{\mu}{}^{\lambda} \right) . \tag{D.7}$$

¹⁷³⁹ Let us consider the contravariant components of (D.6), i.e.

$$\gamma^{\bar{\lambda}\bar{\mu}}\gamma^{\bar{\sigma}\bar{\nu}}\tilde{R}_{\bar{\mu}\bar{\nu}} - \frac{1}{2}\gamma^{\bar{\lambda}\bar{\sigma}}\tilde{R} = \kappa\gamma^{\bar{\lambda}\bar{\mu}}\gamma^{\bar{\sigma}\bar{\nu}}T_{\bar{\mu}\bar{\nu}} .$$
(D.8)

¹⁷⁴⁰ Using (B.1), (B.2), (B.4), (B.5) and (D.7) the $\lambda\sigma$ components of the l.h.s. of equation ¹⁷⁴¹ (D.8) can be rewritten as follows:

$$\gamma^{\lambda\bar{\mu}}\gamma^{\sigma\bar{\nu}}\tilde{R}_{\bar{\mu}\bar{\nu}} - \frac{1}{2}\gamma^{\lambda\sigma}\tilde{R} = \begin{bmatrix} g^{\lambda\mu}g^{\sigma\nu}\tilde{R}_{\mu\nu} + g^{\lambda\mu}\gamma^{\sigma5}\tilde{R}_{5\mu} + \gamma^{\lambda5}g^{\sigma\nu}\tilde{R}_{5\nu} + \gamma^{\lambda5}\gamma^{\sigma5}\tilde{R}_{55} \end{bmatrix} - \frac{1}{2}g^{\lambda\sigma}\tilde{R},$$
$$= R^{\lambda\sigma} - \frac{1}{2}g^{\lambda\sigma}R - \kappa g^{\lambda\mu}g^{\sigma\nu}T^{em}_{\mu\nu}.$$
(D.9)

1742 Following Rosenfeld we define

$$T^{\lambda\sigma} = \gamma^{\lambda\bar{\mu}}\gamma^{\sigma\bar{\nu}}T_{\bar{\mu}\bar{\nu}} , \qquad (D.10)$$

and the $\lambda\sigma$ components of (D.8) read ([Léon Rosenfeld, 1927a]; p. 313):

$$R^{\lambda\sigma} - \frac{1}{2}g^{\lambda\sigma}R = \kappa \left(T_{em}^{\lambda\sigma} + T^{\lambda\sigma}\right) , \qquad (D.11)$$

that correspond to Einstein equations coupled to the electromagnetic and the matter stress-energy tensor. Maxwell equations emerge conversely as follows. If we contract (D.6) with $\gamma^{\bar{\rho}\bar{\mu}}$, we get:

$$\gamma^{\bar{\rho}\bar{\mu}}\tilde{R}_{\bar{\mu}\bar{\nu}} - \frac{1}{2}\delta^{\bar{\rho}}{}_{\bar{\nu}}\tilde{R} = \kappa\gamma^{\bar{\rho}\bar{\mu}}T_{\bar{\mu}\bar{\nu}} .$$
(D.12)

¹⁷⁴⁷ The ρ 5 components of the l.h.s. of equation (D.12) now read⁷¹:

$$\gamma^{\rho\bar{\mu}}\tilde{R}_{\bar{\mu}\bar{\nu}} = \gamma^{\rho\mu}\tilde{R}_{\mu5} + \gamma^{\rho5}\tilde{R}_{55} ,$$

$$= \frac{\alpha\beta}{2}\nabla_{\lambda}F^{\rho\lambda} .$$
(D.13)

1748 Remembering that $\kappa = \frac{\alpha \beta^2}{2}$, following Rosenfeld, we define

$$T^{\rho}{}_{5} = \gamma^{\rho\bar{\mu}} T_{\bar{\mu}5} , \qquad (D.14)$$

and equation (D.12) now reads:

$$\nabla_{\lambda} F^{\rho\lambda} = \beta T^{\rho}{}_5 . \tag{D.15}$$

Equation (D.15) correspond to Maxwell equations coupled to a current density as written by Rosenfeld ([Léon Rosenfeld, 1927a]; p. 313).

1752 D.4 Four-dimensional and five-dimensional curvature scalar

In the main text, we have written that using (D.6) Rosenfeld obtained a particular relation for the curvature scalars R and \tilde{R} , namely

$$R = -\kappa \left[\gamma^{\nu \bar{\mu}} T_{\bar{\mu}\nu} - \gamma^{\mu\rho} A_{\rho} \nabla_{\lambda} \left(\gamma_{\mu\sigma} F^{\sigma\lambda} \right) \right] \quad \text{and} \tag{D.16}$$

$$\tilde{R} = -\kappa \left[\gamma^{\nu \bar{\mu}} T_{\bar{\mu}\nu} + \frac{F_{\sigma\lambda} F^{\sigma\lambda}}{2} - \gamma^{\mu\rho} A_{\rho} \nabla_{\lambda} \left(\gamma_{\mu\sigma} F^{\sigma\lambda} \right) \right].$$
(D.17)

In order to obtain these relations, we set $\bar{\rho} = \bar{\nu} = \nu$ in equation (D.12) and it reads:

$$\gamma^{\nu\bar{\mu}}\tilde{R}_{\bar{\mu}\nu} - 2\tilde{R} = \kappa T^{\nu}{}_{\nu} , \qquad (D.18)$$

where we have defined $T^{\nu}{}_{\nu} = \gamma^{\nu\bar{\mu}}T_{\bar{\mu}\nu}$. Using the definition of \tilde{R} (Eq. (B.7)), equation (D.18) can be rewritten as

$$\tilde{R} - \gamma^{55} \tilde{R}_{55} - \gamma^{5\mu} \tilde{R}_{\mu 5} - 2\tilde{R} = \kappa T^{\nu}{}_{\nu} .$$
(D.19)

Inserting (B.2), (B.4), (B.5) and (B.7), and isolating R, we obtain equation (D.16) and using again (B.7) we obtain (D.17).

⁷¹Remember that $\delta^{\bar{\rho}}_{\bar{\nu}} = 0$ when $\bar{\rho} \neq \bar{\nu}$.

1760 D.5 The retarded potentials

After having linearised Einstein equation (D.6), Rosenfeld integrated it and obtained the retarded potentials, equation (71). Using modern notation the retarded potentials read:

$$h_{\bar{\mu}\nu}(t;\mathbf{x}) = -\frac{\kappa}{2\pi} \int_{\Sigma} \bar{T}_{\bar{\mu}\nu} \left(t - \frac{|\mathbf{x} - \mathbf{y}|}{c}; \mathbf{y} \right) \frac{d^3 y}{|\mathbf{x} - \mathbf{y}|} \quad , \tag{D.20}$$

where the radial distance is defined by $r = |\mathbf{x} - \mathbf{y}|$ and the integration is carried on a three-dimensional hypersurface Σ at the retarded time $t - \frac{|\mathbf{x} - \mathbf{y}|}{c}$. The retarded potential are functions of \mathbf{x} and t.

1767 D.6 The isotropic coordinate system and the "mean distance"

¹⁷⁶⁸ In this last appendix we show how Rosenfeld was inspired by his knowledge of ¹⁷⁶⁹ Eddington's book on GR.

Given a bounded charged matter distribution of radius ϵ , the RN metric is an exact solution of equation (D.11), with $T^{\lambda\sigma}$ being the stress-energy tensor associated to the classical spherical symmetric mass distribution. In polar coordinates the line element has the following form:

$$ds^{2} = -\left(1 - \frac{2mG}{c^{2}r} + \frac{GQ^{2}}{c^{4}r^{2}}\right)c^{2}dt^{2} + \left(1 - \frac{2mG}{c^{2}r} + \frac{GQ^{2}}{c^{4}r^{2}}\right)^{-1}dr^{2} + r^{2}\left(d\theta^{2} + \sin^{2}\theta d\varphi^{2}\right),$$
(D.21)

where m and Q are the mass and the charge of the particle respectively and the coor-1774 dinate r has the following range: $\epsilon \leq r < +\infty$. If Q = 0 the line element describes 1775 the so called exterior Schwartzschild metric. Rosenfeld used the less known isotropic 1776 coordinate system. We do not know if the author would know RN metric in isotropic 1777 coordinates. As stated in Section 2, we know from Kuhn's interview [Kuhn & Heil-1778 bron, 1963 that Rosenfeld studied Eddington's book on GR. In The Mathematical 1779 Theory of Relativity [Eddington, 1923] the British Physicist introduced isotropic coor-1780 dinates for Schwartzschild metric, using both its exact form and its limit at first order 1781 in $\frac{1}{r}$. It is worth noting that at asymptotically large distances from the source, at the 1782 first order in $\frac{1}{r}$, both Schwartzschild and RN metric have the same form. This fact 1783

r, some order in r, some order isotropic and in polar coordinates. 1784 is true both in isotropic and in polar coordinates. 1785 In the so called isotropic Cartesian coordinate system the line element of a

¹⁷⁸⁵ In the so called isotropic Cartesian coordinate system the line element of a ¹⁷⁸⁶ spherically symmetric space-time has the following form:

$$ds^{2} = -A(r) dt^{2} + B(r) \left(dx^{2} + dy^{2} + dz^{2} \right) , \qquad (D.22)$$

where $r = \sqrt{x^2 + y^2 + z^2}$ is the distance from the origin. Following Eddington, at the first order in $\frac{1}{r}$, for a point-particle continually at rest we have ([Eddington, 1923]; p. 101):

$$A(r) \approx 1 - \frac{2mG}{c^2 r}$$
 and $B(R) \approx 1 + \frac{2mG}{c^2 r}$, (D.23)

where the particle need not be at the origin provided that r is the distance from the particle to the point considered. The line element now reads:

$$ds^{2} = -\left(1 - \frac{2mG}{c^{2}r}\right)dt^{2} + \left(1 + \frac{2mG}{c^{2}r}\right)\left(dx^{2} + dy^{2} + dz^{2}\right), \quad (D.24)$$

showing that at large distances the particle's gravitational field is "less different" from the Minkowskian field, as stated by Rosenfeld.

Line element (D.24) and Rosenfeld's line element are different, see e.g. (80). Rosen-1794 feld used the "mean distance" $r_0(\vec{x})$ instead of r: Rosenfeld replaced the distance to 1795 the single particle by the mean distance to the cloud. In order to understand this 1796 fact, we remember that inspecting the semi-classical limit of his quantum metric the 1797 particle is represented by a wave function that is zero outside a volume V. For this 1798 reason, following Eddington, we consider the transition to continuous matter. Sum-1799 ming the fields of force of a number of particles, Eddington suggested the following 1800 form for the two functions A(R) and B(R): 1801

$$A(r) \approx 1 - \frac{2\Omega}{c^2}$$
 and $B(r) \approx 1 + \frac{2\Omega}{c^2}$, (D.25)

where Ω represents the Newton potential at the point considered and using Eddington notation reads:

$$\Omega = \sum \frac{m}{r} \ . \tag{D.26}$$

Let $\vec{y_i}$, with i = 1, ..., N, be the position of the i-th particle, m_i its mass and let \vec{x} be an arbitrary point of the space-time. Using modern notation, equation (D.26) reads:

$$\Omega = \sum_{i=1}^{N} \frac{m_i}{|\vec{y}_i - \vec{x}|} \ . \tag{D.27}$$

For a homogeneous system of mass m with volume V the Newton potential reads:

$$\Omega = \frac{m}{V} \int_{V} \frac{dxdydz}{|\vec{y} - \vec{x}|} , \qquad (D.28)$$

where \mathbf{y} is a point of the volume V. The mean value theorem states that:

 $\frac{1}{V} \int_{V} \frac{dxdydz}{|\mathbf{x} - \mathbf{y}|} = \frac{1}{r_0(\mathbf{x})}$ (D.29)

where $r_0(\vec{x})$ is the mean distance to the cloud. Equation (D.29) is equivalent to Rosenfeld's condition (86), namely $\frac{V}{r_0(\mathbf{x})} = \int_V \frac{dV}{|\mathbf{x} - \mathbf{y}|}$ and the line element to be compared with the semi-classical limit of the quantum metric reads:

$$ds^{2} = -\left(1 - \frac{2mG}{c^{2}r_{0}(\vec{x})}\right)dt^{2} + \left(1 + \frac{2mG}{c^{2}r_{0}(\vec{x})}\right)\left(dx^{2} + dy^{2} + dz^{2}\right).$$
 (D.30)

1811 References

- Ashtekar, Abhay and Robert Geroch. 1974. Quantum theory of gravitation, *Reports on Progress in Physics*, 37: 1211–56.
- Bacciagaluppi, Guido and Antony Valentini. 2009. Quantum Theory at the Crossroads.
 Reconsidering the 1927 Solvay Conference. Cambridge: Cambridge University Press.
- Bahrami, Mohammad, André Grossardt, Sandro Donadi and Angelo Bassi. 2014. The
 Schrödinger-Newton equation and its foundations, *New Journal of Physics*, 16: 115007.
- Birrel, Nicholas D. and Paul C. W. Davies. 1982. Quantum fields in Curved Space.
 Cambridge: Cambridge University Press.
- Blum, Alexander, Martin Jähnert, Christoph Lehner and Jürgen Renn. 2017. Translation
 as heuristics: Heisenber's turn to matrix mechanics, *Studies in History and Philosophy of Science Part B: Studies in History and Philosophy of Modern Physics*: 1–20.
- Bronstein, Matvei P. 1935. Quantentheorie schwacher Gravitationsfelder. Physikalische
 Zeitschrift der Sowjetunion, 9: 140–157 (1936).
- Carlip, Steven. 2008. Is Quantum Gravity necessary? Classical and Quantum Gravity, 25:
 154010.
- 1827 Chadwick, James. 1932. Possible existence of a neutron. Nature, 129: 312 1932.
- Darrigol, Olivier. 1992. From c-Numbers to q-Numbers: The Classical Analogy in the History
 of Quantum Theory. Berkeley: University of California Press.
- de Broglie, Louis. 1927. La mecanique ondulatoire et la structure atomique de la matière et du rayonnement. Comptes rendus hebdomadaires des séances de l'Académie des sciences, 1832 185: 380–382.
- de Broglie, Louis. 1927. L'univers a cinq dimensions et la mécanique ondulatoire. *Le Journal de Physique et le Radium*, Tome VIII: 65–73. Série VI.
- De Donder, Theophile. 1926. Application de la relativité aux le systèmes atomiques et moléculaires. Comptes rendus hebdomadaires des séances de l'Académie des sciences, 182: 1380–1382.
- De Donder, Theophile and Frans H. van den Dungen. 1926. La quantification déduite de la
 Gravifique einsteinienne. Comptes rendus hebdomadaires des séances de l'Académie des
 sciences, 183: 22-24.
- De Donder, Theophile. 1926. Application de la quantification déduite de la Gravifique einsteinienne. Comptes rendus hebdomadaires des séances de l'Académie des sciences, 183:
 594–595
- 1844 De Donder, Theophile. 1927. The Mathematical Theory of Relativity. Cambridge, MA: MIT.
- De Donder, Théophile. 1927. Le Principe de Corrispondance déduit de la Gravifique et la
 Mécanique ondulatoire. (Quatrième communication). Bulletin de l'Académie royale de
 Belgique [Classe des Sciences], 13: 504–509. Serie 5.
- De Donder, Théophile. 1930. Einsteinian gravity. Annales de l'Institut Henri Poincaré, 1:
 77–116.
- De Donder Théophile. 1930. Théorie invariantive du calcul des variations. Paris: Gauthier Villars.
- Dirac, Paul A. M. 1928. The Quantum Theory of the Electron. Proceedings of the Royal
 Society A: Mathematical, Physical and Engineering Sciences, 117: 610–624.
- Duff, Michael J. 1973. Quantum Tree Graphs and the Schwarzschild Solution. *Physical Review D*, 7: 2317–2326.
- Duff, Michael J., B.E.W. Nilsson and C.E. Pope. 1986. Kaluza-Klein Supergravity. *Physics Reports*, 130: 1–142.
- Eddington, Arthur S. 1923. The Mathematical Theory of Relativity. Cambridge: Cambridge
 University Press.
- Einstein, Albert. 1916. N\"aherungsweise Integration der Feldgleichungen der Gravitation.
 Preussische Akademie der Wissenschaften, Berlin, pp. 688–696.
- Giulini, Domenico and André Grossardt. 2012. The Schrödinger–Newton equation as a non relativistic limit of self-gravitating Klein–Gordon and Dirac fields. *Classical and Quantum Gravity*, 29: 215010.

- Gordon, Walter. 1927. Der Comptoneffekt nach der Schrdingerschen Theorie. Zeitschrift für
 Physik, 40: 117–133.
- Gorelik, Gennady E. and Viktor Frenkel. 1994. Matvei Petrovich Bronstein and Sovjet
 Theoretical Physics in the Thirties. Basel: Birkhäuser.
- Hagar, Amit. 2014. Discrete or Continuous? The Quest for Fundamental Length in Modern
 Physics. Cambridge University Press.
- Hawking, Stephen W. 1975. Particle Creation by Black Holes. Communications in Mathe matical Physics, 43: 199–220.
- Heisenberg, Werner. 1927. Über den anschaulichen Inhalt der quantentheoretischen Kinematik und Mechanik, Zeitschrift für Physik, 43: 172–198.
- Hilbert, David. 1900. Mathematische Probleme Vortrag, gehalten auf dem internationalen Mathematiker-Kongreß zu Paris 1900. Nachrichten von der Gesellschaft der Wissenschaften zu Göttingen, Mathematisch-Physikalische Klasse: 253–297.
- Jacobsen, Anja S. 2012. Léon Rosenfeld. Physics, Philosophy, and Politics in the Twentieth
 Century. Singapore: World Scientific.
- Jaffé, George. 1922. Bemerkungen über die relativistischen Keplerellipsen. Annalen der
 Physik, 372: 212.
- Jeffery, George B. 1921. The Field of an Electron on Einstein's Theory of Gravitation.
 Proceedings of the Royal Society of London A, 99: 123–134.
- Jordan, Pascual and Oslar Klein. 1927. Zum Mehrkörperproblem der Quantentheorie.
 Zeitschrift für Physik, 45: 751–765.
- Jordan, Pascual. 1947. Erweiterung der projektiven Relativitätstheorie. Annalen der Physik,
 1: 219–228.
- Kaluza, Theodor. 1984. On the Unification Problem in Physics. In: Lee, H. C., editor, An
 Introduction to Kaluza-Klein Theories Workshop on Kaluza-Klein Theories, p. 1. Chalk
 River, Ontario, Canada: World Scientific. Translated by Taizo Muta.
- Kaluza, Theodore. 1921. Zum Unitätsproblem in der Physik. Sitzungsberichte der Koniglich
 Akademieder Preussischen Akademie der Wissenschaften, 1: 966–972.
- Kanatchikov, Igor V. 2014. On precanonical quantization of gravity. Nonlinear Phenomena in Complex Systems, 17: 372–376.
- 1897 Kiefer, Claus. 2004. Quantum Gravity. Oxford: Claredon Press.
- 1898 Klein, Oskar. 1991. From my Life of Physics. In: The Oskar Klein Memorial Lectures. Vol.
- 1: Lectures by C. N. Yang and S. Weinberg with translated reprints by O. Klein. Editor:
 Gösta Ekspong. Singapore: World Scientific Publishing Co. Pte. Ltd.
- ¹⁹⁰¹ Klein, Oskar. 1984. "Quantum Theory and five-dimensional Relativity" by Oskar Klein. In:
- Lee, H. C., editor, An Introduction to Kaluza-Klein Theories Workshop on Kaluza-Klein
 Theories: 10–21. Chalk River, Ontario, Canada: World Scientific. Traduzione a cura di
 Taizo Muta.
- Klein, Oskar. 1926. Quantentheorie und fünfdimensionale Relativitätstheorie. Zeitschrift für
 Physik, 37: 895–906.
- ¹⁹⁰⁷ Klein, Oskar. 1926. The atomicity of electricity as a quantum theory law. *Nature*, 118: 516.
- Klein, Oskar. 1927. Sur l'article de M. L. de Broglie: L'univers a cinq dimensions et la
 mécanique ondulatoire. Le Journal de Physique et le Radium, Tome VIII: 242–243. Série
 VI.
- Klein, Oskar. 1927. Zur fünfdimensionalen Darstellung der Relativitätstheorie. Zeitschrift
 für Physik, 46: 188–208.
- Klein, Oskar. 1927. Elektrodynamik und Wellenmechanik vom Standpunkt des Korrespon denzprinzip, Zeitschrift für Physik, 41: 407–442.
- Kramers, Hendrik A. 1922. On the application of Einstein's theory of gravitation to a stationary field of gravitation, *Proceedings Koninklijke Akademie van Wetenschappen*, 23:
 1052–1073.
- Kuhn, Thomas S. and John L. Heilbron. 1963. Interview with Dr. Leon Rosenfeld by Thomas
 S. Kuhn and John L. Heilbron At Carlsberg. July 1, 1963. College Park, MD USA: Niels
- ¹⁹²⁰ Bohr Library & Archives, American Institute of Physics. Session I.

- Landau, Lev D. and Evgenij M. Lifšhitz. 1951. The Classical Theory of Fields. Cambridge:
 Addison-Wesley.
- Berestetskii, Valdimir, Evgenij M. Lifšhitz and Lev Pitaevskii. 1971. Relativistic Quantum
 Theory. Oxford: Pergamon Press.
- Lodge, Oliver. 1921. The Gravitational Field of an Electron. Nature, 107: 392.
- Mehra, Jagdish and Helmut Rechenberg. 2001. The Historical Development of Quantum
 Theory 1–6. New York: Springer-Verlag.
- Mehra, Jagdish and Helmut Rechenberg. 2001. The Probability Interpretation and the
 Statistical Transformation Theory, the Physical Interpretation, and the Empirical and
 Mathematical Foundations of Quantum Mechanics 1926–1932. The Historical Develop ment of Quantum Theory, Vol. 6, The Completion of Quantum Mechanics 1926–1941,
- Part I. New York: Springer-Verlag.
- Misner, Charles W. and Kip S. Thorne and John A. Wheeler. 1973. Gravitation. W.H.
 Freeman and Company.
- Møller, Christian. 1962. The energy-momentum complex in general relativity and related
 problems. In Les théories relativistes de la gravitation. (ed. A. Lichnerowicz and M. A.
 Tonnelat), Paris: Editions du Centre National de la Recherche Scientifique.
- Nordström, Gunnar. 1914. Über die Möglichkeit, das elektromagnetische Feld und das
 Gravitationsfeld zu Vereinigen. *Physlische Zeitschrift*, 15: 504–506.
- O'Raifeartaigh, Lochlain and Norbert Straumann. 2000. Gauge theory: Historical origins
 and some modern developments. *Reviews of Modern Physics*, 72: 1–23.
- Overduin, James M. and Paul S. Wesson. 1997. Kaluza-Klein Gravity. *Physics Reports*, 283:
 303–380.
- Pais, Abraham. 1982. "Subtle is the Lord ...". The Science and Life of Albert Einstein.
 Oxford: Oxford University Press.
- Pais, Abraham. 2000. The Genius of Science: A Portrait Gallery. Oxford: Oxford University
 Press.
- Pauli, Wolfgang. 1927. Zur Quantenmechanik des magnetischen Elektrons Zeitschrift für
 Physik, 43: 601–623.
- Pauli, Wolfgang. 1993. Wissenschaftlicher Briefwechsel mit Bohr, Einstein, Heisenberg u.a.
 Band III: 1940–1949/Scientific Correspondence with Bohr, Einstein, Heisenberg a.o.
- Volume III: 1940–1949. Edited by Karl von Meyenn. Berlin Heidelberg: Springer-Verlag.
 Penrose, Roger. 1996. On gravity's role in quantum state reduction. General Relativity and
 Gravitation, 28: 581–600.
- Rickles, Dean. 2013. "Pourparlers for Amalgamation: Some Early Sources of Quantum Grav ity Research". In: Shaul Katzir, Christoph Lehner and Renn Jürgen, editors, Traditions
- and Transformations in the History of Quantum Physics, Chapter 6. Max Planck Research
 Library for the History and Development of Knoledge. Proceedings 5. Third Interna tional Conference on the History of Quantum Physics, Berlin, June 28–July 2, 2010;
 http://www.edition-open-access.de/proceedings/5/index.html.
- Rickles, Dean. 2005. "Pioneers of Quantum Gravity". Talk presented at the Third Conference
 on History of Quantum Physics (HQ3).
- Robertson, Baldwin. 1972. Introduction to field operators in quantum mechanics. American
 Journal of Physics, 41:678–690.
- Rocci, Alessio. 2013. On first attempts to reconcile quantum principles with gravity. Journal
 of Physics: Conference Series, 470: 12004.
- Rocci, Alessio. 2015. Oliver in Quantum-Gravity-land. http://www.oliverlodge.org/oliver in-quantum-gravity-land/. Based on talk given at 3rd Making Waves Workshop. October,
 31 Liverpool.
- Rocci, Alessio. 2015. History of Quantum Gravity: from the birth of General Relativity to
 the end of WWII 1915–1945. http://paduaresearch.cab.unipd.it/8916, Language: Italian.
- Rosenfeld, Léon. 1927. L'Univers a cinq dimensions et la Mécanique ondulatoire. Bulletin
 de l'Académie royale de Belgique [Classe des Sciences], 13: 304–325. Serie 5.
- 1974 Rosenfeld, Léon. 1927. L'Univers a cinq dimensions et la Mécanique ondulatoire. (Deuxième
- communication). Bulletin de l'Académie royale de Belgique [Classe des Sciences], 13:
 447-458. Serie 5.

- Rosenfeld, Léon. 1927. L'Univers a cinq dimensions et la Mécanique ondulatoire. (Troisième communication). Bulletin de l'Académie royale de Belgique [Classe des Sciences], 13:
 573-579. Serie 5.
- Rosenfeld, Léon. 1927. L'électron magnétique et la mécanique ondulatoire. Comptes rendus
 hebdomadaires des séances de l'Académie des sciences, T184: 1540–1541.
- Rosenfeld, Léon. 1927. L'Univers a cinq dimensions et la Mécanique ondulatoire. (Quatrième
- communication). Bulletin de l'Académie royale de Belgique [Classe des Sciences], 13:
 661–682. Serie 5.
- Rosenfeld, Léon. 1930. Über die Gravitationswirkungen des Lichtes. Zeitschrift für Physik,
 65: 589–599.
- 1987 Rosenfeld, Léon. 1930. Zur Quantelung der Wellenfelder. Annalen der Physik, 5: 113–152.
- 1988 Rosenfeld, Léon. 1963. On quantization of fields. Nuclear Physics, 40: 353–356.
- Rosenfeld, Léon. 2017. On the quantization of wave fields. The European Physical Journal H, 42: 63–94.
- Salisbury, Donald and Kurt Sundermeyer. 2017. Léon Rosenfeld's general theory of
 constrained Hamiltonian dynamics. The European Physical Journal H, 42: 23–61.
- Schrödinger, Erwin. 1926. Quantisierung als Eigenwertproblem. (Erste Mitteilung). Annalen
 der Physik, 79: 361–376.
- Schrödinger, Erwin. 1927. Der Energieimpulssatz der Materiewellen. Annalen der Physik,
 82: 265–272.
- 1997 Solomon, Jacques. 1938. Gravitation et quanta. Journal de Physique et le Radium, 9: 479–
 485.
- Stachel, John. 1999. Introduction. In: Tian Yu Cao, editor, *Conceptual foundations of quantum field theory*, Chapter V, Quantum field theory and space-time. Cambridge: Cambridge University Press. pp. 166–175.
- Rickles, Dean, and Steven Weinstein. 2016. "Quantum Gravity", The Stanford
 Encyclopedia of Philosophy (Winter 2016 Edition), Edward N. Zalta (ed.),
 https://plato.stanford.edu/archives/win2016/entries/quantum-gravity/.
- Thiry, Yves. 1948. Les équations de la théorie unitaire de Kaluza. Comptes rendus
 hebdomadaires des séances de l'Académie des sciences, T226: 216–218.
- Vallarta, Manuel Sandoval. 1924. Bohr's Atomic Model from the Standpoint of the General Theory of Relativity and of the Calculus Of Perturbations. Ph.D. thesis, Cambridge, MA, USA: Massachusetts Institute of Technology.
- von Borzeszkowski, Horst-Heino and Hans J. Treder. 1988. The Meaning of Quantum
 Gravity. Dordrecht: D. Reidel Publishing Company.

Author Query

- **Q1** Please provide volume number for Refs. "Blum et al. (2017)" and "Hilbert (1900)".
- Q2 Please provide publisher name for Ref. "Einstein (1916)."
- **Q3** Please provide publisher location for Refs. "Hagar (2014) and Misner et al. (1973)."
- Q4 Please provide an update for Ref. "Kanatchikov (1998)", if available.