

Erratum

Structural properties of disk galaxies

I. The intrinsic equatorial ellipticity of bulges

J. Méndez-Abreu^{1,2,4}, J. A. L. Aguerri³, E. M. Corsini⁴, and E. Simonneau⁵

¹ INAF - Osservatorio Astronomico di Padova, vicolo dell'Osservatorio 5, 35122 Padova, Italy
e-mail: jairo.mendez@oapd.inaf.it

² Universidad de La Laguna, Av. Astrofísico Francisco Sánchez s/n, 38206 La Laguna, Spain

³ Instituto Astrofísico de Canarias, Calle vía Láctea s/n, 38200 La Laguna, Spain
e-mail: jalfonso@iac.es

⁴ Dipartimento di Astronomia, Università di Padova, vicolo dell'Osservatorio 3, 35122 Padova, Italy
e-mail: enricomaria.corsini@unipd.it

⁵ Institut d'Astrophysique de Paris, CNRS, 98bis Bd Arago, 75014 Paris, France

A&A 478, 353–369 (2008), DOI: 10.1051/0004-6361:20078089

Key words. galaxies: bulges – galaxies: formation – galaxies: fundamental parameters – galaxies: photometry – galaxies: statistics – errata, addenda

Figures 4 and 6 of the original paper are incorrect since the data were plotted with the axes inverted. These mistakes do not influence any result or conclusion of the paper. The corrected figures appear below.

Acknowledgements. We thank A. Kalnajs for bringing this to our attention.

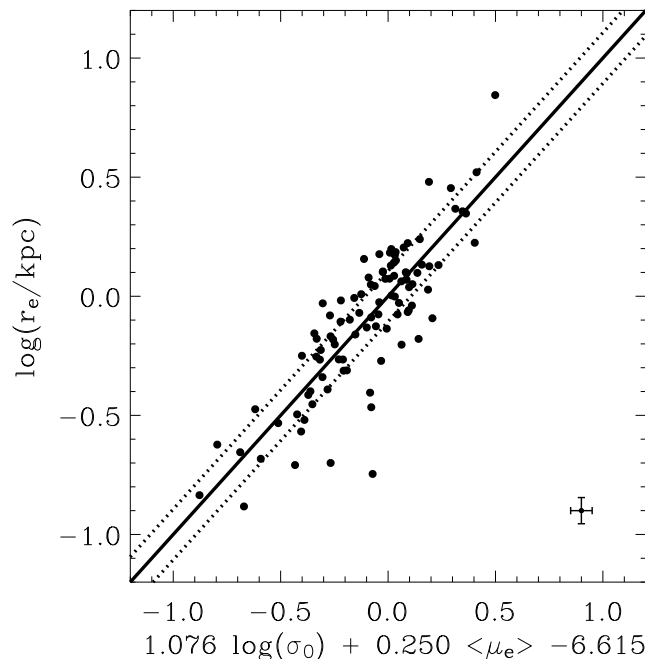


Fig. 4. Edge-on view of the FP for the 98 early-to-intermediate type bulges of our sample with measured velocity dispersion. The solid line represents the linear fit to the data. The dotted lines represent the 1σ deviation from the fit. The error bars in the lower right corner of the panel indicate the mean errors of the data.

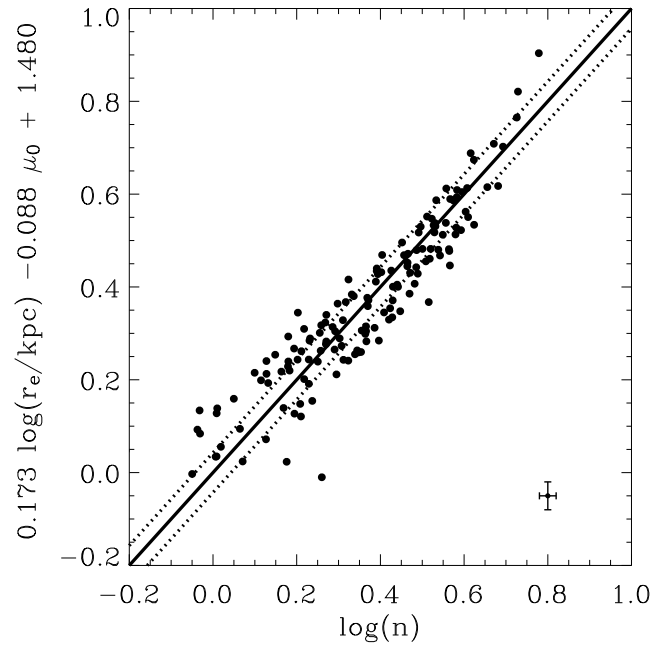


Fig. 6. Edge-on view of the PP for early-to-intermediate bulges of our sample. The solid line represents the linear fit to the data. The dotted lines represent the 1σ deviation from the fit. The error bars in the lower right corner of the panel indicate the mean errors of the data.