

Insights into Juina super-deep diamonds

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Inclusions in diamonds may provide direct samples from the otherwise inaccessible Earth's lower mantle, as was first suggested almost thirty years ago. Nevertheless, it is only in the last two decades that diamonds of a deeper origin have been identified as minor components from a number of kimberlitic and alluvial diamond deposits worldwide [1, and references therein]. These so-called super-deep diamonds may originate between 300 and 800 km depth and thus provide important natural evidence on phases and phase compositions that allow unique insights into the processes and states in the extreme depths of our planet. They mainly contain ferropericlase, enstatite (believed to be derived from MgSi-perovskite) and CaSiO₃-walstromite (believed to be derived from CaSiO₃-perovskite).

In order to obtain some information on their depth of origin, two alluvial diamonds coming from São Luiz (Juina, Brazil) were investigated both by *in-situ* microRaman spectroscopy and *in-situ* single-crystal X-ray diffraction. In total we analyzed more than sixty inclusions and we found evidence of CaSiO₃-walstromite, larnite, CaSi₂O₅-titanite, perovskite, orthoenstatite, siderite, phase EGG and probably ferropericlase; other phases are present but they still need to be identified. In particular, at least one inclusion showed the coexistence of CaSiO₃-walstromite, larnite and CaSi₂O₅-titanite, indicating that CaSiO₃-walstromite can effectively form in sublithospheric conditions from the back transformation of larnite + CaSi₂O₅-titanite. Our preliminary thermodynamic calculations, however, suggested us that single inclusions of CaSiO₃-walstromite cannot derive from CaSiO₃-perovskite. This is due to the high volume change from the high-pressure to the low-pressure polymorph, that would require ~30% of expansion in the host. Further analyses are in progress in order to determine the unit-cell of ferropericlase and phase EGG.

[1] Kaminsky (2008) *Earth Sc Rev* **110**, 127-147.